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COMPUTER PROGRAMS FOR THE ANALYSIS OF
SPACECRAFT MAGNETISM

Milton H. Lackey

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White Oak, Maryland

28 September 1973

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13. ABSTRACT This report describes five computer programs that are being used at the Naval Ordnance Laboratory for analyzing satellite magnetism. The programs contain numerical analysis algorithms for the spherical harmonic analysis of the magnetic field emanating from a satellite. The analysis is directed at determining the components of the magnetism which correspond to the dipole moment, quadrupole moment, etc. The first three programs were devised to analyze data from magnetic field measurements around the satellite. The fourth program was devised to generate simulated measurement data for a specified system of multipole magnets. The last program is a combination of the data generation and the data analysis programs. Sample problems are included in the discussion to illustrate the techniques of using the programs with a CDC 6400 Computer including the INTERCOM time-sharing system. A brief description is also included of data acquisition techniques and of principal subprograms.			

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Computer Programs for the Analysis of Spacecraft Magnetism

During the past several years techniques have been developed for measuring and analyzing the magnetism of spacecraft. The techniques are directed, first, at estimating the spacecraft's magnetic dipole moment and, second, at compensating the dipole moment to allow the spacecraft to maintain a stable orientation while in orbit. This procedure requires accurate measurements of the magnetic field emanating from the spacecraft. The Naval Ordnance Laboratory is currently involved in the development of a facility to conduct sophisticated magnetic tests of spacecraft. The facility will contain instrumentation which will automatically record and analyze the test data. This report has been published to document the numerical techniques to be used in the analysis of the data. Techniques are also described which can be used to predict the accuracy of different types of measurement and analysis techniques.

The development and testing of the computer programs required a considerable amount of effort. Mr. H. W. Korab contributed much to this effort, including the development of the BASIC version of the analysis procedure. He also assisted in the development of the illustrations included in this report.

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ROBERT WILLIAMSON II
Captain, USN
Commander


R. B. KNOWLES
By direction

CONTENTS

Chapter	Page
1. INTRODUCTION.....	1
2. DATA ACQUISITION TECHNIQUES.....	3
3. MAIN PROGRAMS.....	7
Programs SA1024 and SA2024.....	7
Introduction.....	7
Input Data.....	7
Output Data.....	8
Sample Problem for SA2024.....	9
Program SA3024.....	10
Introduction.....	10
Input Data.....	10
Output Data.....	12
Integrating Schemes.....	14
Units.....	19
Sample Problems.....	19
Program SA4024.....	20
Introduction.....	20
Input Data.....	20
Output Data.....	22
Sample Problems.....	23
Program SA5024.....	24
Introduction.....	24
Input Data.....	24
Output Data.....	26
Sample Problems.....	26

NOLTR 73-191

CONTENTS (CONT)

Chapter	Page
4. PRINCIPAL SUBPROGRAMS.....	31
Introduction.....	31
Subroutine AMPMNT.....	31
Subroutine AMPFLD.....	32
Functions PNM, SPNM, and ANF.....	35
Subroutine DATPLT.....	36
Subroutine FNCTON.....	37
5. CONCLUSIONS.....	39
APPENDIX A GLOSSARY OF SYMBOLS AND TERMS.....	A-1
APPENDIX B LISTING OF SA1024.....	B-1
APPENDIX C LISTING OF SA2024.....	C-1
APPENDIX D SAMPLE PROBLEM FOR SA2024.....	D-1
APPENDIX E LISTING OF SA3024.....	E-1
APPENDIX F SAMPLE PROBLEMS FOR SA3024.....	F-1
APPENDIX G LISTING OF SA4024.....	G-1
APPENDIX H SAMPLE PROBLEMS FOR SA4024.....	H-1
APPENDIX I LISTING OF SA5024.....	I-1
APPENDIX J SAMPLE PROBLEMS FOR SA5024.....	J-1
APPENDIX K NOL TECHNICAL NOTE 9726.....	K-1

ILLUSTRATIONS

Figure	Title	Page
1	Illustration of Spherical Polar Coordinates.....	40
2	Degrees of Freedom for Measurements.....	41
3	Sphere of Measurements.....	42
4	Satellite Initial positions for NO = 6.....	43
5	Illustration of Sample Problem.....	44
6	Typical Data Curves.....	45
7	Example of Surface Area Elements Assigned to Data Point.....	46
8	Examples of Intervals of Colatitude.....	47
9	Data Curves from SA3024 for Sample Problem.....	48
10	Data Curves from SA4024 for Sample Problem.....	49

ILLUSTRATIONS (CONT)

Figure	Title	Page
11	Dipole Curves for $A_1^0 = D_3 = 1000$ gauss-cm ³	50
12	Dipole Curves for $A_1^1 = D_1 = 1000$ gauss-cm ³	51
13	Dipole Curves for $B_1^1 = D_2 = 1000$ gauss-cm ³	52
14	Quadrupole Curves for $A_2^0 = 1000$ gauss-cm ⁴	53
15	Quadrupole Curves for $A_2^1 = 1000$ gauss-cm ⁴	54
16	Quadrupole Curves for $A_2^2 = 1000$ gauss-cm ⁴	55
17	Quadrupole Curves for $B_2^1 = 1000$ gauss-cm ⁴	56
18	Quadrupole Curves for $B_2^2 = 1000$ gauss-cm ⁴	57
19	Quadrupole Curves for $Q_{11} = 1000$ gauss-cm ⁴ and $Q_{22} = Q_{33} = -500$ gauss-cm ⁴	58
20	Quadrupole Curves for $Q_{22} = 1000$ gauss-cm ⁴ and $Q_{11} = Q_{33} = -500$ gauss-cm ⁴	59
21	Quadrupole Curves for $Q_{33} = 1000$ gauss-cm ⁴ and $Q_{11} = Q_{22} = -500$ gauss-cm ⁴	60
22	Quadrupole Curves for $Q_{12} = 1000$ gauss-cm ⁴	61
23	Quadrupole Curves for $Q_{13} = 1000$ gauss-cm ⁴	62
24	Quadrupole Curves for $Q_{23} = 1000$ gauss-cm ⁴	63
25	Data Curves from SA5024 for Sample Problem.....	64
26	Multipole Component of Degree 1 (Dipole Component).....	65
27	Multipole Component of Degree 2 (Quadrupole Component).....	66
28	Multipole Component of Degree 3.....	67
29	Multipole Component of Degree 4.....	68
30	Multipole Component of Degree 5.....	69
31	Multipole Component of Degree 6.....	70
32	Multipole Component of Degree 7.....	71
33	Multipole Component of Degree 8.....	72
34	Summation of First Eight Multipole Components.....	73
35	Individual Curves Demonstrating Approximation Accuracy.....	74
36	Data Curves for Radially Offset Dipole.....	75
37	Data Curves for Transversely Offset Dipole.....	76
38	Sample Graph From Electrostatic Plotter.....	77

NOLTR 73-191

TABLES

Table	Title	Page
1	Examples of Minimum Values of NO and NI.....	15
2	Peak Data for the Multipole Components of the Sample Problem.....	27
3	Maximum Error Versus Parametric Variation in Computing the Dipole Moment for the Sample Problem.....	29

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Chapter 1

INTRODUCTION

1. A major effort has been applied to the development of computer programs and subprograms to assist in the analysis of satellite magnetism. The analysis is part of a procedure directed at determining and compensating the magnetic dipole moment of a satellite. The procedure involves four steps. The first step is to measure the normal component of the satellite's magnetic field on the surface of a sphere enclosing the satellite. The second step involves the spherical harmonic analysis of the measured data to determine the magnitude and the direction of the satellite's dipole moment. The third step is to attach an opposing dipole to the satellite with equal magnitude and opposite direction. The final step is to repeat the measurement and analysis procedures to verify the compensation.

2. The five programs described in this report have been devised to perform a variety of data analysis tasks. One of the programs is coded in the BASIC computer language. The other four are coded in FORTRAN IV. The first three programs are used to analyze data representing the normal component of the magnetic field on the surface of a sphere enclosing the satellite. The fourth program is used to generate simulated measurement data for a specified system of multipole magnets. The program also allows the simulation of measurement errors. The last program is a combination of the data generation and the analysis programs. The programs have been designed to allow a variety of options including:

- a. Reading the input data from paper tape or from a data file
- b. Using either an algebraic or geometric integrating scheme to perform the analysis
- c. Printing the data after preliminary data processing
- d. Interpolating and plotting the processed data curves.

3. The discussion begins with a brief description of data acquisition techniques. This is followed by a description of the five programs including sample problems which demonstrate the use of the programs. Finally, a description is given of some of the primary subprograms. These are listed separately since they can be used as building blocks for other programs. The discussion is supplemented by several illustrations, many of which are the actual computer outputs from the execution of the computer programs. The majority of the data curves are output from the CALCOMP 570 digital incremental plotter. A description of the data units for the programs has also been included in paragraph 33. Appendix A contains a glossary of symbols and terms used in the report.

Chapter 2

DATA ACQUISITION TECHNIQUES

4. A brief description of spherical coordinates and the techniques used in the data acquisition will assist in the description of the parameters used in the data analysis and the computer programs. Figure 1 illustrates the relationship between rectilinear and spherical coordinates. An arbitrary point \bar{R} (or vector) in space can be defined in terms of the spherical coordinates of R , θ , and φ representing the radial distance, the colatitude and the (easterly) longitude respectively. (θ is sometimes called the polar angle, and φ is sometimes called the dihedral angle.) Appendix A gives the definitions of these coordinates in terms of rectilinear coordinates.

5. Figure 2 shows a simplified diagram of the test setup for conducting the measurements of a satellite's magnetism. The degrees of freedom illustrated in the figure are defined to correspond to the spherical angles. The rotation axis, provided by a horizontal turntable, allows a variation in colatitude. The tilt axis, provided by a gimbaling fixture attached to the satellite, allows a variation in longitude. Analog curves are made of the sensor reading versus the colatitude for a fixed set of positions of longitude. The curves are recorded while the satellite is in a zero magnetic field environment. Notice that the sensor in Figure 2 is radially aligned. This setup generates analog curves representing the normal component of the satellite magnetism along great circles of longitude as shown in Figure 3.

6. Let the parameter NO represent the number of curves of data to be recorded. These curves correspond to measurements along great circles spaced $(180/NO)$ degrees apart. Figure 3 shows an example when $NO = 6$. There will be six analog curves of data corresponding to the six great circles. Figure 4 shows the initial positions for each of the curves.

7. As a sample problem consider the system of dipoles illustrated in Figure 5. The first curve begins with the +z-axis directed at the magnetic sensor and the +y-axis pointed up. The turntable is rotated clockwise. Prior to beginning each succeeding curve, the satellite is tilted ($180/N_0 = 30$) degrees around the z-axis. Figure 6 shows a typical set of six data curves for the sample problem. Each curve is marked with 25 data points. The parameter N_1 is used to represent the number of data points per curve ($N_1 = 25$ in the example). The points are spaced ($360/(N_1-1)$) = 15 degrees apart. The value at each measurement point is automatically recorded via analog-to-digital (A/D) conversion equipment in the satellite measurements.

8. Notice that the curves in Figure 6 do not all begin with the same value, although, theoretically they should. This is typical of the type of measurements that are made on satellites. The curves represent relative measurements instead of absolute measurements. It would be possible to insure that the curves all begin with the same value, but it would not necessarily be the correct value. Nevertheless, the data analysis is independent of the starting value of each curve. Therefore, no concern is given to this problem in the acquisition of the data.

9. Reference (a) lists several considerations in the determination of the parameters for the data acquisition including N_0 , N_1 , and the radius R_1 of the measurement sphere. The considerations include:

- a. The smoothness of the data curves
- b. The analysis method
- c. The accuracy of the measuring apparatus
- d. The round-off errors of the computing machine.

The parameters are not all independent. They must be determined by making certain compromises. For instance, the data curves can be smoothed by increasing R_1 (i.e., by increasing the minimum distance between the surface of the measurement sphere and the satellite), but this will decrease the relative accuracy of the data. Also, the total number of data points $N_0 \cdot N_1$ may be increased, but computing machine errors will become more significant. An increase in the number of data points will also increase the computing time and data storage.

10. Experience has shown that the dipole moment for most problems can be approximated accurately enough if the following conditions hold:

a. Let the minimum distance between the surface of the measurement sphere and the satellite be at least half the maximum diameter of the satellite to insure sufficient smoothness of the data curves. For example, the radius of the measurement sphere for a body with maximum diameter of 50 inches, and centered at the origin, should be at least 50 inches.

b. Let N_0 be ≥ 8 and N_1 be ≥ 16 to insure that the errors from the numerical approximation are smaller than the measurement errors. (This condition can be relaxed in cases where the satellite magnetism gives simple sine-cosine curves. In the past, $N_0 = 4$ and $N_1 = 25$ has been sufficient in many cases.)

c. Use the simplest analysis procedure (i.e., the geometric integrating scheme) for the preliminary analysis since it is relatively accurate and easy to use.

Chapter 3
MAIN PROGRAMS

PROGRAMS SA1024 AND SA2024

Introduction

11. The programs labeled SA1024 and SA2024 (Appendices B and C) are BASIC and FORTRAN IV versions, respectively, of the basic one-term data analysis procedure. The procedure is simple because only the dipole moment term is computed for the satellite magnetism and only the approximate method of numerical integration is used. The programs print the measurement data and the computed dipole moment in rectilinear and spherical coordinates. They also print data which assists in calibrating and aligning the compensating dipoles. The major difference between the two versions is the manner of entering the data. The BASIC version is set up to take the data from DATA statements. The FORTRAN IV version accepts data from paper tape or from a data file.

Input Data

12. The data is entered into the programs in the order of the following definitions:

a. First line of data or data card - format: (3I5)

NO - The number of curves of data (NO is even and ≤ 16 .)

N1 - The number of equally spaced data points per curve from 0 thru 360 degrees colatitude (N1 is odd and ≤ 33 . The first and last data points for each curve correspond to measurements at 0 degrees colatitude.)

IR - The parameter that determines whether or not to read the data from a data file. IR $\neq 0$ means that the data will be read from the data file DAT024. (This parameter is used only with the FORTRAN IV version.)

b. Second line of data or data card - format: (2F6.1)

R1 - The radius of the measurement sphere in inches

C6 - The value of the calibration signal (C7-C8) in gamma

c. Third line of data or data card - format: (I5)

N2 - The parameter that determines whether or not to print the measured data (N2 \neq 0 means that the data will be printed.)

d. Measurement data - format: (9F6.1) (These cards are deleted if IR \neq 0.)

C7 - The static measurement with the calibration signal

C8 - The static measurement without the calibration signal

F(I,J) - The measurement data for I = 1,2,...,N1 and J = 1,2,...,NO.

The input data cycle may be repeated by starting with new data NO, N1, and IR again. The program execution is terminated by setting the new value for NO equal to zero.

Output Data

13. The programs SA1024 and SA2024 then perform a numerical integration of the equation

$$\bar{D} = (3 \cdot R_1^3 / 8\pi) \int_0^\pi \int_0^{2\pi} f(\theta, \varphi) \sin \theta \, d\theta \, d\varphi \quad (1)$$

for the dipole moment \bar{D} . The numerical equation is

$$\bar{D} \approx (3 \cdot R_1^3 / 8\pi) \sum_{J=1}^{NO} \sum_{I=1}^{N1} F(I,J) \cdot \pi(\theta(I), \varphi(J)) \cdot X(I) \quad (2)$$

where

$\bar{D} \equiv (D_1, D_2, D_3)$ is in gauss-centimeters³

$F(I,J) \equiv f(\theta(I), \varphi(J))$ is in gammas

$$\bar{n}(\theta(I), \varphi(J)) = \begin{bmatrix} \sin(\theta(I)) \cdot \cos(\varphi(J)) \\ \sin(\theta(I)) \cdot \sin(\varphi(J)) \\ \cos(\theta(I)) \end{bmatrix}$$

$$\theta(I) \equiv 2\pi(I - 1)/(N1 - 1) \text{ is the colatitude (turntable angle)} \quad (3a)$$

$$\varphi(J) \equiv \pi(J - 1)/N0 \text{ is the longitude (tilt angle)} \quad (3b)$$

$$X(I) \equiv [\sin(\pi/(2 \cdot N1 - 2))]^2/(2 \cdot N0) \text{ if } I = 1 \text{ or } N1 \quad (4a)$$

$$\equiv [\sin(\pi/(2 \cdot N1 - 2))]^2/N0 \text{ if } I = (N1 + 1)/2 \quad (4b)$$

$$\equiv |\sin(\theta(I)) \cdot \sin(\pi/(N1 - 1))|/(2 \cdot N0) \text{ if } I \neq 1, (N1 + 1)/2, \text{ or } N1. \quad (4c)$$

14. Next, the dipole moment is transformed into spherical coordinates as

$$D \equiv (D_1^2 + D_2^2 + D_3^2)^{1/2}$$

$$\theta \equiv \tan^{-1} [(D_1^2 + D_2^2)^{1/2}/D_3] \cdot 180/\pi = \text{colatitude}$$

$$\varphi \equiv \tan^{-1} (D_2/D_1) \cdot 180/\pi = \text{longitude}$$

15. Finally, the program computes the values of two compensating magnets: one in the xy-plane and one along the z-axis. The value of the dipole field at one meter is also given to assist in charging the magnets to the desired values.

Sample Problem for SA2024

16. Appendix D contains the input and output data for SA2024 using the sample problem represented in Figures 5 and 6. The file BN2024, which is used for execution, is the binary version of SA2024. This example was executed in two different ways on the INTERCOM time-sharing system using the CDC 6400 computer. The information typed in at the teletype terminal has been underlined. Data tapes representing the measurement data are generated by the A/D equipment during the measurements. The program automatically reads the data in the proper format; converts the data into gammas using the calibration parameters C6, C7, and C8; and adjusts the data so that the beginning and ending points for all the curves have nearly the same value. The result of this procedure is visible in the data print-out on page 2 of Appendix D.

17. The first method of data input in Appendix D was from the file DAT024. This complete problem required only the first three lines of data. The measurement data was read into the file from an earlier execution of SA2024. (The program execution automatically generates the file DAT024 if the data is read in from tape.) The remaining lines of data initiated the computation of the solution to the same problem except that the measurement data was read in from a data tape.

PROGRAM SA3024

Introduction

18. The program SA3024, listed in Appendix E, allows a more complete analysis to be performed on the data. The program can be used to compute the dipole, quadrupole, and higher order multipole terms. Also included is an optional method of integrating the data and a data plotting option. The optional integrating scheme is labeled "exact" although it is exact only for magnetic data from a finite number of multipole magnets centered at the origin. This scheme is discussed in more detail in paragraphs 25 through 32 and in reference (a).

19. The program uses several special subprograms to perform such tasks as:

- a. The computation of spherical harmonic coefficients for magnetic field data
- b. The generation of values of associated Legendre polynomials and Schmidt functions
- c. The inversion of an $n \times n$ matrix
- d. The interpolation and plotting of data.

Some of these subprograms will be discussed in more detail in later sections.

Input Data

20. The method for entering the data into SA3024 is very similar to the method for SA2024 with several additional variables. The data is entered in the following order:

- a. First line of data or data card - format: (6I5)

NO - The number of curves of data (NO is even and ≤ 16 .)

- N1 - The number of equally spaced data points per curve from 0 thru 360 degrees colatitude (N1 is odd and ≤ 33 . The first and last data points for each curve correspond to measurements at 0 degrees colatitude.)
- NH - The highest degree spherical harmonic term to be computed from the data (NH = 1 for dipoles, 2 for quadrupoles, etc.)
- IR - The parameter that determines whether or not to read the data from a data file (IR \neq 0 means that the data will be read from the file DAT024.)
- IP1 - The parameter that determines whether or not the data is to be interpolated and plotted (IP1 \neq 0 means that the data will be interpolated and plotted.)
- IW - The parameter that determines which integrating scheme is to be used (IW = 0 means that the exact, algebraic scheme is to be used.)

b. Second line of data or data card - format: (3F8.4)

- R1 - The radius of the measurement sphere in inches
- CV - The value of the calibration signal (CS-CZ) in gamma
- PY - The scale factor (gammas/inch) for the y-axis if the data is to be plotted (If PY = 0.0 a factor will be computed from the data.)

c. Third line of data or data card - format: (I5)

- IP2 - A parameter that determines whether or not to print the measured data (IP2 \neq 0 means that the data will be printed.)

d. Measurement data - format: (9F6.1) (These cards are deleted if IR \neq 0.)

- CS - The static measurement with the calibration signal
- CZ - The static measurement without the calibration signal

$F(I, J)$ - The measurement data for $I = 1, 2, \dots, N1$ and $J = 1, 2, \dots, N0$.

The input data cycle may be repeated by starting with new data $N0, N1, \dots$, etc. The program is terminated by setting the new value for $N0$ equal to zero.

Output Data

21. The program then computes the spherical harmonic coefficients A_n^m and B_n^m based on the equation

$$\begin{pmatrix} A_n^m \\ B_n^m \end{pmatrix} = \left[(2n+1) \cdot R1^{n+2} / (4\pi \cdot (n+1)) \right] \int_0^\pi \int_0^{2\pi} f(\theta, \varphi) \cdot P_n^m(\cos \theta) \begin{Bmatrix} \cos(m\varphi) \\ \sin(m\varphi) \end{Bmatrix} \cdot \sin \theta d\theta d\varphi \quad (5)$$

for $n = 0, 1, \dots, NH$ and $m = 0, 1, \dots, n$. The numerical equivalent of equation (5) is

$$\begin{pmatrix} A_n^m \\ B_n^m \end{pmatrix} = \left[(2n+1) \cdot R1^{n+2} / (4\pi \cdot (n+1)) \right] \sum_{J=1}^{N0} \sum_{I=1}^{N1} F(I, J) \cdot P_n^m(\cos(\theta(I))) \cdot \begin{Bmatrix} \cos(m\varphi(J)) \\ \sin(m\varphi(J)) \end{Bmatrix} \cdot Y(I) \quad (6)$$

where

n is the degree of the spherical harmonic term

m is the order of the spherical harmonic term

$P_n^m(\cos(\theta(I)))$ is the Schmidt function of degree n and order m . (These functions are discussed in Appendix K.)

$Y(I)$ is the array of weighting factors for the numerical integration. (One of two methods can be used to determine the values for $Y(I)$ depending on the parameter IW .)

$\theta(I) \equiv 2\pi(I-1)/(N1-1)$ is the colatitude (turntable angle)

$\varphi(J) \equiv \pi(J-1)/N0$ is the longitude (tilt angle).

Equation (5) gives an expansion for the function $f(\theta, \varphi)$ as

$$f(\theta, \varphi) = \sum_{n=0}^{\infty} (n+1)/(R_1^{n+2}) \sum_{m=0}^n [A_n^m \cos(m\varphi) + B_n^m \sin(m\varphi)] P_n^m(\cos(\theta)). \quad (7)$$

If the expansion is written as $f(\theta, \varphi) = \sum_{n=0}^{\infty} f_n(\theta, \varphi)$ then each $f_n(\theta, \varphi)$ represents the field from a multipole magnet of degree n . (Reference (a) contains more details on the expansion of the function $f(\varphi, \theta)$.)

22. In general, the coefficients A_n^m and B_n^m can be stored as two-dimensional arrays $A(n, m)$ and $B(n, m)$, or, they may be packed into single-dimensional arrays as

$$A(m+1 + (n^2 + n)/2) \equiv A_n^m \text{ for } n = 0, 1, 2, \dots \text{ and } m = 0, 1, \dots, n$$

$$B(m + (n^2 - n)/2) \equiv B_n^m \text{ for } n = 1, 2, \dots \text{ and } m = 1, 2, \dots, n.$$

23. The program SA3024 is set up to conserve storage by computing and storing the coefficients for only one degree term at a time (i.e., for each fixed n) using the equations

$$A(m+1) \equiv A_n^m \text{ for } m = 0, 1, \dots, n$$

$$B(m) \equiv B_n^m \text{ for } m = 1, 2, \dots, n. \quad (8)$$

It should be noted that the coefficient A_0^0 , corresponding to the monopole moment, would be zero if the numerical integration was exact and the data $F(I, J)$ was correct. If the monopole moment is not zero, then the data is corrected prior to being printed and prior to further analysis. Also for $n = 1$ and $Y(I) = X(I)$ it can be shown that Eqs. (1) (or (2)) and (5) (or (6)) are identical if

$$\begin{aligned} D_1 &= A_1^1 \\ D_2 &= B_1^1 \\ D_3 &= A_1^0. \end{aligned} \quad (9)$$

24. The dipole moment $\bar{D} = (D_1, D_2, D_3)$ is computed and printed out separately in the program in both rectilinear and spherical coordinates. The program is also

set up to compute the quadrupole moment Q_{ij} for i and $j = 1, 2$, and 3 . This is computed and printed separately if $NH \geq 2$. The computation is based on the equations

$$Q_{11} = \sqrt{3} \cdot A_2^2 - A_2^0 \quad (10a)$$

$$Q_{22} = -\sqrt{3} \cdot A_2^2 - A_2^0 \quad (10b)$$

$$Q_{33} = 2A_2^0 \quad (\text{i.e., } Q_{11} + Q_{22} + Q_{33} = 0) \quad (10c)$$

$$Q_{12} = \sqrt{3} \cdot B_2^2 \quad (10d)$$

$$Q_{13} = \sqrt{3} \cdot A_2^1 \quad (10e)$$

$$Q_{23} = \sqrt{3} \cdot B_2^1 \quad (10f)$$

$$Q_{ij} = Q_{ji} \text{ for } i \text{ and } j = 1, 2, \text{ and } 3. \quad (10g)$$

The higher degree coefficients for $n = 3, 4, \dots, NH$ are printed only in their spherical form.

Integrating Schemes

25. The program SA3024 is set up to allow a choice between two numerical integration schemes for Eq. (6). These two schemes are based on two different methods for determining the weighting factors $\{Y(I)\}$. The factors actually correspond to the elements of spherical surface area assigned to each data point $F(I, J)$. Therefore, a geometrical description of the two different sets of weighting factors will provide some insight into the two methods of numerical integration. Figure 7 shows an example of the areas assigned to a data point for both integrating methods. The dashed lines are boundaries for areas when the parameter $IW \neq 0$. The areas have values $\{X(I)\}$ as defined in Eq. (4). (The weighting factors $\{Y(I)\}$ in Eq. (6) are then set equal to $\{X(I)\}$.) The dashed lines are equally spaced between data points with equal intervals of longitude ($= \pi/NO$) and equal intervals of colatitude ($= 2\pi/N1$). The weighting factors $\{Y(I) = X(I)\}$ are easier and faster to calculate by computer than the factors when $IW = 0$. The resulting integration has good numerical stability and gives fairly accurate answers.

26. If the parameter $IW = 0$, then a more exact integrating scheme is used. The weights $\{Y(I)\}$ are still elements of area on the surface of the unit sphere, and they still consist of equal intervals of longitude. But the intervals of colatitude (dotted lines in Figure 7) are varied to make certain surface integrals exact if the integrand consists of a finite number of spherical harmonic terms. The highest degree term that can be contained in the integrand (or highest degree magnet that can be represented by the data $F(I,J)$) and still be exact depends on the parameters NO and $N1$. The relationship of NO and $N1$ to the degree n and order m of the coefficients A_n^m and B_n^m are

$$\begin{aligned} NO &\geq m + 1 \\ N1 &\geq 4m + 1. \end{aligned} \quad (11)$$

Since $m \leq n$, as seen in Eqs. (5) and (7), the coefficients for a multipole magnet of degree n can be approximated accurately only if

$$\begin{aligned} NO &\geq n + 1 \\ N1 &\geq 4n + 1. \end{aligned} \quad (12)$$

This means that the dipole terms ($n = 1$) require that $NO \geq 2$ and $N1 \geq 5$. The following table shows values for several multipole magnets.

TABLE 1 EXAMPLES OF MINIMUM VALUES OF NO AND $N1$

n	$NO \geq$	$N1 \geq$
1 (dipole)	2	5
2 (quadrupole)	3	9
3	4	13
4	5	17
5	6	21
6	7	25
7	8	29
8	9	33

27. It should be noted here that setting $NO, N1 = 2, 5$ will not generally give the dipole term very accurately unless the higher degree terms are all zero. For example, consider a problem which has only dipole and quadrupole terms, and assume

that only the dipole coefficients are to be computed. Then $N_0, N_1 \geq 3$, 9 is required to insure the accuracy of the dipole computations. In general a problem contains an infinite number of terms. The only cases when the expansion in Eq. (7) contains a finite number of terms are when there are only a pure dipole (of insignificant length), a quadrupole, and/or, a finite number of other multipole magnets which are centered at the origin. An offset dipole or a dipole of considerable length (or any offset multipole magnet of degree n) requires an infinite number of terms for representation by Eq. (7). In general, it is best to select N_0 and N_1 as large as possible with

$$N_0 = (N_1 - 1)/2. \quad (13)$$

28. For $IW = 0$, the weights $Y(I)$ are composed of two factors $D(I)$ and C , i.e.,

$$Y(I) = D(I) \cdot C \quad (14)$$

for $I = 1, 2, \dots, N_1$ and $J = 1, 2, \dots, N_0$ (see reference (a)). The constant C represents the equally spaced intervals of longitude with value

$$C = \pi/N_0. \quad (15)$$

The factors $D(I)$ are determined by solving a set of simultaneous linear equations. The equations can be set up in a number of ways since the factors are symmetric with respect to the values of colatitude of $\theta = \pi/2$ and $\theta = \pi$. The method used in SA3024 is to set up and solve the equations for factors representing the intervals between $\theta = 0$ and $\pi/2$, and then to use symmetry to determine the other weights. This procedure involves two different cases based on the odd integer N_1 . Figure 8 shows the intervals of colatitude for two examples; one when $(N_1 + 1)$ is a multiple of 4 and one when it is not. The examples give rise to two different sets of equations for the factors $D(I)$ as follows:

a. Equations when $(N_1 + 1)$ is a multiple of 4

Let

$$\theta(I) \equiv 2\pi(I - 1)/(N_1 - 1) \quad (3a)$$

$$N_3 \equiv \text{largest integer} \leq (N_1 + 3)/4 \quad (16)$$

Then

$$\sum_{J=1}^{N3} D(J) = 1 \quad (17a)$$

$$\sum_{J=1}^{N3} (\cos \theta(J))^{(2I - 2)} \cdot D(J) = 1/(2I - 1) \quad (17b)$$

for $I = 2, 3, \dots, N3$

$$D[(N1 + 3)/2 - J] = D(J) \quad (17c)$$

for $J = 2, 3, \dots, N3$

$$D(N1 + 1 - J) = D(J) \quad (17d)$$

for $J = 1, 2, \dots, (N1 - 1)/2$

$$D[(N1 + 1)/2] = 2D(1) \quad (17e)$$

b. Equations when $(N1 + 1)$ is not a multiple of 4

Let $\theta(I)$ and $N3$ be defined as in Eqs. (3a) and (16). Then

$$\sum_{J=1}^{N3} 2D(J) + D(N3 + 1) = 2 \quad (18a)$$

$$\sum_{J=1}^{N3} (\cos \theta(J))^{(2I - 2)} \cdot D(J) = 1/(2I - 1) \quad (18b)$$

for $I = 2, 3, \dots, N3 + 1$. Equations (17c), (17d), and (17e) remain unchanged.

29. In Eqs. (17a) and (17b) the I and J "subscripts" can be considered to designate the row and column for the coefficient matrix for $N3$ simultaneous linear equations in $N3$ unknowns. This also follows in Eqs. (18a) and (18b) except for an $(N3 + 1) \times (N3 + 1)$ system of linear equations. The system of equations for two examples are given below.

30. Let $N1 = 7$. Then $(N1 + 1) = 8$ is a multiple of 4, and

$$N3 = 2 \leq (7 + 3)/4 = 2.5$$

$$\theta(1), \theta(2) = 0, 60^\circ$$

$$\cos(60) = 1/2.$$

Therefore Eqs. (17a) and (17b) imply that

$$D(1) + D(2) = 1$$

$$D(1) + D(2)/4 = 1/3.$$

These equations are satisfied by

$$D(1) = 1/9, \quad D(2) = 8/9.$$

The remaining equations, (16c), (16d), and (16e), imply that

$$D(1), D(2), \dots, D(7) = 1/9, 8/9, 8/9, 2/9, 8/9, 8/9, 1/9.$$

31. Next, let $N1 = 9$. Then $(N1 + 1) = 10$ is not a multiple of 4, and

$$N3 = 3 \leq (9 + 3)/4 = 3.0$$

$$\theta(1), \theta(2), \theta(3) = 0, 45^\circ, 90^\circ$$

$$\cos(45^\circ) = 1/\sqrt{2}.$$

Therefore, Eqs. (18a) and (18b) imply that

$$2D(1) + 2D(2) + D(3) = 2$$

$$D(1) + D(2)/2 = 1/3$$

$$D(1) + D(2)/4 = 1/5.$$

The solution to these equations is

$$D(1), D(2), D(3) = 1/15, 8/15, 4/5.$$

The remaining equations, (17c), (17d), and (17e), imply that

$$D(1), D(2), \dots, D(9) = 1/15, 8/15, 4/5, 8/15, 2/15, 8/15, 4/5, 8/15, 1/15.$$

32. Reference (a) gives a more theoretical approach to the methods of integration and describes another method for determining the factors $D(I)$ using a polynomial approach. This method was compared with the direct inversion of the Eqs. (16) and (17). The results indicated that the direct inversion method was more stable numerically and took less computer time.

Units

33. All the programs use a mixed system of units. In general, the magnetic units are in the cgs system, e.g., the magnetic field intensity \bar{H} ($f(\theta, \varphi)$ in Eq. (5)) is in gammas (10^{-5} oersted with a permeability of one). The units for a multipole moment of degree n is pole-centimeter ^{n} , e.g., monopole moment is in poles, dipole moment is in pole-centimeters (cm), quadrupole moment is in pole-cm². The unit pole is equivalent to gauss-cm². The unit of length is inches. This is used for the measurement radius R_1 and the multipole position vectors P (for programs SA4024 and SA5024). The unit of gammas is used for the data array $F(I, J)$ and for the error parameters EG and ED (for programs SA4024 and SA5024). Any parameters that represent angles or angular errors are in degrees.

Sample Problems

34. Appendix F contains the input and output data for SA3024 using the same measurement data as in SA2024 (Figures 5 and 6). The file BN3024, used in the execution, is the binary version of SA3024. Only the first 16 binary records (subprograms) are used when executing via INTERCOM. As before, the example was executed in two different ways on the INTERCOM System. The problem was also submitted to BATCH processing via INTERCOM to demonstrate a method of using the plotting option of the program. One of the two identical plots that resulted from the BATCH processing is included in Figure 9. Although the example in Section II of Appendix F was set up to use the GOULD electrostatic plotter, the data was later plotted on the CALCOMP to simplify reproduction problems.

35. Several characteristics of the analysis can be observed by comparing Appendices D and F and Figures 6 and 9. The first characteristic is the data processing. This involves the conversion of the data units and the adjustment of the curves so that all the curves begin and end as near as possible to the beginning and ending of the first curve. The curve adjustment can be observed by comparing the print out of the data tape on page D-1 and the data printed on page D-2. The first points of curves 1 and 2 on the data tape are -504.3 and -902.8. After the data processing these points are -504.3 and -502.8. Another data adjustment in SA3024 has to do with the monopole moment. Since the monopole moment should be zero if the data has absolute accuracy, then the data is adjusted to produce this condition. The procedure involves the computation of the monopole moment, and then, the subtraction of the magnetic field of the monopole component from the data. A comparison of the data printed on page D-2 and on pages F-2 and

F-3 show a difference of about 24 gammas. This appears as the monopole moment of 14.3... printed on page F-2. The data plotted in Figure 9 represents the data that would have been measured if absolute accuracy was attainable.

36. An example of the method of printing out the spherical harmonic coefficients is shown on page F-3. For each n th degree harmonic term, the coefficients A_n^m for $m = 0, 1, \dots, n$ are printed on the first line, and the coefficients B_n^m for $m = 1, 2, \dots, n$ are printed on the second line. It should be noted that in the definitions related to Eqs. (5) and (6), the coefficients A_n^m and B_n^m are coefficients for the Schmidt polynomials and not for the associated Legendre polynomials. The relationship is discussed in more detail in Appendix K.

PROGRAM SA4024

Introduction

37. The program SA4024, listed in Appendix G, was devised to assist in the conduct of error studies relating to the analysis of satellite magnetism. The program is set up to generate data simulating measurements around a specified system of multipole magnets representing the satellite. Simulated errors can be inserted into the generated data to represent position and instrumentation errors. The generated data is written on a data file (DAT024) in the same format as the one used in the data acquisition procedures with the punched paper tape. The data file is generated in a form that can be used with programs SA2024 and SA3024. The program SA4024 also contains the plotting option.

38. Special subprograms used in the program perform such tasks as:

- a. The computation of the magnetic field vector at a remote location from a multipole magnet
- b. The generation of values of associated Legendre polynomials and Schmidt functions
- c. The interpolation and plotting of the magnetic data.

Input Data

39. The input data for SA4024 varies considerably from the preceding programs since it includes specifications for multipole magnets and data errors. The data is entered in the following order:

a. First line of data or data card - format: (4I5)

NO - The number of curves of data (NO is even and ≤ 16 .)

N1 - The number of equally spaced data points per curve from 0 thru 360 degrees colatitude (N1 is odd and ≤ 33 . The first and last data points for each curve correspond to measurements at 0 degrees colatitude.)

NH - The total number of different harmonics (degrees) of multipole magnets to be considered

IP - The parameter that determines whether or not the data is to be interpolated and plotted (IP $\neq 0$ means that the data will be interpolated and plotted.)

b. Second line of data or data card - format: (5F8.4)

R1 - The radius of the measurement sphere in inches

EG - The error (in gamma) to be randomly inserted into the data to represent instrumentation inaccuracies

EA - The angular error (in degrees) to be randomly inserted into the data to represent measurement position errors

ED - The constant error (in gammas) to be inserted into the data to represent offset in the instrumentation. (This will be analyzed as monopole moment.)

PY - The scale factor (gammas/inch) for the y-axis if the data is to be plotted. (If PY = 0.0 a factor will be computed from the data.)

c. Third line of data or data card - format: (7A10)

F9 - The format for reading and printing the spherical coefficients A(I) and B(I), e.g., (1H, 7E10.4)

All of the following data is repeated "NH" times:

d. Next line of data or data card - format: (2I5)

NN - The harmonic number (degree) of the multipole data being read in
(NN = 1 for dipoles, 2 for quadrupole, etc.)

NM - The number of multipoles with harmonic number NN

The following data is repeated "NM" times:

e. Next lines of data or data cards - format: (3F8.4)

P - The position vector (in inches) of the multipole in rectilinear
coordinates (P_x, P_y, P_z .)

f. Next lines of data or data cards - format: F9

A(I) - The spherical coefficients A_{NN}^{I-1} for the multipole of degree
NN where $I = 1, 2, \dots, NN + 1$ (The order of the Ith coefficient
is $I - 1$.)

g. Next lines of data or data cards - format: F9

B(I) - The spherical coefficients B_{NN}^I for the multipole of degree NN
where $I = 1, 2, \dots, NN$ (The order of the Ith coefficient is I .)

(The relationships between spherical and rectilinear coefficients for dipoles and
quadrupoles are given in Eqs. (9) and (10).)

Output Data

40. The program computes data representing measurements of the specified
system of multipole magnets. There are NO curves of data computed, each containing
N1 data points. The computations are made by using Eq. (7). For instance assume
that the system has NM multipoles of degree N and that $f_{nj}(\theta, \varphi)$ represents (as in
Eq. (7)) the normal component of the magnetic field from the jth multipole of
degree N. Then the total magnetic field $f_n(\theta, \varphi)$ for all multipoles of degree N is

$$f_n(\theta, \varphi) = \sum_{j=1}^{NM} f_{nj}(\theta, \varphi) \quad (19)$$

If NH is the total number of different degrees for the system of multipoles then the normal component of the magnetic field for the total system is

$$f(\theta, \varphi) = \sum_{N=1}^{NH} f_n(\theta, \varphi). \quad (20)$$

The program sets up the data array F(I,J) in this manner. (F(I,J) is defined in Eq. (2).) The data array is then written onto the data file DAT024. The file can then be used with either program SA2024 or program SA3024. A data tape can also be made of this file by listing the file under SYSTEM/BASIC with the tape punch unit on.

Sample Problems

41. SA4024 was used to compute the data for several examples. These are listed in Appendix H and in Figures 10 through 24. BN4024 is the binary version of SA4024. When executing the program via INTERCOM only the first nine binary records are used. The first section in Appendix H presents the input and output data for the sample problem illustrated in Figure 5. The instrumentation inaccuracy was assumed to be ± 1.0 gamma by setting EG = 1.0. The positions for the measurements of the data were considered to have inaccuracies of ± 0.2 degrees (EA = 0.2). The set of data curves were assumed to be offset by 25.0 gammas (ED = 25.0). The data curves representing the simulated measurements are listed on page H-3. Section II of Appendix H contains the data file DAT024 that was generated with this data in a format like the data tape. In Section III of Appendix H the sample problem and several other problems were submitted to BATCH. The resulting data was also plotted. The data in Figure 10 for the sample problem is nearly identical to the data in Figure 9. The remaining examples in section III represent the individual and combined data for the first eight spherical harmonics ($n = 1, 2, \dots, 8$) of the sample problem. These will be discussed in more detail in later paragraphs.

42. SA4024 was also used to show data curves for individual components of dipole and quadrupole moments. These are included in Figures 11 through 24. A small diagram is also included on each figure to illustrate the particular component that is represented by the curves. Curves are shown for components in both spherical and rectilinear coordinates according to Eqs. (9) and (10).

PROGRAM SA5024Introduction

43. The program SA5024, listed in Appendix I, was devised to simplify the error studies. It performs functions similiar to both programs SA3024 and SA4024, i.e., SA5024 is both a data generation and a data analysis program. A system of multipole magnets is specified in the input data to represent the satellite as in SA4024. Since the data is analyzed as it is generated it is not written onto a data file as in SA4024. However, the program does include the plotting option when executing in the BATCH mode. Also, the integration option discussed in paragraphs 25 through 32 is included. The primary subprograms in SA5024 perform most of the functions already described in the sections on SA3024 and SA4024.

Input Data

44. The input data for SA5024 resembles the input data for the program SA4024 since it consists mostly of specifications for multipole magnets and data errors. The data is entered in the following order:

a. First line of data or data card - format: (7I5)

NO - The number of curves of data (NO is even and ≤ 16 .)

N1 - The number of equally spaced data points per curve from
0 through 360 degrees colatitude (N1 is odd and ≤ 33 . The first
and last data points for each curve correspond to measurements
at 0 degrees colatitude.)

NH1 - The total number of different harmonics (degrees) of multipole
magnets to be considered

NH2 - The harmonic number (degree) representing the highest degree
spherical harmonic term to be computed from the data (NH2 = 1
for dipoles, 2 for quadrupoles, etc.)

IP1 - The parameter that determines whether or not the data is to be
interpolated and plotted (IP1 $\neq 0$ means that the data will be
interpolated and plotted.)

IP2 - The parameter that determines whether or not the magnetic data is to be printed (IP2 \neq 0 means that the data will be printed.)

IW - The parameter that determines which integrating scheme is to be used (IW = 0 means that the exact, algebraic scheme is to be used.)

b. Second line of data or data card - format: (5F8.4)

R1 - The radius of the measurement sphere in inches

EG - The error (in gammas) to be randomly inserted into the data to represent instrumentation inaccuracies

EA - The angular error (in degrees) to be randomly inserted into the data to represent measurement position errors

ED - The constant error (in gammas) to be inserted into the data to represent offset in the instrumentation (this will be analyzed as monopole moment)

PY - The scale factor (gammas/inch) for the y-axis if the data is to be plotted (if PY = 0.0 a factor will be computed from the data.)

c. Third line of data or data card - format: (7A10)

F9 - The format for reading and printing the spherical coefficients A(I) and B(I), e.g., (1H ,7E10.4)

All of the following data is repeated "NH" times:

d. Next line of data or data card - format: (2I5)

NN - The harmonic number (degree) of the multipole data being read in (NN = 1 for dipoles, 2 for quadrupole, etc.)

NM - The number of multipoles with harmonic number NN

The following data is repeated "NM" times:

e. Next lines of data or data cards - format: (3F8.4)

P - The position vector (in inches) of the multipole in rectilinear coordinates (P_x, P_y, P_z)

f. Next lines of data or data cards - format: F9

A(I) - The spherical coefficients A_{NN}^{I-1} for the multipole of degree NN where $I = 1, 2, \dots, NN + 1$ (The order of the Ith coefficient is $I - 1$.)

g. Next lines of data or data cards - format: F9

B(I) - The spherical coefficients B_{NN}^I for the multipole of degree NN where $I = 1, 2, \dots, NN$. (The order of the Ith coefficient is 1.).

(The relationships between spherical and rectilinear coefficients for dipoles and quadrupoles are given in Eqs. (9) and (10).)

Output Data

45. Initially, the program SA5024 computes the simulated measurement data as in Eqs. (19) and (20). Next, the spherical harmonic coefficients are computed according to Eq. (6). The program also computes the exact dipole moment from the multipole specifications. This value is compared with the dipole moment computed from the simulated measurement data. The percent error is printed as part of the output data.

Sample Problems

46. Appendix J contains the input and output data for several problems using SA5024. As in the other programs the binary version of SA5024 is the file BN5024. When executing the program via INTERCOM, only the first 15 binary records are used. In the first section of Appendix J an analysis was conducted of the sample problem illustrated in Figure 5. The parameters EG, EA, and ED were all set to zero (page J-1) so that no simulated errors were inserted into the data.

47. Figure 25 shows eight curves of the computed data for the three dipoles in the sample problem. Actually, 16 curves, each containing 33 data points, were used in the computations (based on Eq. (12) and Table 1). These values for the

parameters NO and N1 allowed the computation of the spherical harmonic coefficients for the first eight multipole components. The program SA4024 was used to compute eight data curves for each multipole component to demonstrate the type of curves representing each component. These are presented in Figures 26 through 33. These were also combined and plotted by SA4024. Figure 34 displays the combined data. The first data curve (xz-plane) was chosen to further demonstrate the characteristics of spherical harmonic approximation. Figure 35 contains four curves. The first curve is the original data from the sample problem. Curve 2 represents the dipole component. Curve 3 represents the combination of dipole and quadrupole components. Curve 4 represents the combination of the first eight multipole components. The convergence of the series in Eq. (7) is not obvious from the coefficients computed in Section I of Appendix J. Table 2 summarizes the peak coefficients and the data peaks for each degree n. Although the magnitude of the peak coefficient increases with increasing n, the data peak (magnetic field at 96 inches) decreases. In fact, the first eight multipole components account for about 97 or 98 percent of the original data curves. This indicates that higher degree components are relatively insignificant.

TABLE 2
PEAK DATA FOR THE MULTIPOLE COMPONENTS OF THE SAMPLE PROBLEM

Degree n	(Max { $ A_n^m $, $ B_n^m $ }) $0 \leq m \leq n$	Data Peak (for 8 curves)
1	10^4	13.79
2	$-.9472 \times 10^7$	803.79
3	$.2341 \times 10^{10}$	839.34
4	$.1545 \times 10^{12}$	367.51
5	$-.2812 \times 10^{14}$	211.03
6	$+.4533 \times 10^{16}$	189.44
7	$+.3541 \times 10^{18}$	57.51
8	$-.4517 \times 10^{20}$	35.47
Original		1817.94
Sum of Components		1786.61

48. Section II of Appendix J demonstrates the procedure for submitting the program SA5024 for BATCH execution. The data specifies two problems consisting of

offset dipoles. In the first problem the dipole is aligned parallel to the x-axis and offset along the x-axis. The resultant curves are presented in Figure 36. The second problem is a dipole aligned parallel to the y-axis and offset along the x-axis. These curves are presented in Figure 37. A small diagram is included in each figure to demonstrate the geometric configuration.

49. A small error study was made to demonstrate the relationship between the accuracy of the dipole analysis; the parameters N_0 , N_1 , and R_1 ; and the measurement errors (represented by EG and EA). The sample problem in Figure 5 was again selected. This example has a resultant dipole moment of 1000 gauss-cm^3 , but the magnetic field from this dipole component is embedded in the large field from quadrupole and higher terms. Curve 2 in Figure 35 demonstrates this condition. Also Table 2 shows that the peak of the dipole curves is only 13.79 gammas out of a total of 1817.94 gammas for the total system. Table 3 contains a summary of the study. Several different values of N_0 , N_1 , IW , R_1 , EG and EA were used. The indications are that increasing R_1 doesn't help much if EA and EG are too large. Also, there isn't much improvement in the accuracy of the computations when using the exact integration scheme (i.e., when setting $IW = 0$). Probably the most significant observation is that the parameter EA , representing position errors, has a much greater effect on the analysis accuracy than the parameter EG , representing instrumentation errors.

TABLE 3
MAXIMUM ERROR VERSUS PARAMETRIC VARIATIONS IN COMPUTING THE DIPOLE MOMENT FOR THE SAMPLE PROBLEM

Dipole Moment Error (gauss-cm³)

NO, NL	Parameters IW, RL, EG, EA						
	1,96,0,0	1,96,1,1	1,96,5,2	1,72,1,1	1,72,5,2	1,120,1,1	0,72,1,1
6, 13	-128 (x)	-135 (y)	-239 (y)	-2450 (x)	-2510 (x)	868 (x)	-2430 (x)
8, 17	-.63 (x)	653 (y)	1320 (y)	1760 (y)	3590 (x)	395 (y)	1780 (y)
4, 25	431 (x)	887 (x)	1470 (y)	7420 (x)	8390 (x)	500 (y)	7440 (x)
6, 25	-54 (x)	-734 (y)	1440 (y)	-1160 (y)	-2170 (y)	-462 (y)	-1160 (y)
8, 25	6.3 (x)	-407 (x)	-785 (x)	220 (y)	-505 (x)	-298 (x)	223 (y)
10, 25	-2.8 (x)	-515 (x)	-1030 (x)	-1080 (x)	-2200 (x)	-306 (x)	-1080 (x)
12, 25	-2.7 (x)	-884 (x)	-1730 (x)	-1620 (x)	-3250 (x)	-570 (x)	-1620 (x)
Minimum detectable dipole based on RL & EG		73	363	31	153	142	31

NO, NL, IW, RL = 12, 25, 1, 96

Note: An error of -1620 gauss-cm³ means that the program computed a dipole moment of -620 gauss-cm³ instead of +1000 gauss-cm³. The letter contained in parentheses indicates the dipole component containing the error.

EG	ERROR	EA	ERROR
8	-66 (x)	4	1003 (y)
4	23 (x)	2	-854 (x)
2	-6.8 (x)	1	505 (y)
1	-10.9 (y)	.5	362 (x)
.5	-2.3 (y)	.25	-197 (x)
.25	-2.8 (x)	.125	56 (x)
0	-2.7 (x)	0	-2.7 (x)

Chapter 4

PRINCIPAL SUBPROGRAMS

INTRODUCTION

50. The programs described in the preceding sections were constructed to utilize several algorithms which were incorporated into functions and subroutines. These include algorithms for computing the spherical harmonic coefficients, computing the vector magnetic field emanating from a multipole magnet, computing the associated Legendre Polynomials, computing the Schmidt polynomials, and interpolating and plotting data curves. The subprograms have been code in a manner which makes them useful as building blocks for other programs. They are discussed separately in the following paragraphs.

SUBROUTINE AMPMNT

51. The subroutine AMPMNT is used to compute the spherical coefficients A_n^m and B_n^m for the data array $F(I,J)$. The computation is based on Eq. (6). The subroutine is constructed to compute the coefficients for only one degree at a time. The calling sequence is

CALL AMPMNT (NO, N1, N, R1, F, P, A, B)

where the arguments are defined as follows:

a. Input

NO - The number of curves of data

N1 - The number of equally spaced data points per curve from
0 through 360 degrees colatitude

N - The degree of the spherical harmonic coefficients to be computed

R1 - The radius of the measurement sphere in inches

$F(IJ)$ - The measured data for $IJ = I + (J - 1) \cdot N1$, $I = 1, 2, \dots, N1$, and $J = 1, 2, \dots, N0$ (The data is stored in a one-dimensional array to conserve storage, i.e., equating the one-dimensional array to the two-dimensional array in Eq. (6) gives $F(IJ) \equiv F(I, J)$.)

$P(I)$ - The first $(N1 + 1)/2$ weighting factors $Y(I)$ in Eq. (6), i.e., $P(I) = Y(I)$ for $I = 1, 2, \dots, (N1 + 1)/2$.

b. Output

$A(M + 1)$ - The spherical coefficients A_n^m for $M = 0, 1, 2, \dots, N$ (see Eq. (8))

$B(M)$ - The spherical coefficients B_n^m for $M = 1, 2, 3, \dots, N$ (see Eq. (8)).

52. There are several other subroutines that are used to support AMPMNT in addition to the standard FORTRAN IV machine routines. AMPMNT internally calls a subroutine labeled POLVAL which in turn uses functions labeled SPNM, PNM, and ANF. These routines are used to compute the appropriate values for the Schmidt polynomials $P_n^m(\cos(\theta(I)))$ in Eq. (6). Also, the weights $P(I)$ are set up externally by special subroutines before being input to the AMPMNT subroutine. This process requires the use of subroutines labeled WGT1, WGT2, WGT3, and GAUSEL. Listings of all of these subroutines are included in Appendix I.

SUBROUTINE AMPFLD

53. The subroutine AMPFLD is used to compute the magnetic field at the point \bar{R} from a multipole magnet of degree n centered at the point \bar{P} . The magnet is specified in terms of the spherical harmonic coefficients A_n^m and B_n^m . The calling sequence is

CALL AMPFLD (R, P, N, A, B, F)

where the arguments are defined as follows:

a. Input

$R(I)$ - The rectilinear coordinates (in inches) of the point at which the field is to be computed ($R(1) = x\text{-coord.}$, $R(2) = y\text{-coord.}$, and $R(3) = z\text{-coord.}$)

$P(I)$ - The rectilinear coordinates (in inches) of the center position of the multipole magnet ($P(1) = x\text{-coord.}$, $P(2) = y\text{-coord.}$, and $P(3) = z\text{-coord.}$)

N - The harmonic degree of the multipole magnet

$A(M + 1)$ - The spherical coefficients A_n^m for $M = 0, 1, 2, \dots, N$ (see Eq. (8))

$B(M)$ - The spherical coefficients B_n^m for $M = 1, 2, \dots, N$ (see Eq. (8))

b. Output

$F(I)$ - The total magnetic field in rectilinear coordinates ($H_x = F(1)$, $H_y = F(2)$, and $H_z = F(3)$).

54. The subroutine computes the total magnetic field vector for a single multipole magnet. The computations are based on relationships that are similar to Eq. (7) except that only a single multipole magnetic (N is fixed) is considered. Also, the subroutine gives all three rectilinear components of the magnetic field instead of just the radial component. The relationships are presented below using terminology that can be found in references (b) and (c). Let \bar{R}_1 be defined as the vector from the magnet to the point \bar{R} , i.e.,

$$\bar{R}_1 = \bar{R} - \bar{P} \quad (21)$$

where \bar{R} and \bar{P} are as defined above. Next, let

$$\bar{R}_1 = (R_1, \varphi, \theta) \quad (22)$$

be the representation of \bar{R}_1 in spherical coordinates where R_1 is the vector length, θ is the colatitude, and φ is the longitude relative to a system of coordinates centered at the multipole position. Now, define an orthonormal set of vectors \bar{r} , $\bar{\theta}$, and $\bar{\varphi}$ in spherical coordinates as

$$\bar{r} = \cos \varphi \sin \theta \bar{i} + \sin \varphi \sin \theta \bar{j} + \cos \theta \bar{k} \quad (23a)$$

$$\bar{\theta} = \cos \varphi \cos \theta \bar{i} + \sin \varphi \cos \theta \bar{j} - \sin \theta \bar{k} \quad (23b)$$

$$\bar{\varphi} = -\sin \varphi \bar{i} + \cos \varphi \bar{j} \quad (23c)$$

where \bar{i} , \bar{j} , and \bar{k} represent the orthonormal set of vectors along the x, y, and z-axis, respectively (see Figure 1). It can be seen that

$$\bar{r} = \bar{r}_1 / R_1. \quad (24)$$

Then, the total field vector $\bar{H}_n(\bar{R})$ for the n th degree multipole is given as

$$\bar{H}_n(\bar{R}) = H_{nr}(\bar{R})\bar{r} + H_{n\theta}(\bar{R})\bar{\theta} + H_{n\phi}(\bar{R})\bar{\phi} \quad (25)$$

where

$$H_{nr}(\bar{R}) = [(n+1)/R_1^{n+2}] \sum_{m=0}^n (A_n^m \cos m\phi + B_n^m \sin m\phi) P_n^m(\cos \theta) \quad (26a)$$

$$H_{n\theta}(\bar{R}) = (1/R_1^{n+2}) \sum_{m=0}^n (A_n^m \cos m\phi + B_n^m \sin m\phi) [C_n^m P_n^{m+1}(\cos \theta) - m \cos \theta P_n^m(\cos \theta) / \sin \theta] \quad (26b)$$

$$H_{n\phi}(\bar{R}) = (1/R_1^{n+2}) \sum_{m=0}^n m (A_n^m \sin m\phi - B_n^m \cos m\phi) P_n^m(\cos \theta) / \sin \theta. \quad (26c)$$

The coefficient C_n^m , in Eq. (26b), has the value

$$C_n^m = [(n-m)(n+m+1)]^{\frac{1}{2}} \text{ for } (m > 0) \\ = [n(n+1)/2]^{\frac{1}{2}} \text{ for } (m = 0).$$

Also, the terms which have $\sin \theta$ as a divisor are computed using the following identities:

$$2m P_{n,m}(\cos \theta) / \sin \theta = (n-m+1)(n-m+2) P_{n+1,m-1}(\cos \theta) + P_{n+1,m+1}(\cos \theta) \quad (27) \\ \text{for } (m > 0) \\ = 0 \text{ for } (m = 0)$$

$$P_n^m(\cos \theta) = \left\{ 2 \frac{(n-m)!}{(n+m)!} \right\}^{\frac{1}{2}} P_{n,m}(\cos \theta) \text{ for } (m > 0) \quad (28) \\ = P_{n,0}(\cos \theta) \text{ for } (m = 0).$$

The function $P_{n,m}(\cos \theta)$ represents the associated Legendre polynomial of degree n and order m . $P_n^m(\cos \theta)$ represents the Schmidt polynomial of degree n and order m . These polynomials are discussed in more detail in Appendix K. The radial component $f_n(\varphi, \theta)$ of the magnetic field at the point \bar{R} is computed from the total field $\bar{H}_n(\bar{R})$ as

$$f_n(\varphi, \theta) = \bar{H}_n(\bar{R}) \cdot \bar{R} / (\bar{R} \cdot \bar{R})^{\frac{1}{2}}. \quad (29)$$

55. The subroutine AMPFLD uses several external subprograms including SPNM, PNM, ANF, SPCOOR, and SUM. These are listed in Appendix I. The subprograms SPNM, PNM, and ANF are used to compute the Schmidt and associated Legendre polynomials. The subroutines SPCOOR and SUM are used in mathematical operations with vectors. SPCOOR transforms a vector from rectilinear to spherical coordinates. SUM multiplies vectors and scalars, and then, adds the products.

FUNCTIONS PNM, SPNM, AND ANF

56. The function subprograms PNM, SPNM, and ANF are used in the generation of values for spherical polynomials. PNM computes the value of the associated Legendre polynomial $P_{n,m}(x)$ of degree n and order m at the point x for $|x| \leq 1$. (If $m = 0$ the subprogram computes the value of the regular Legendre polynomial.) SPNM converts this value into the Schmidt function $P_n^m(x)$. ANF simply computes the factorial value for the integer N . Let y represent a variable used in a computer program which is to be set equal to a polynomial value. Then the calling sequence for the subprograms is defined as follows:

$y = P_n(x)$ is written as $Y = \text{PNM}(N, 0, X)$ for $N \geq 0$ and $|X| \leq 1$

$y = P_{n,m}(x)$ is written as $Y = \text{PNM}(N, M, X)$ for $N \geq 0$, $M \geq 0$, and $|X| \leq 1$

$y = P_n^m(x)$ is written as $Y = \text{SPNM}(N, M, X)$ for $N \geq 0$, $M \geq 0$, and $|X| \leq 1$.

The factorial subprogram is used as follows:

$y = n!$ is written as $Y = \text{ANF}(N)$ for $N \geq 0$.

The development and use of these subprograms was reported in more detail in an internal technical note. This is included in Appendix K.

SUBROUTINE DATPLT

57. The subroutine DATPLT is used to interpolate and plot the curves of data. It is written to make use of either the GOULD electrostatic plotter or the CALCOMP 570 pen plotter. Most of the data curves in the enclosed figures were plotted on the CAPCOMP plotter. Figure 38 shows a set of data curves plotted on the GOULD plotter. These correspond to the curves in Figure 6. The routine sets up x and y-coordinate arrays for the magnetic field data $F(IJ)$. A plotting symbol (the curve number) is used to mark each data point along the data curves. The data is plotted as magnetic field versus the angle of colatitude as shown in Figure 6. The calling sequence is

CALL DATPLT (NO, N1, F, YDIST)

where the argument list is defined of follows:

NO - The number of curves of data

N1 - The number of data points per curve

$F(IJ)$ - The array of data curves for $IJ = J + (I - 1) \cdot N1$;
 $J = 1, 2, \dots, N1$; and $I = 1, 2, \dots, NO$

YDIST - The value in gamma to be used for one inch of the y-axis (If YDIST = 0.0 the subroutine PKY is used to determine a suitable scale for the y-axis.).

58. The subroutine uses the external subprograms CALCML, FNCTON, and PKY for execution. CALCML is contained in special plotting packages of subprograms that is in the NOL subroutine library. There are two separate packages that can be loaded when using the subroutines DATPLT and CALCML. The package labeled GOULD1 is loaded when plotting with the GOULD electrostatic plotter. The package labeled CALCML is loaded when plotting with the CALCOMP plotter. Also a tape (TAPE99) must be loaded with the CALCML package. The subprograms FNCTON and PKY are listed in Appendix I. FNCTON computes the Fourier coefficients for each curve of data. These are used to increase the point density of the data to be plotted to one point for every five degrees of colatitude. This turns out to be 12 points per inch on the graphs or a total of 73 points per curve. The subprogram PKY is used to determine a suitable scale for the y-axis if YDIST is zero.

SUBROUTINE FNCTON

59. The subroutine FNCTON is used to perform a Fourier analysis on a single-dimensioned array of data that represents the values along a curve at equally spaced intervals. The program is used primarily for interpolating data. The calling sequence is

CALL FNCTON (F, N, C, CO, PHI, N2, DEV)

where the argument list is defined as follows:

a. Input

F(I) - The data array representing the values of the curve at N equally space values $\theta(I) = 2\pi(I - 1)/N$ for $I = 1, 2, \dots, N$

N - The number of data points

b. Output

C(I) - The array of Fourier amplitudes for $I = 1, 2, \dots, N2$

CO - The zero degree Fourier amplitude

PHI(I) - The array of Fourier phase angles for $I = 1, 2, \dots, N2$

N2 - The number of terms in the approximating Fourier series

DEV - The maximum difference between the data points and the curve defined by the Fourier series.

60. The coefficients are obtained by the following equations:

$$C_0 = (1/N) \sum_{I=1}^N F(I) \quad (30)$$

$$S1_j = (2/N) \sum_{I=1}^N F(I) \cdot \sin [2\pi j(I - 1)/N] \quad (31a)$$

$$C1_j = (2/N) \sum_{I=1}^N F(I) \cdot \cos [2\pi j(I - 1)/N] \quad (31b)$$

$$\begin{aligned}
 C_j &\equiv (S1_j^2 + C1_j^2)^{\frac{1}{2}} \text{ for } j \neq N/2 \\
 &\equiv (S1_j^2 + C1_j^2)^{\frac{1}{2}}/2 \text{ for } j = N/2
 \end{aligned}
 \tag{32}$$

$$PHI_j \equiv \tan^{-1}(C1/S1) \tag{33}$$

for $j = 1, 2, \dots, N/2 + 1$.

61. This results in the approximation

$$F(I) \approx C_0 + \sum_{j=1}^{N/2} C_j \sin [2\pi j(I - 1)/N + PHI_j] \text{ for } I = 1, 2, \dots, N. \tag{34}$$

Any other set of N' equally spaced values which begin at $\theta(1) = 0$ can be computed by substituting N' in place of N in Eq. (34). DATPLT uses the value of 72 for N' .

Chapter 5

CONCLUSIONS

62. The computer programs presented in this report are part of a major effort at NOL to develop and improve magnetic test procedures on spacecraft. The effort involves the development and evaluation of both data acquisition and data analysis procedures. The Spherical Field Coil Facility at NOL is being instrumented to handle both the acquisition and analysis of the test data. The proposed specifications for the instrumentation were presented in reference (d). More detailed specifications for the data acquisition and analysis system were forwarded to NRL as an enclosure to reference (e). The facility is scheduled to be operational by the beginning of fiscal year '75. A technical report describing the facility should be published shortly thereafter.

63. There are other analysis projects to be completed. One is the compilation and publication of a reference manual containing the mathematical equations for dipole and quadrupole computations. This will simplify slide-rule calculations which are made during testing. Another project involves the conduct of a more elaborate error analysis using the computer programs described above. This will be necessary to evaluate the minimum detectable dipole moment under several different conditions.

64. It should be noted here that there are several other methods for estimating the dipole moment of spacecraft. The most effective method to use in any particular case depends on a number of things including the size of the spacecraft, the availability of a gimbaling system, the desired accuracy of the compensation, the complexity and permanence of the spacecraft's magnetism, etc. The procedure based on the spherical harmonic analysis of the spacecraft's magnetism has proven to be both fast and accurate for NRL satellites. Many of the tests have been completed in less than four hours and with an accuracy which allows dipole moment compensation of all but about 100 out of 6000 gauss-cm³.

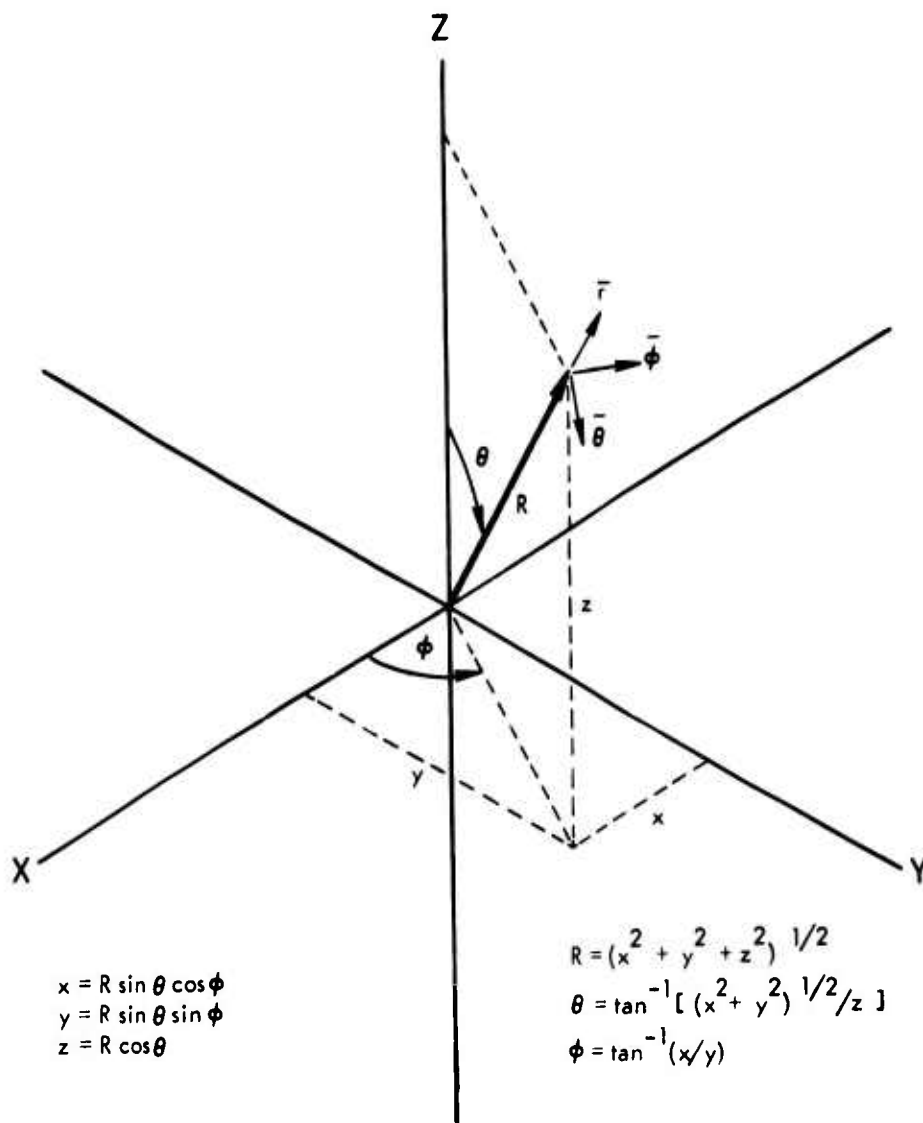


FIG. 1 ILLUSTRATION OF SPHERICAL POLAR COORDINATES

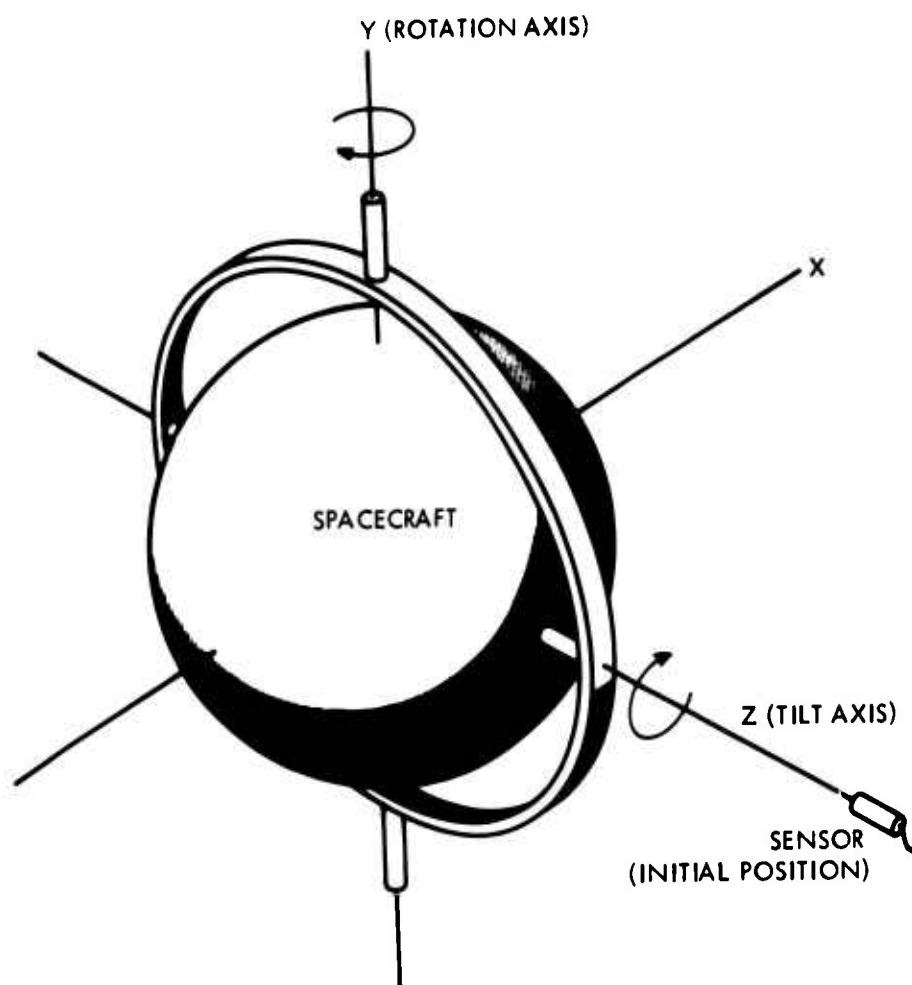


FIG. 2 DEGREES OF FREEDOM FOR MEASUREMENTS

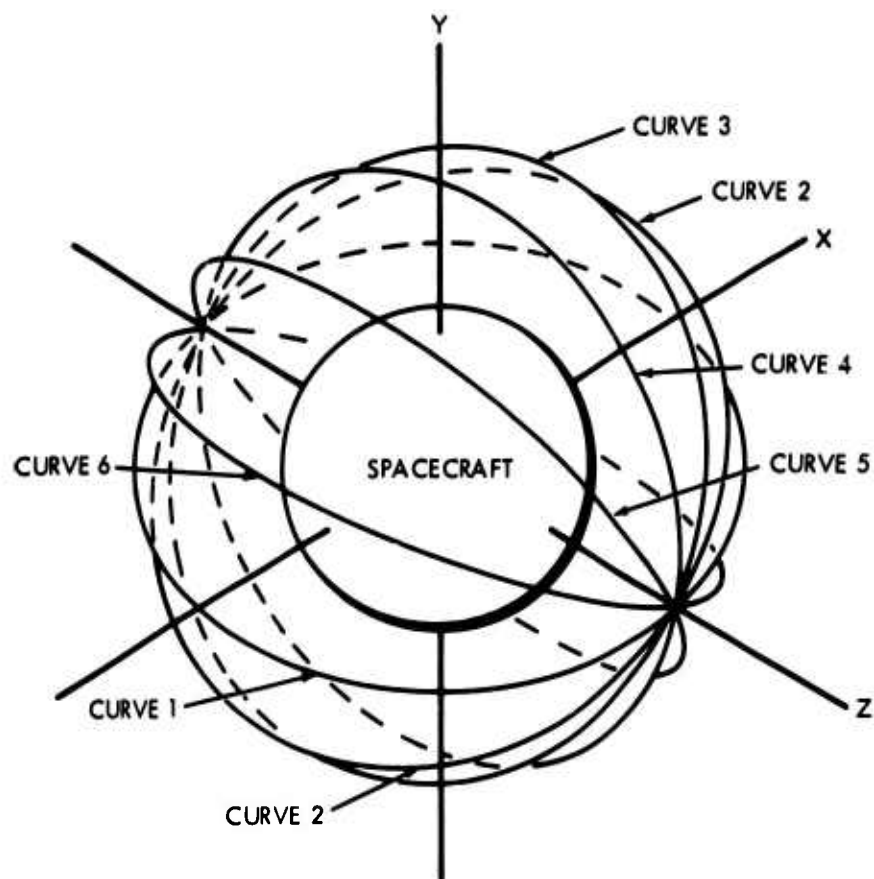
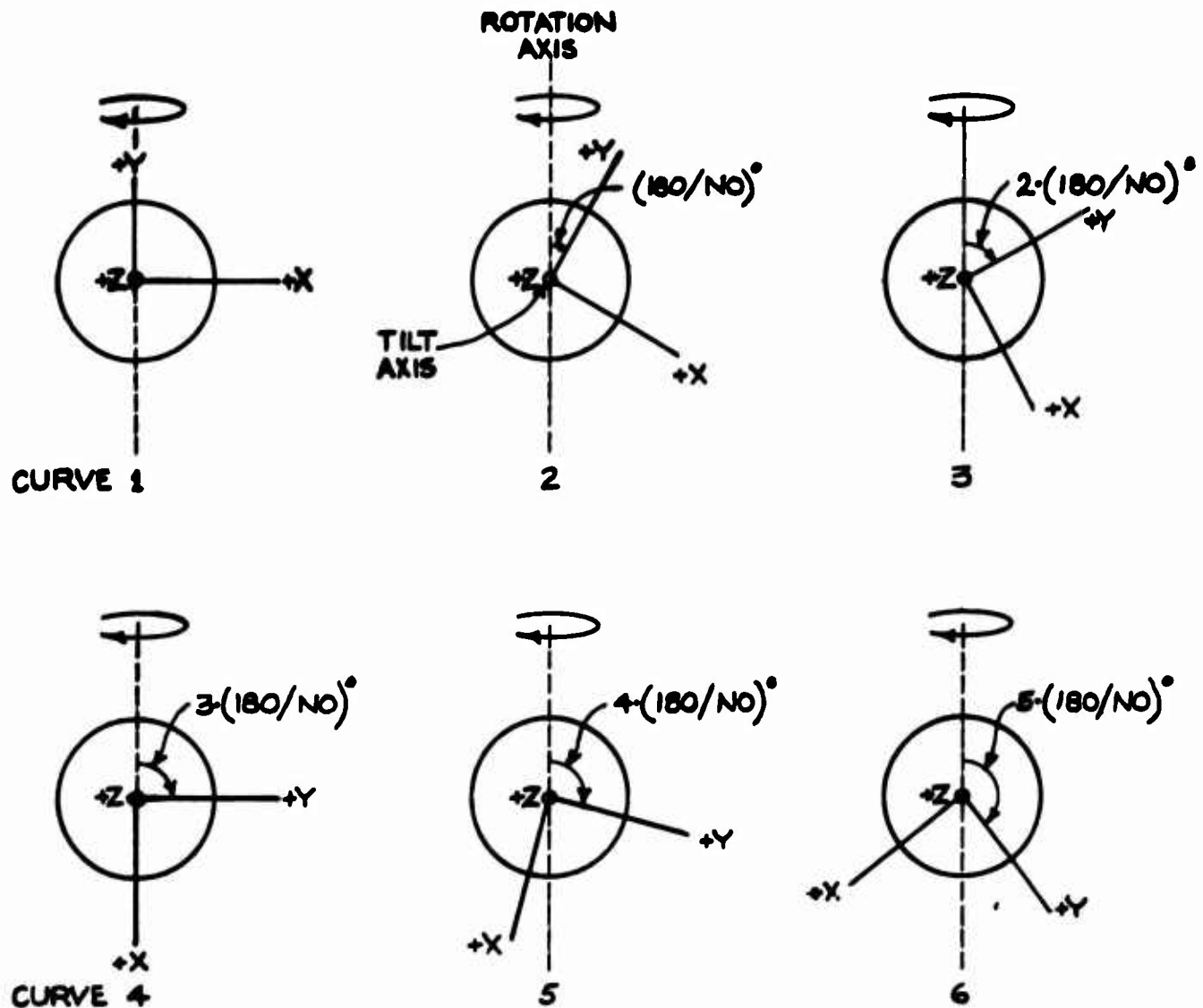


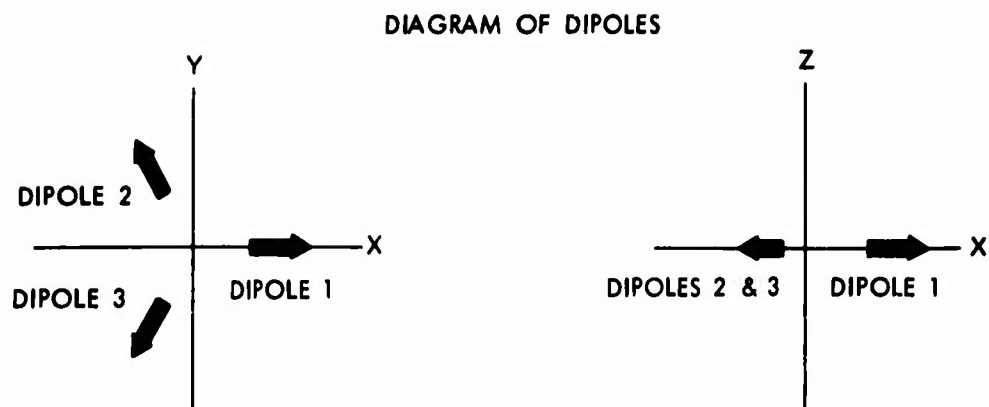
FIG. 3 SPHERE OF MEASUREMENTS



Notes:

1. The diagrams are viewed from the sensor position.
2. The turntable rotation is clockwise.
3. Each curve begins with +z-axis directed at the sensor.

FIG. 4 SATELLITE INITIAL POSITIONS FOR NO = 6
(Viewed from sensor position)



PARAMETRIC VALUES

PARAMETER	COMP.	DIPOLE		
		1	2	3
POSITION VECTOR	X	39.000	-21.000	-21.000
	Y	0.000	36.373	-36.373
	Z	0.000	0.000	0.000
DIPOLE MOMENT	$D_x (A(2))$	31,000.	-15,000.	-15,000.
	$D_y (B(1))$	0.	25,980.	-25,980.
	$D_z (A(1))$	0.	0.	0.

FIG. 5 ILLUSTRATION OF SAMPLE PROBLEM

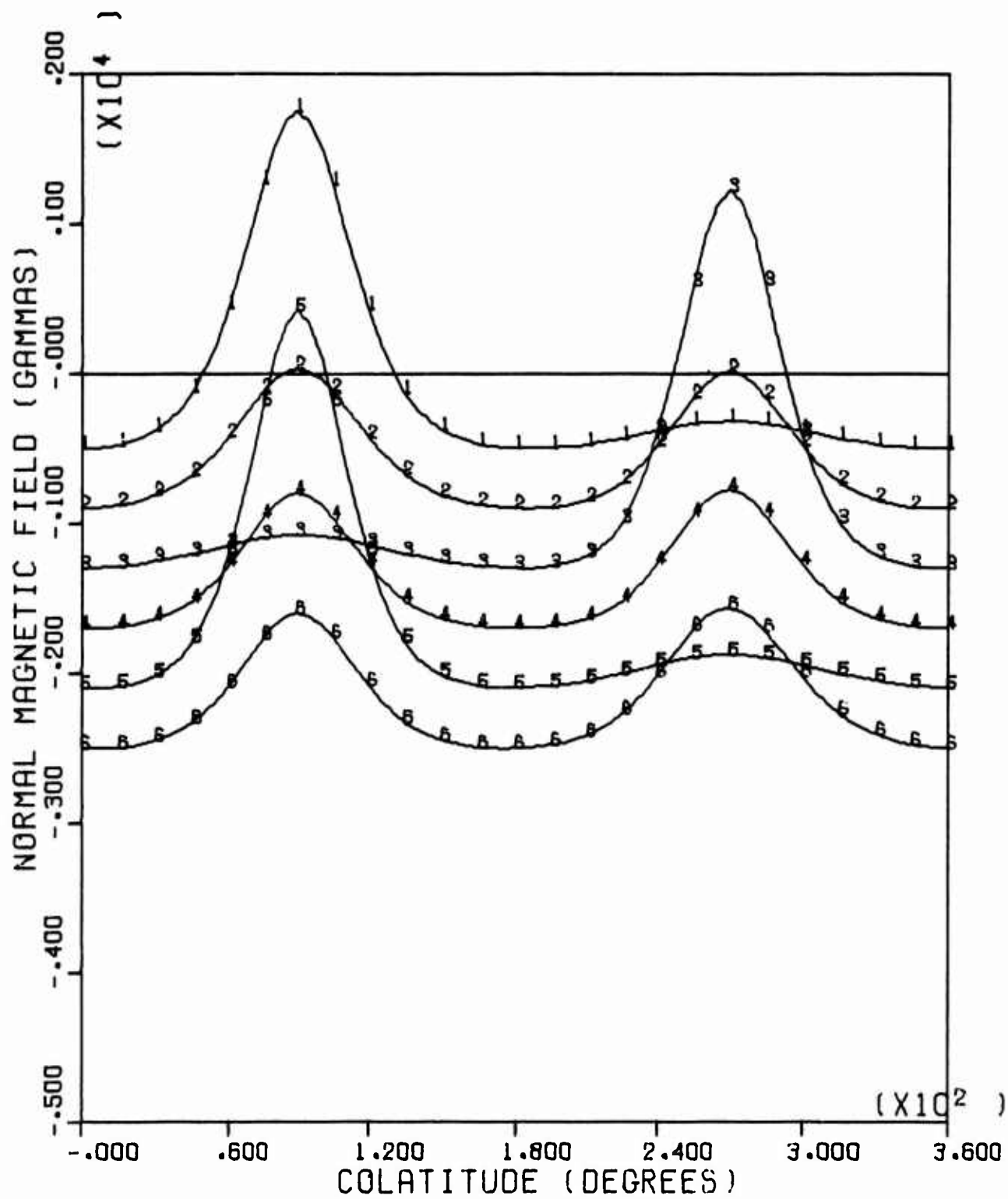


FIG. 6 TYPICAL DATA CURVES

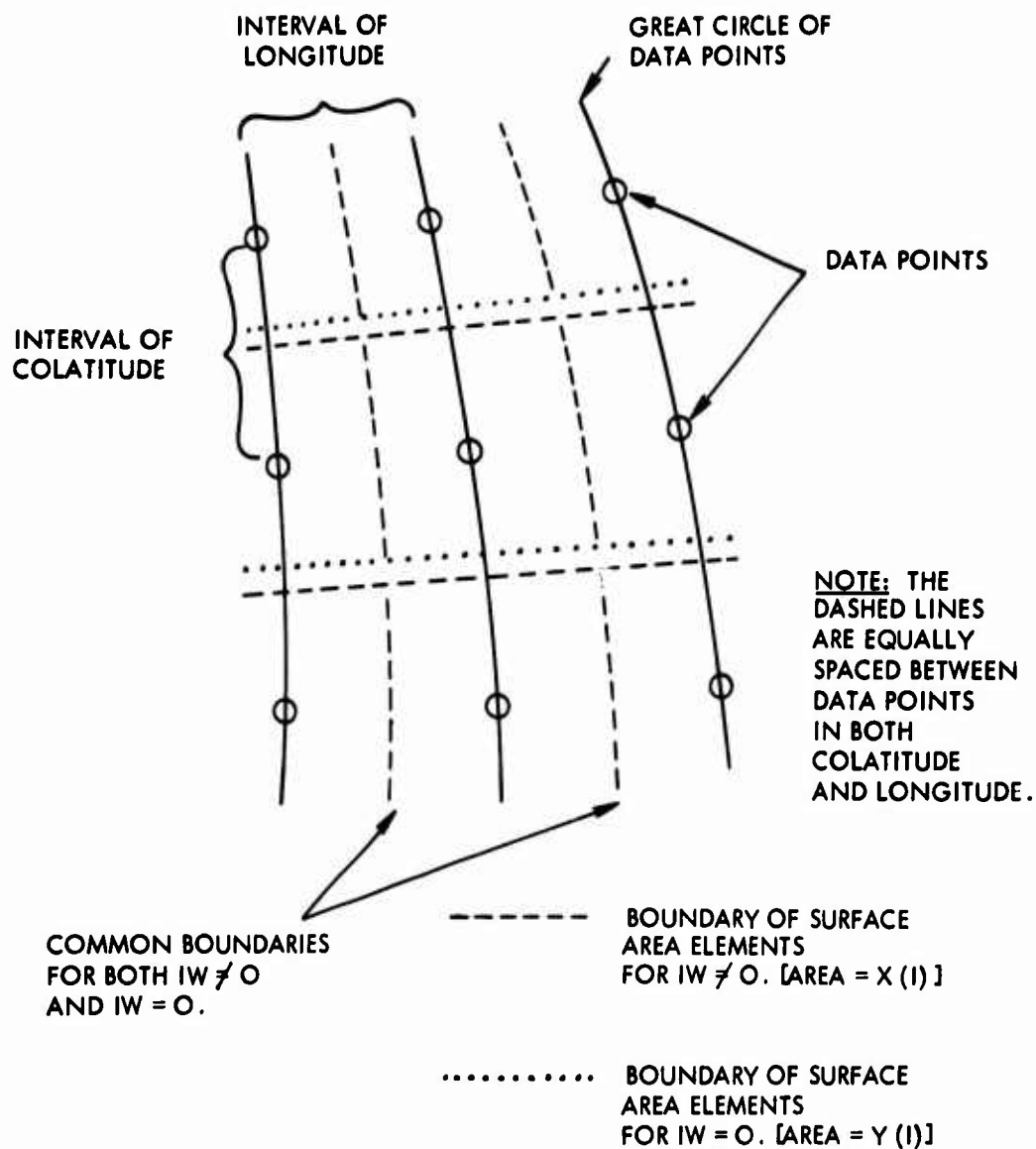


FIG. 7 EXAMPLE OF SURFACE AREA ELEMENTS ASSIGNED TO DATA POINT

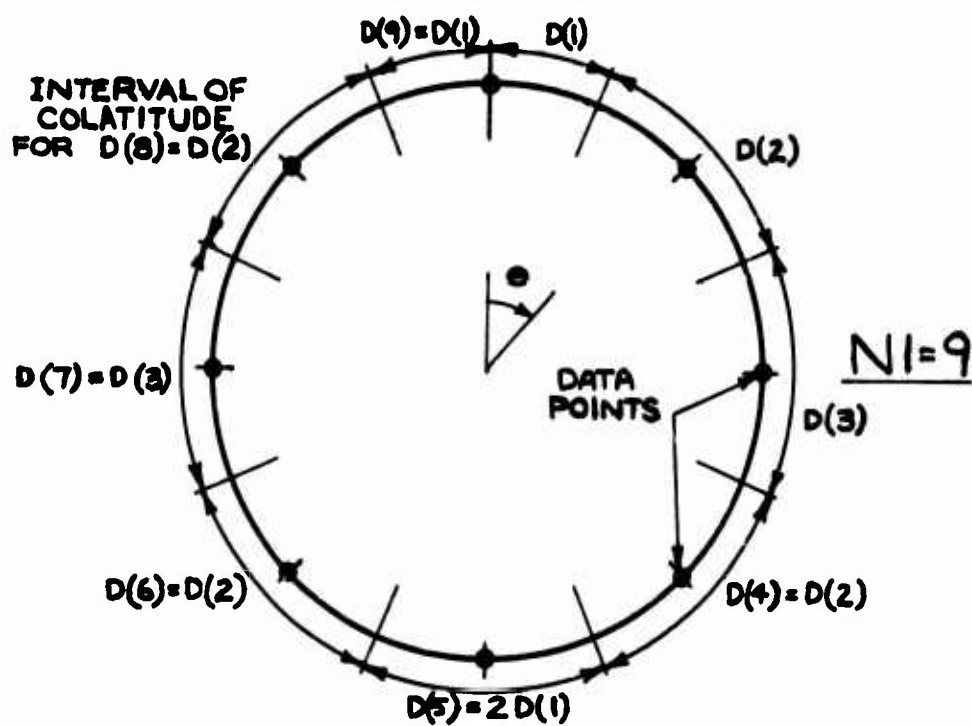
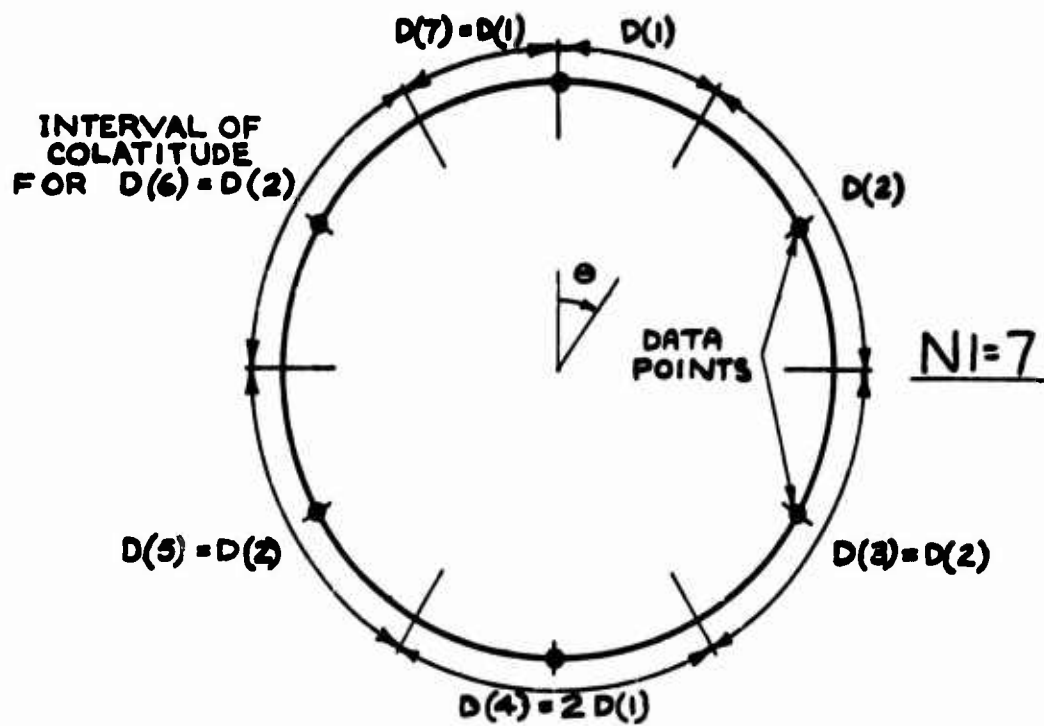


FIG. 8 EXAMPLES OF INTERVALS OF COLATITUDE

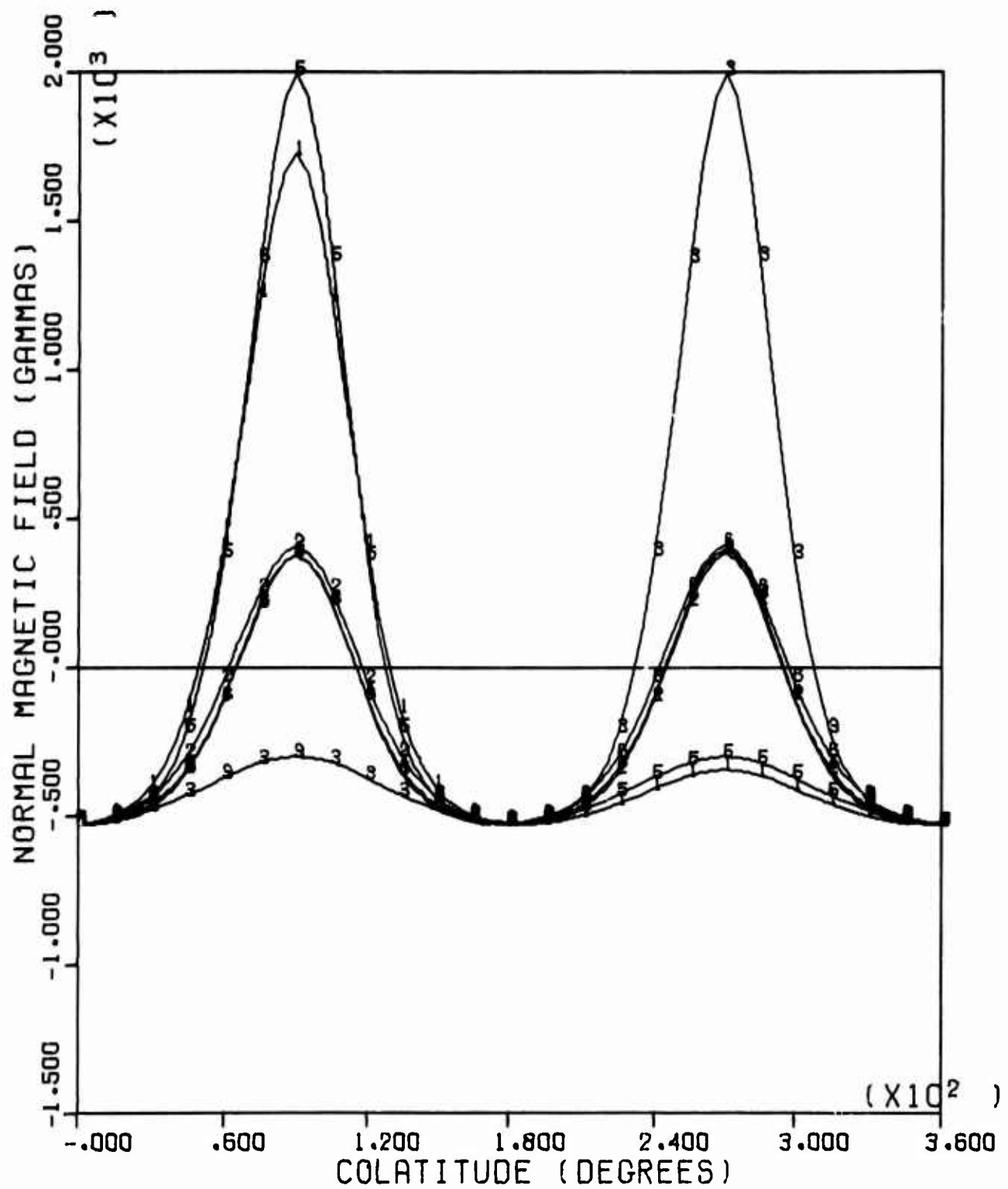


FIG. 9 DATA CURVES FROM SA3024 FOR SAMPLE PROBLEM

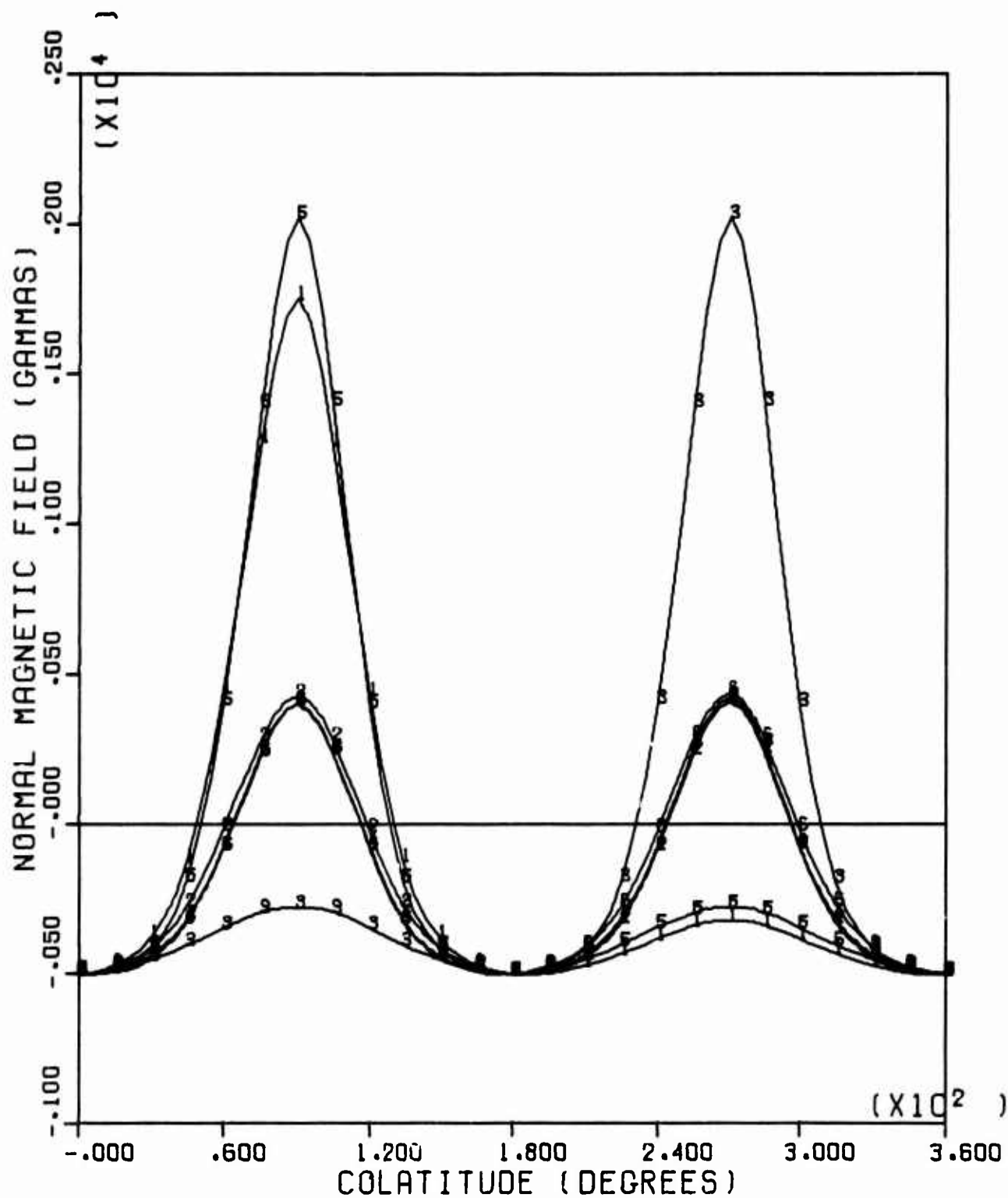


FIG. 10 DATA CURVES FROM SA4024 FOR SAMPLE PROBLEM

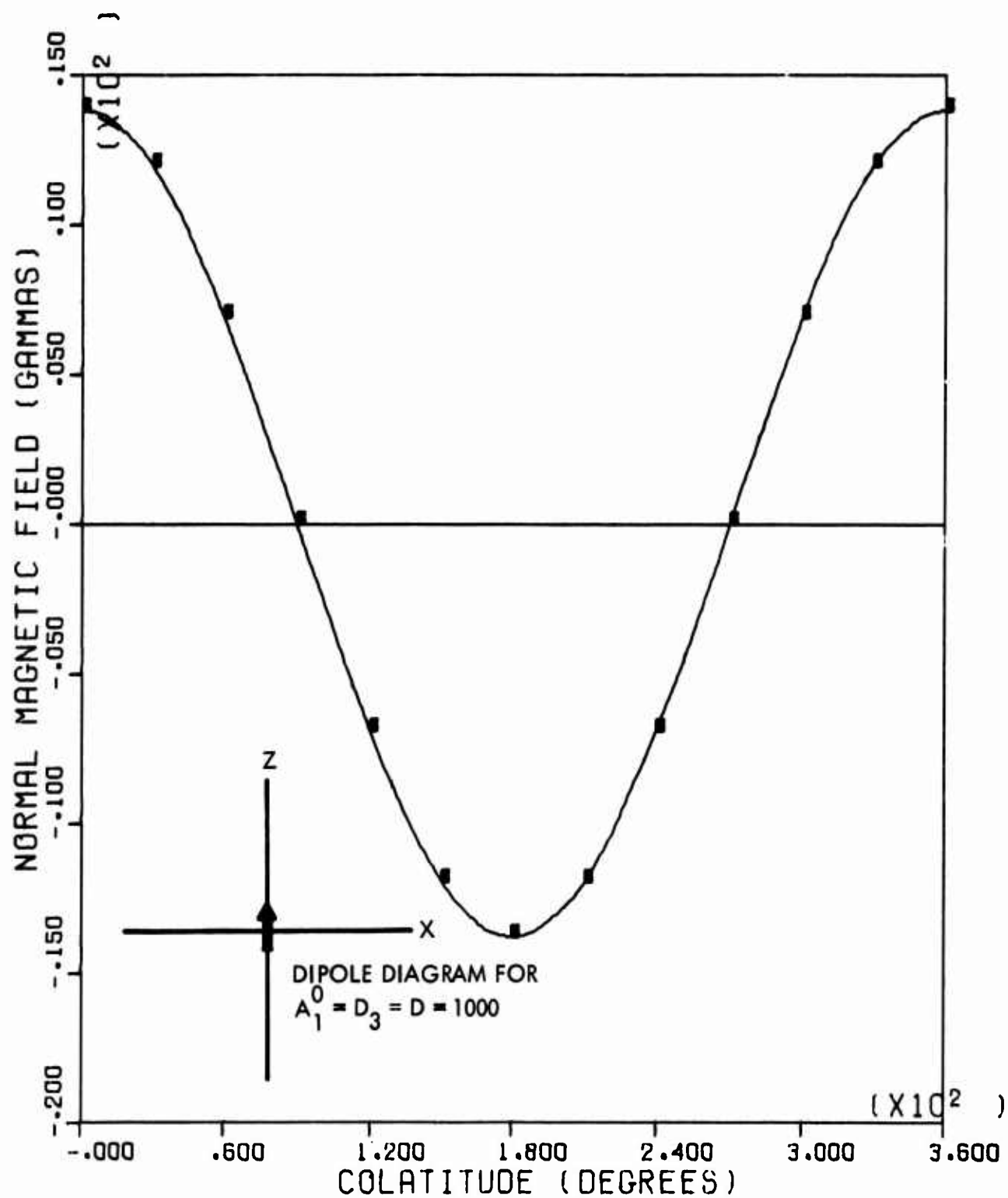


FIG. 11 DIPOLE CURVES FOR $A_1^0 = D_3 = 1000$ GAUSS-CM³

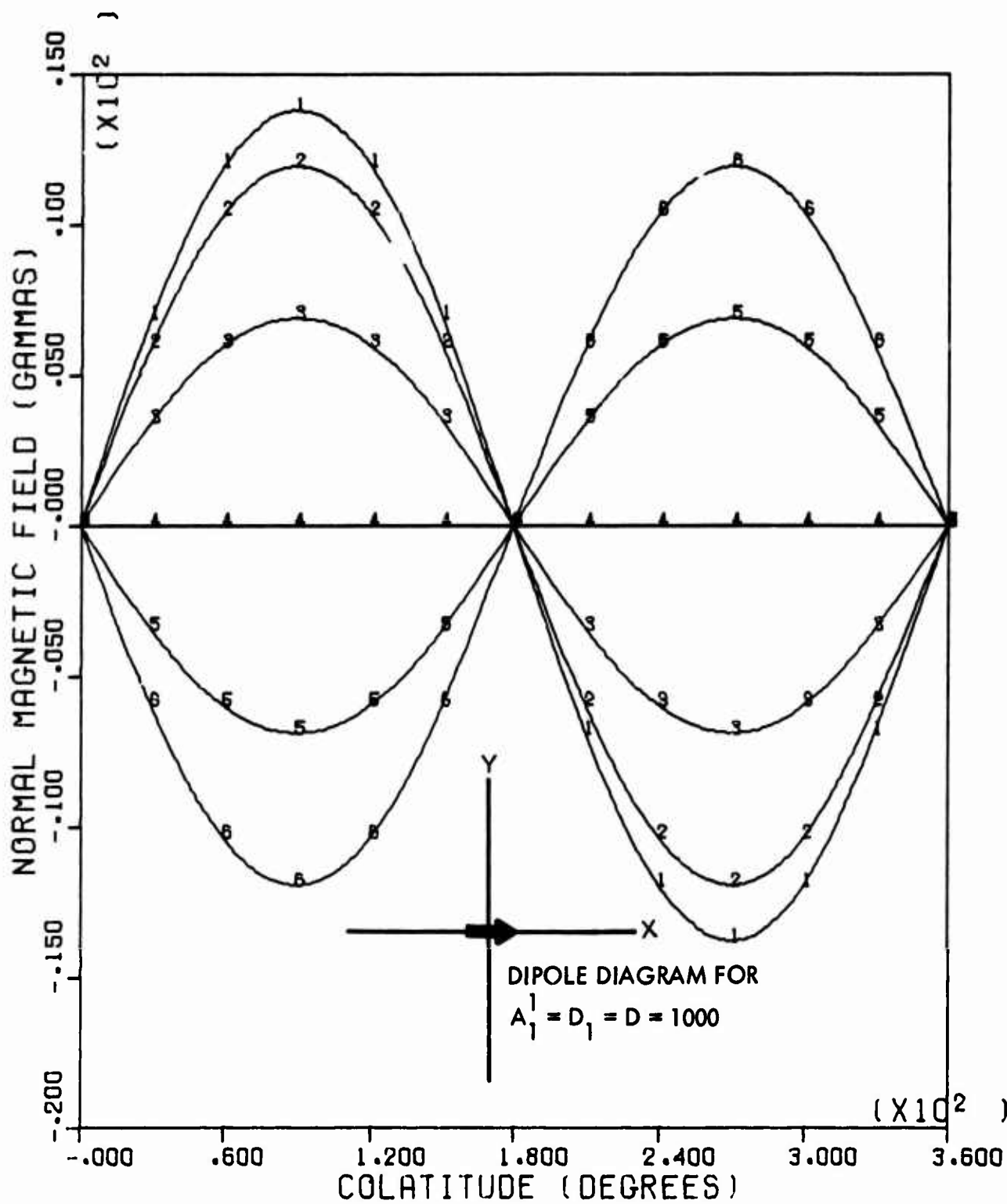


FIG. 12 DIPOLE CURVES FOR $A_1^1 = D_1 = 1000$ GAUSS-CM³

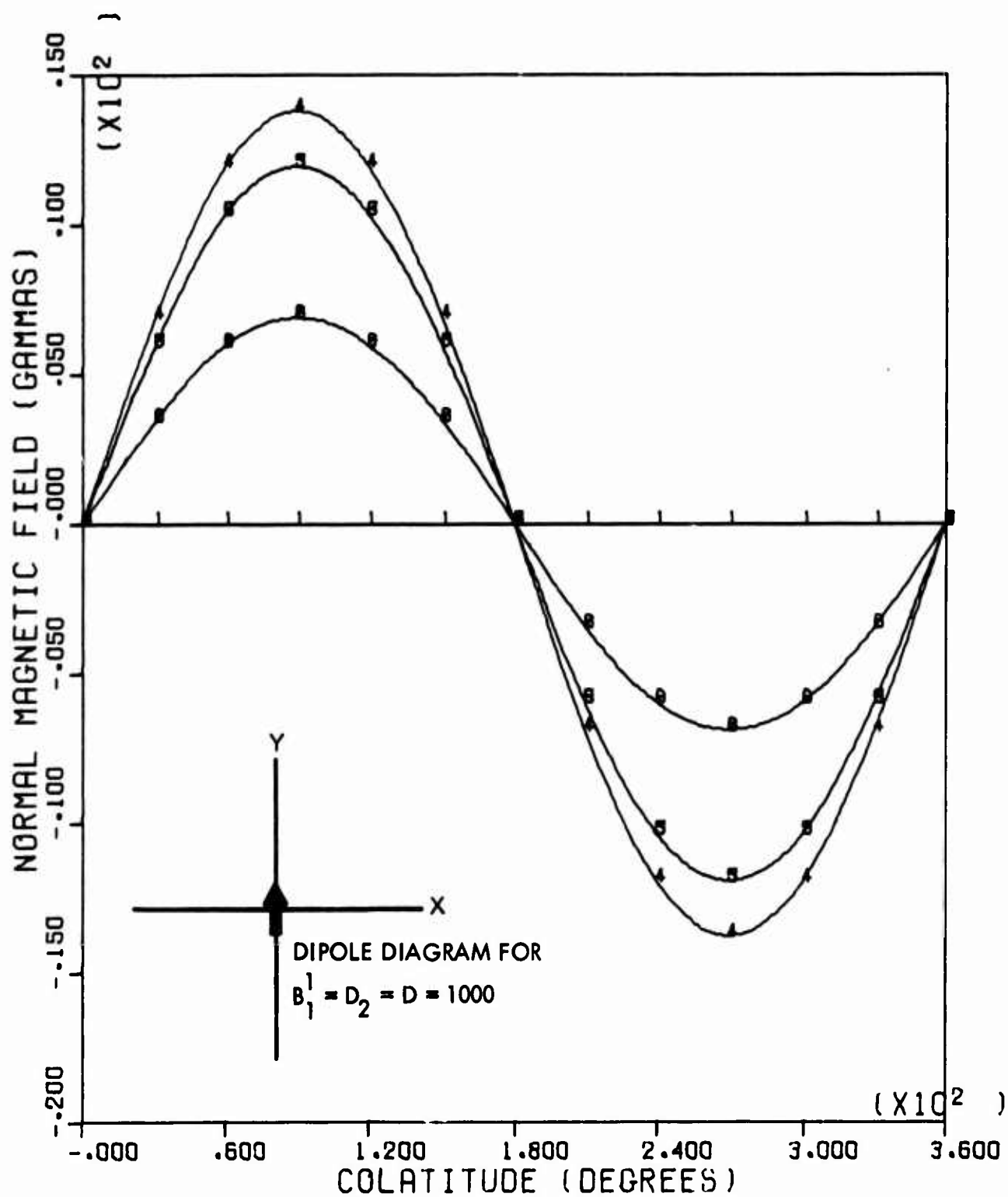


FIG. 13 DIPOLE CURVES FOR $B_1^1 = D_2 = 1000$ GAUSS-CM³

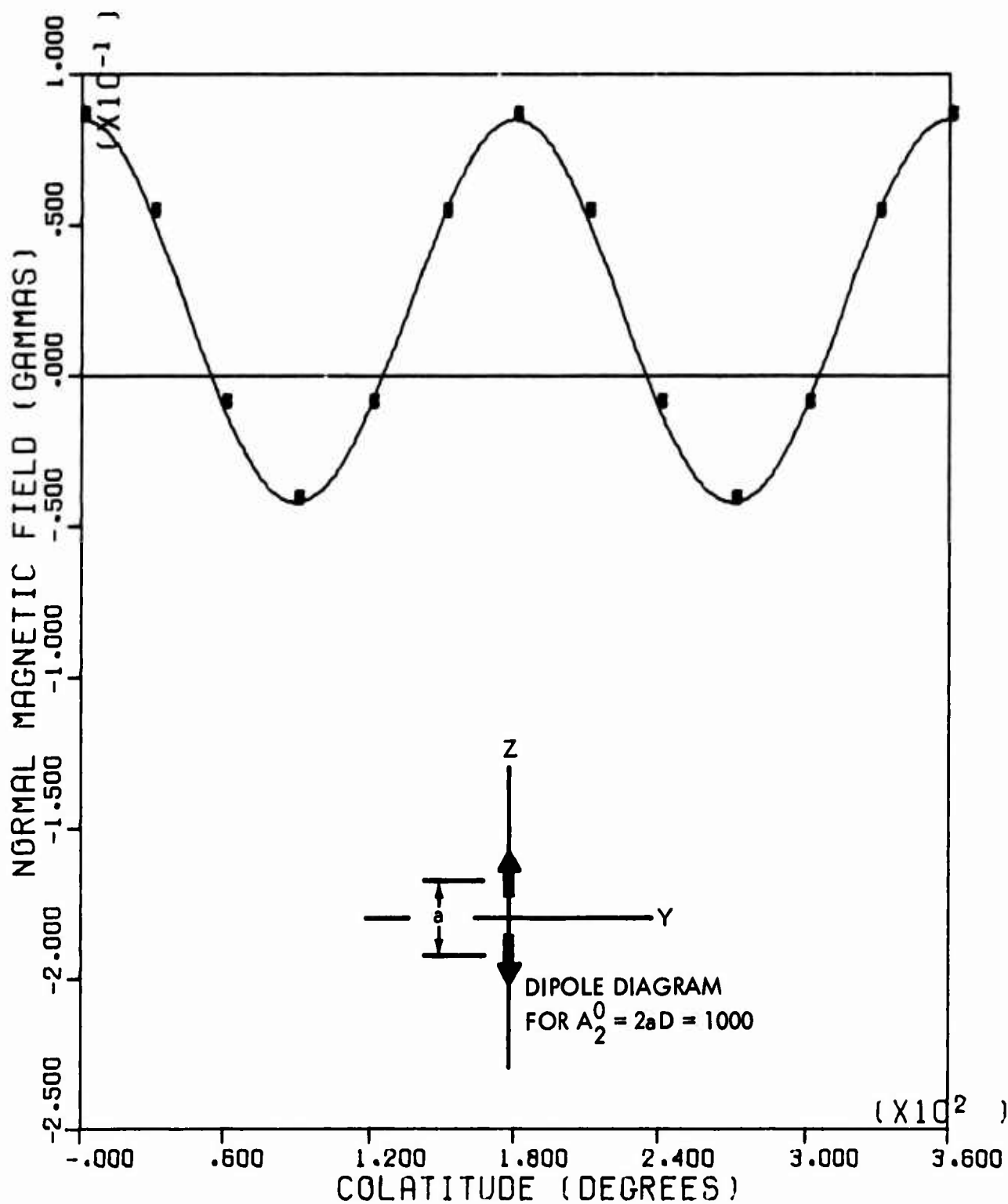


FIG. 14 QUADRUPOLE CURVES FOR $A_2^0 = 1000$ GAUSS-CM⁴

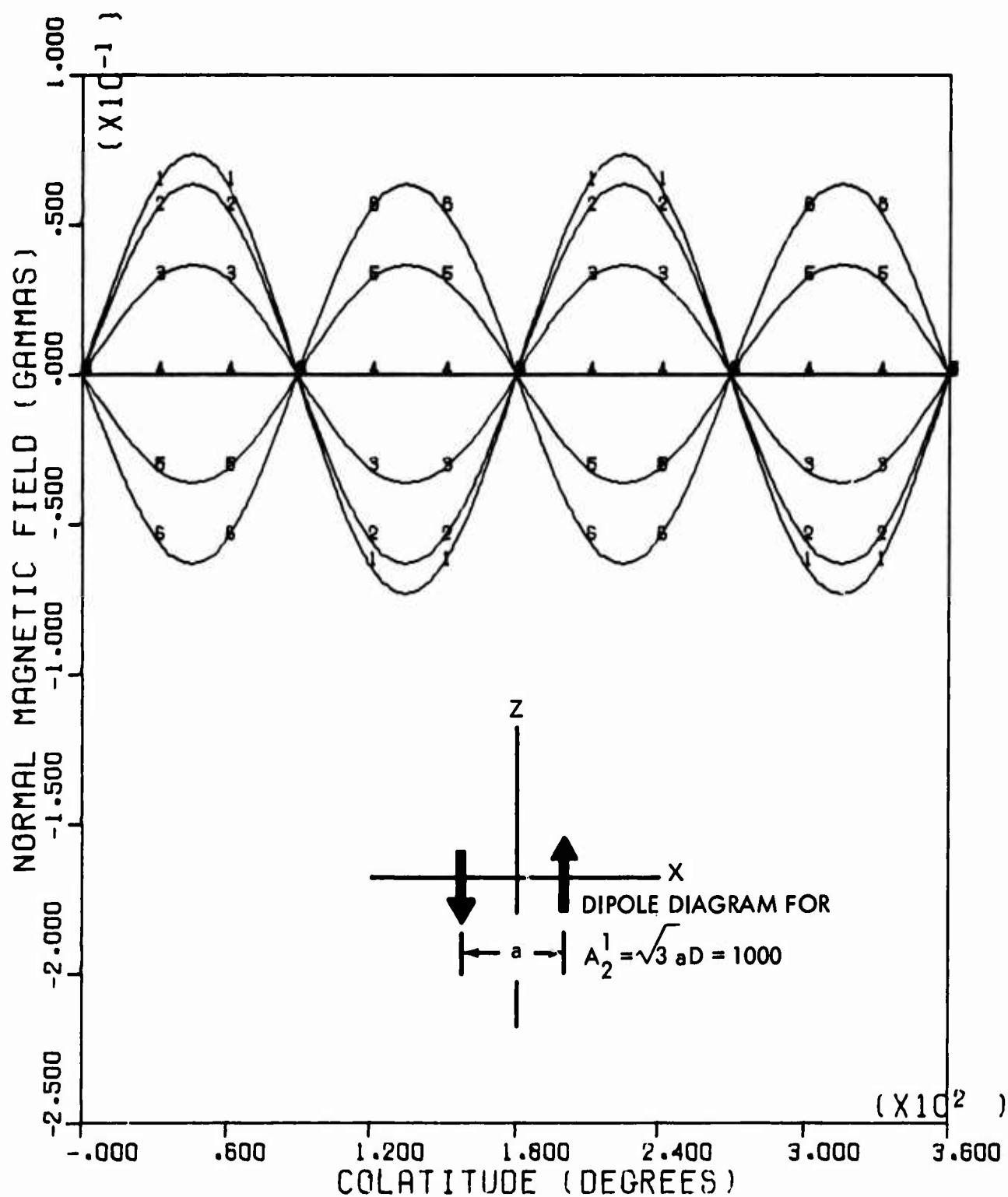


FIG. 15 QUADRUPOLE CURVES FOR $A_2^1 = 1000 \text{ GAUSS-CM}^4$

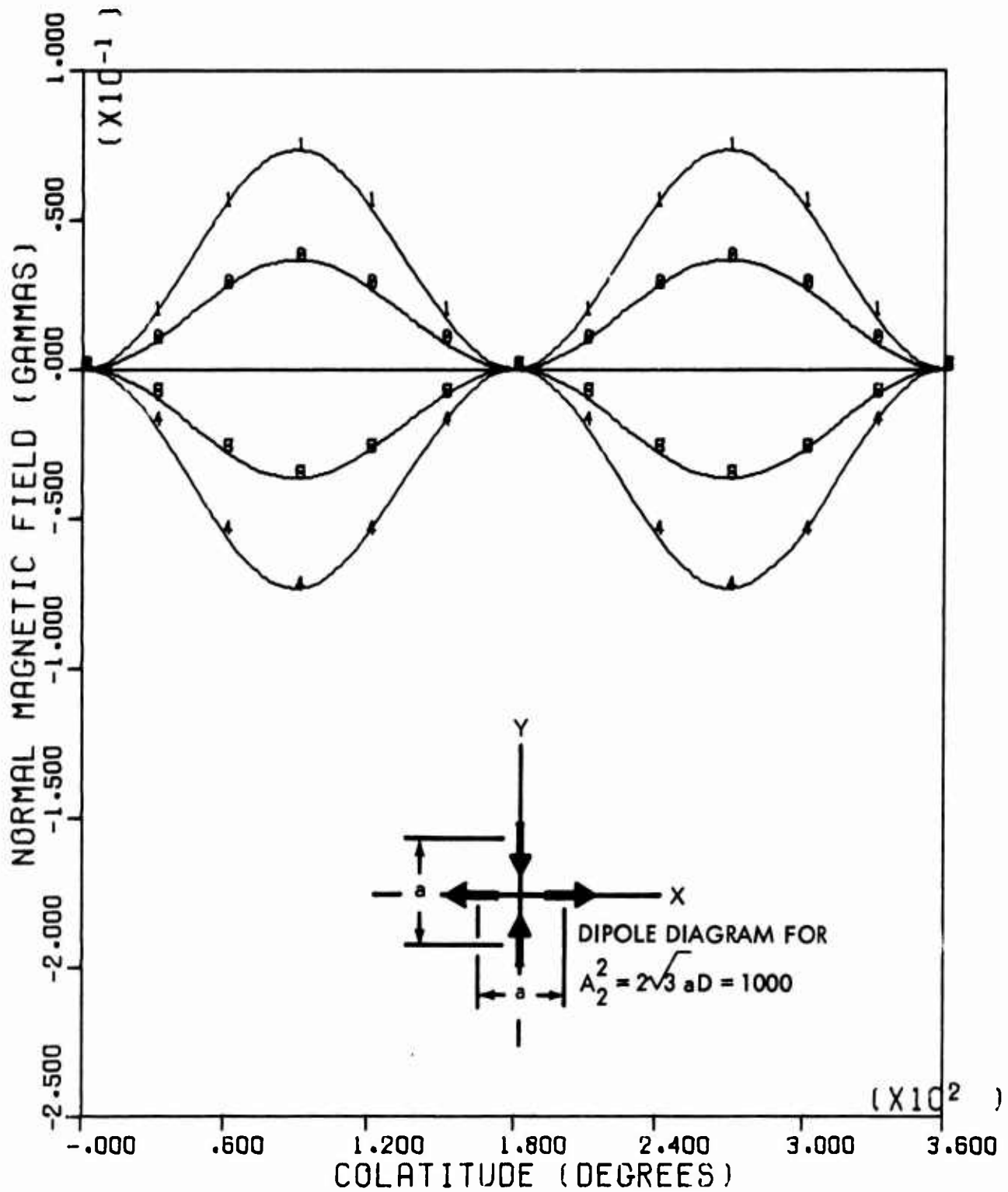


FIG. 16 QUADRUPOLE CURVES FOR $A_2^2 = 1000$ GAUSS-CM⁴

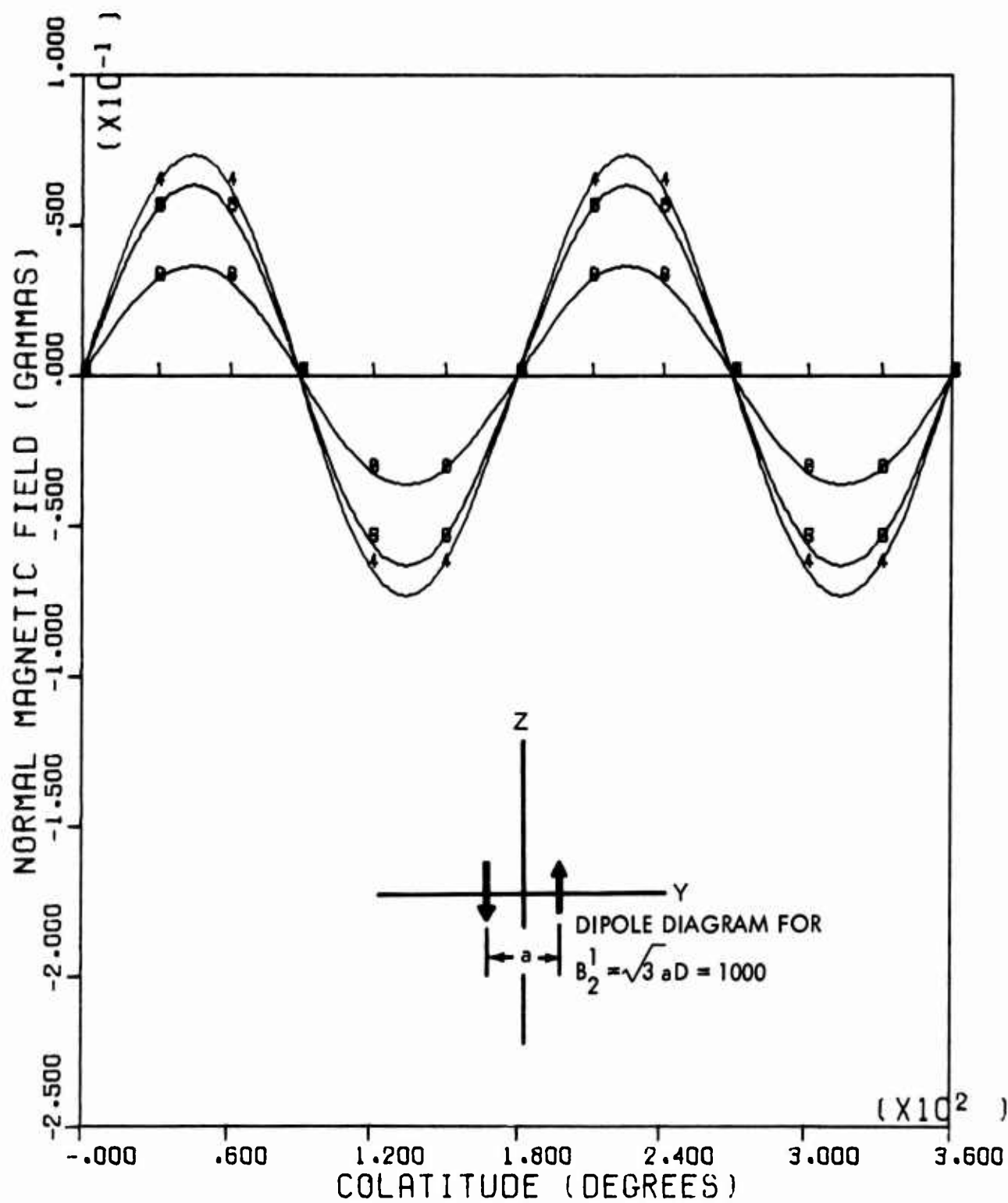


FIG. 17 QUADRUPOLE CURVES FOR $B_2^1 = 1000$ GAUSS-CM⁴

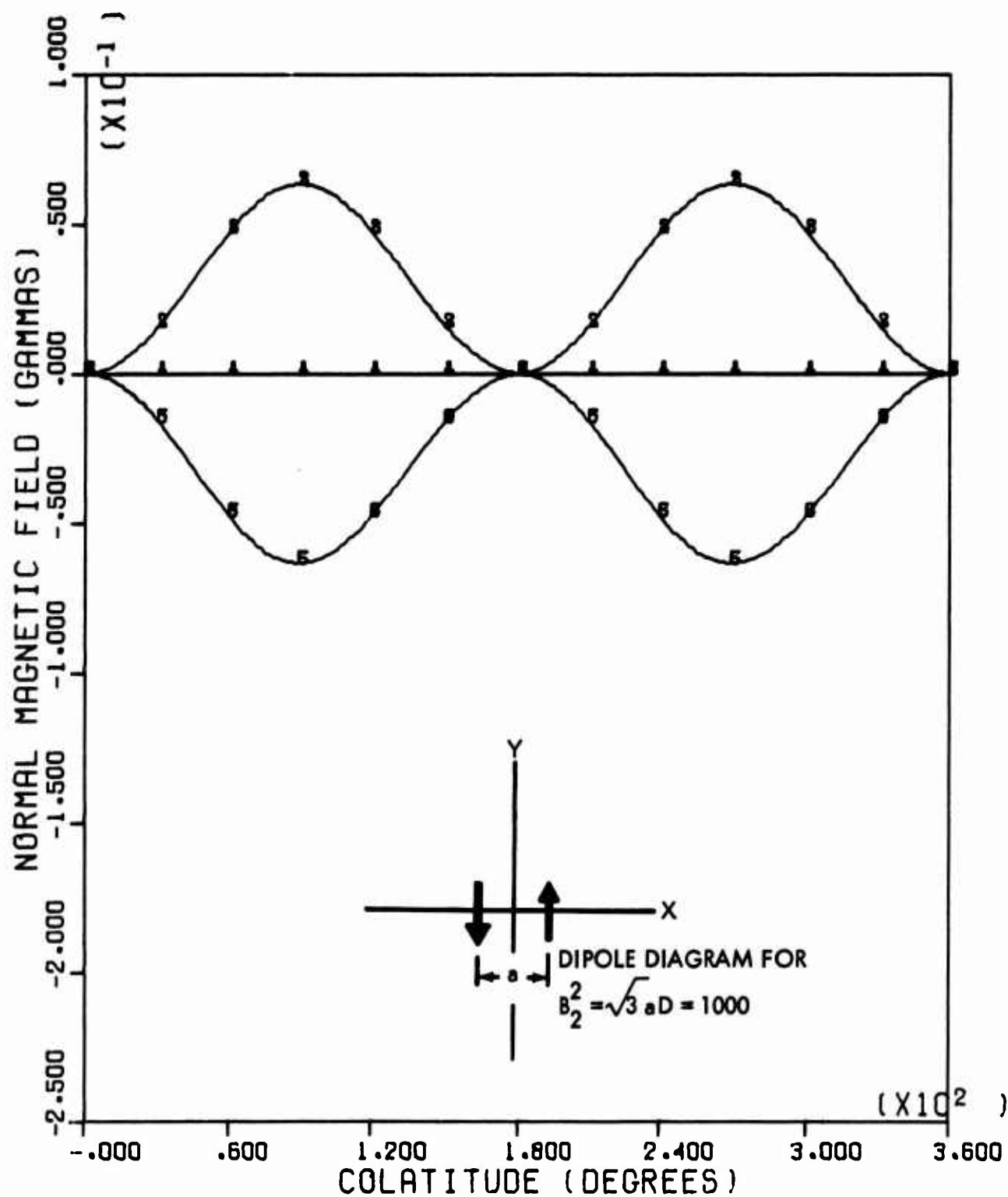


FIG. 18 QUADRUPOLE CURVES FOR $B_2^2 = 1000 \text{ GAUSS-CM}^4$

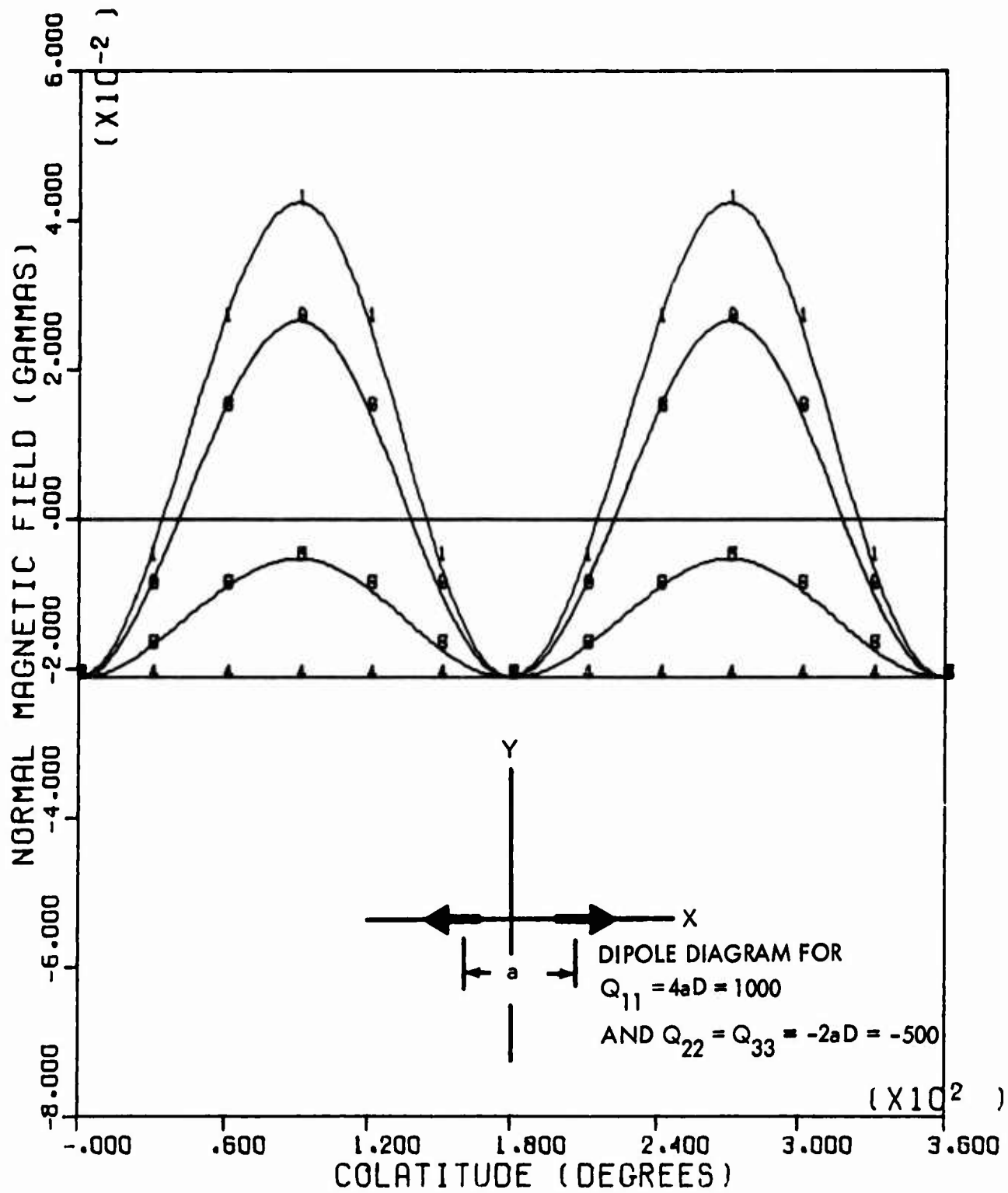


FIG. 19 QUADRUPOLE CURVES FOR $Q_{11} = 1000 \text{ GAUSS-CM}^4$
AND $Q_{22} = Q_{33} = -500 \text{ GAUSS-CM}^4$

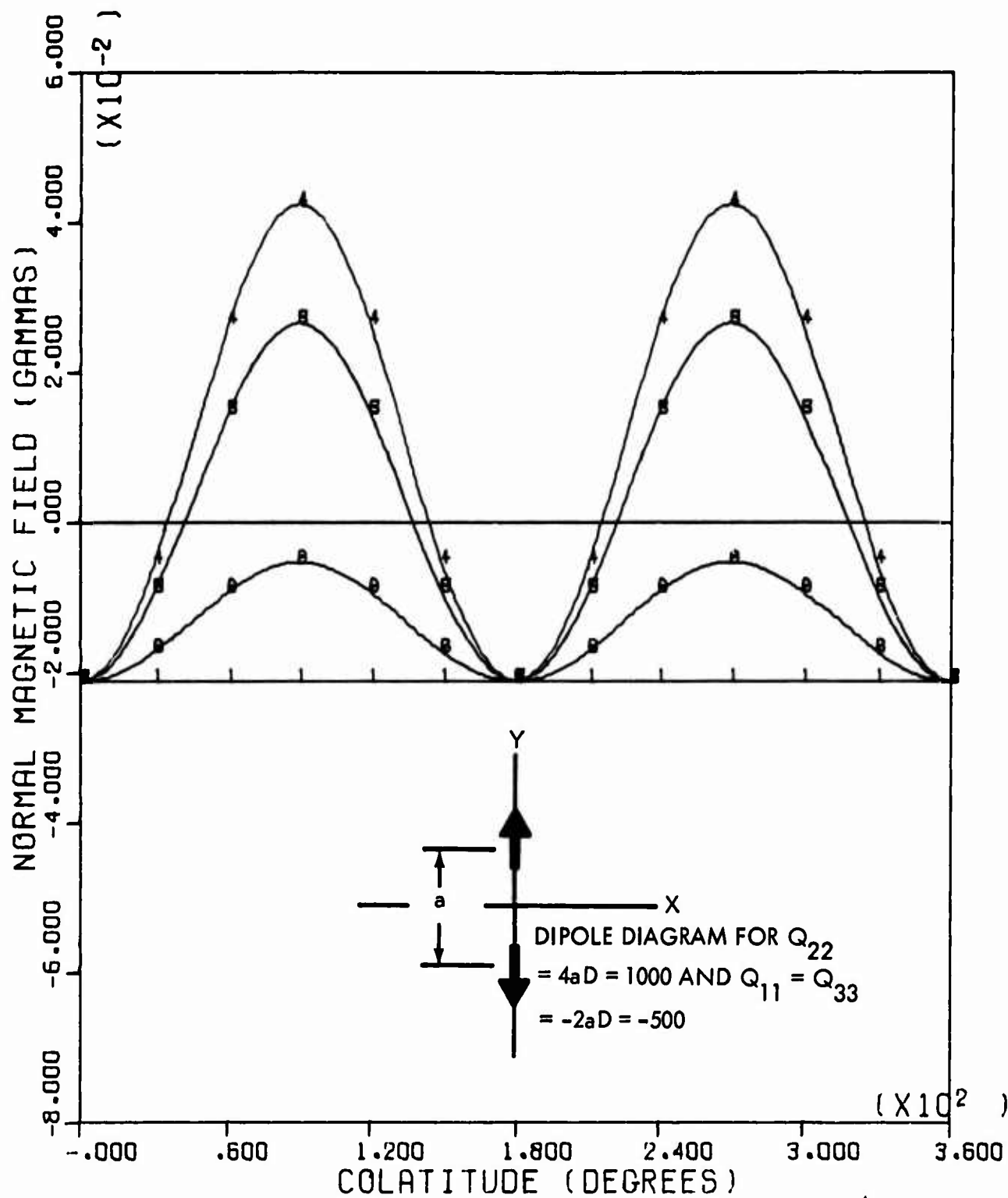


FIG. 20 QUADRUPOLE CURVES FOR $Q_{22} = 1000 \text{ GAUSS-CM}^4$
AND $Q_{11} = Q_{33} = -500 \text{ GAUSS-CM}^4$

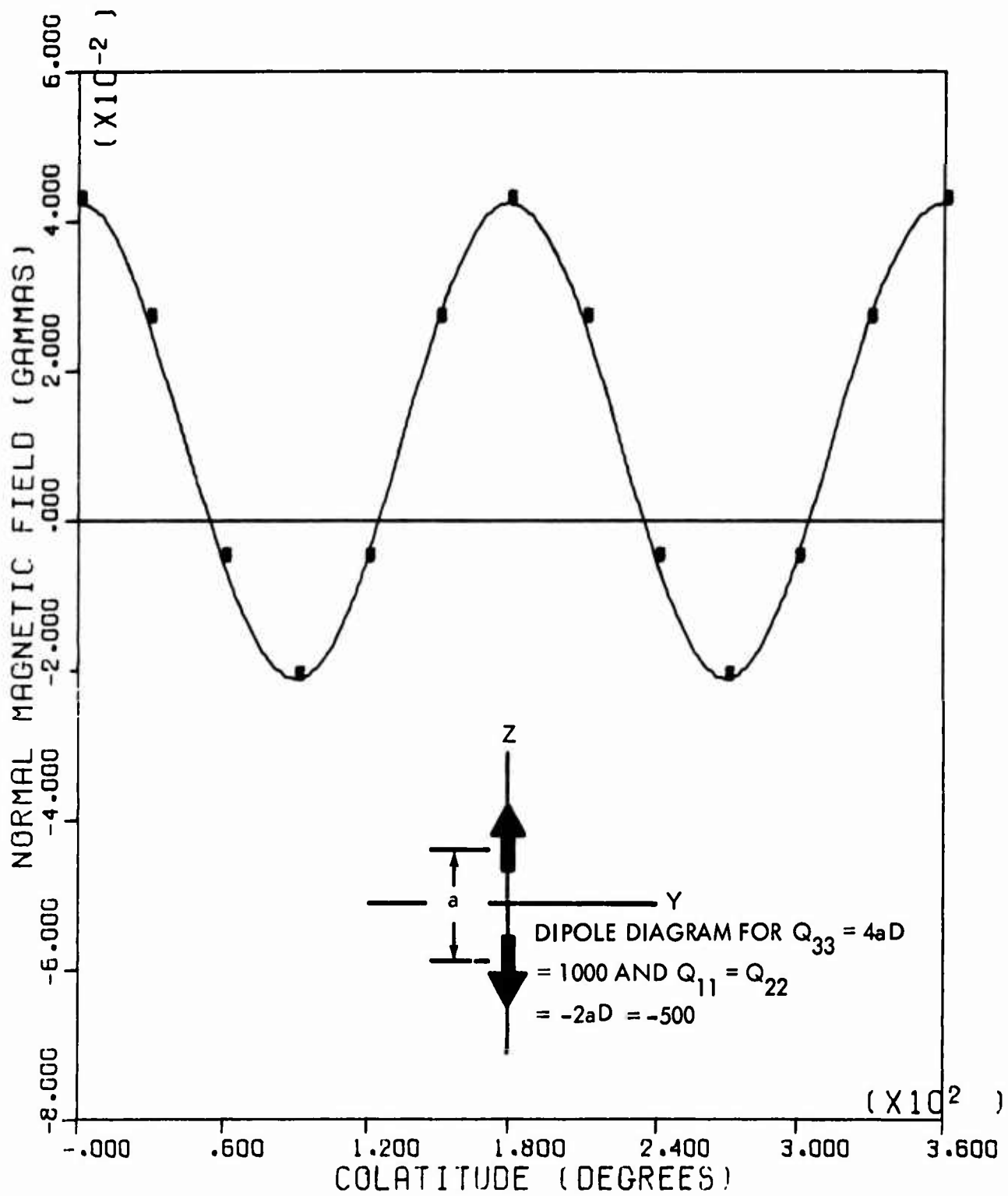


FIG. 21 QUADRUPOLE CURVES FOR $Q_{33} = 1000 \text{ GAUSS-CM}^4$
 AND $Q_{11} = Q_{22} = -500 \text{ GAUSS-CM}^4$

NOLTR 73-191

.500 2.000
(X10³)



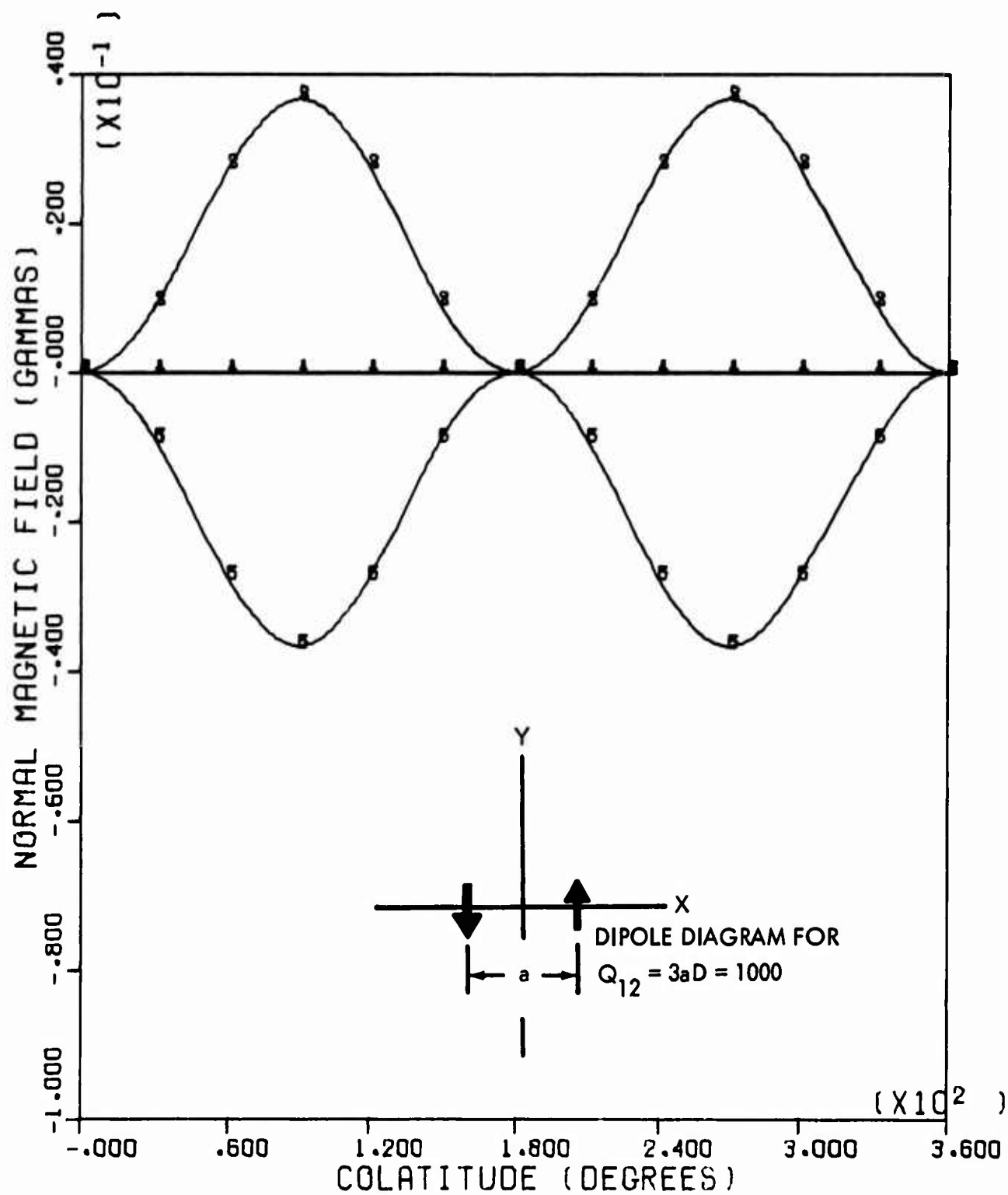
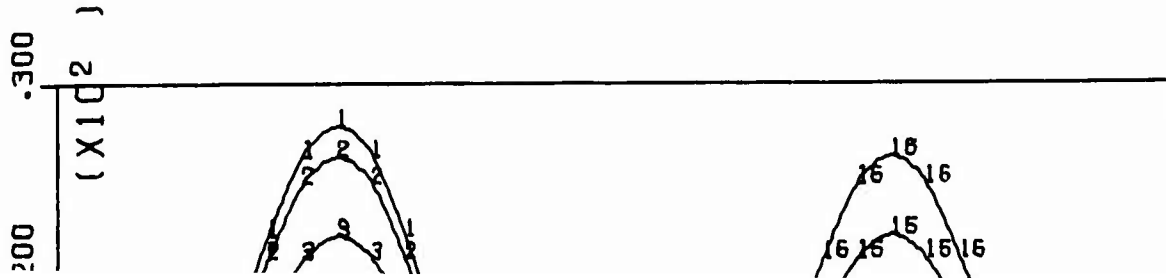


FIG. 22 QUADRUPOLE CURVES FOR $Q_{12} = 1000 \text{ GAUSS-CM}^4$

NOLTR 73-191



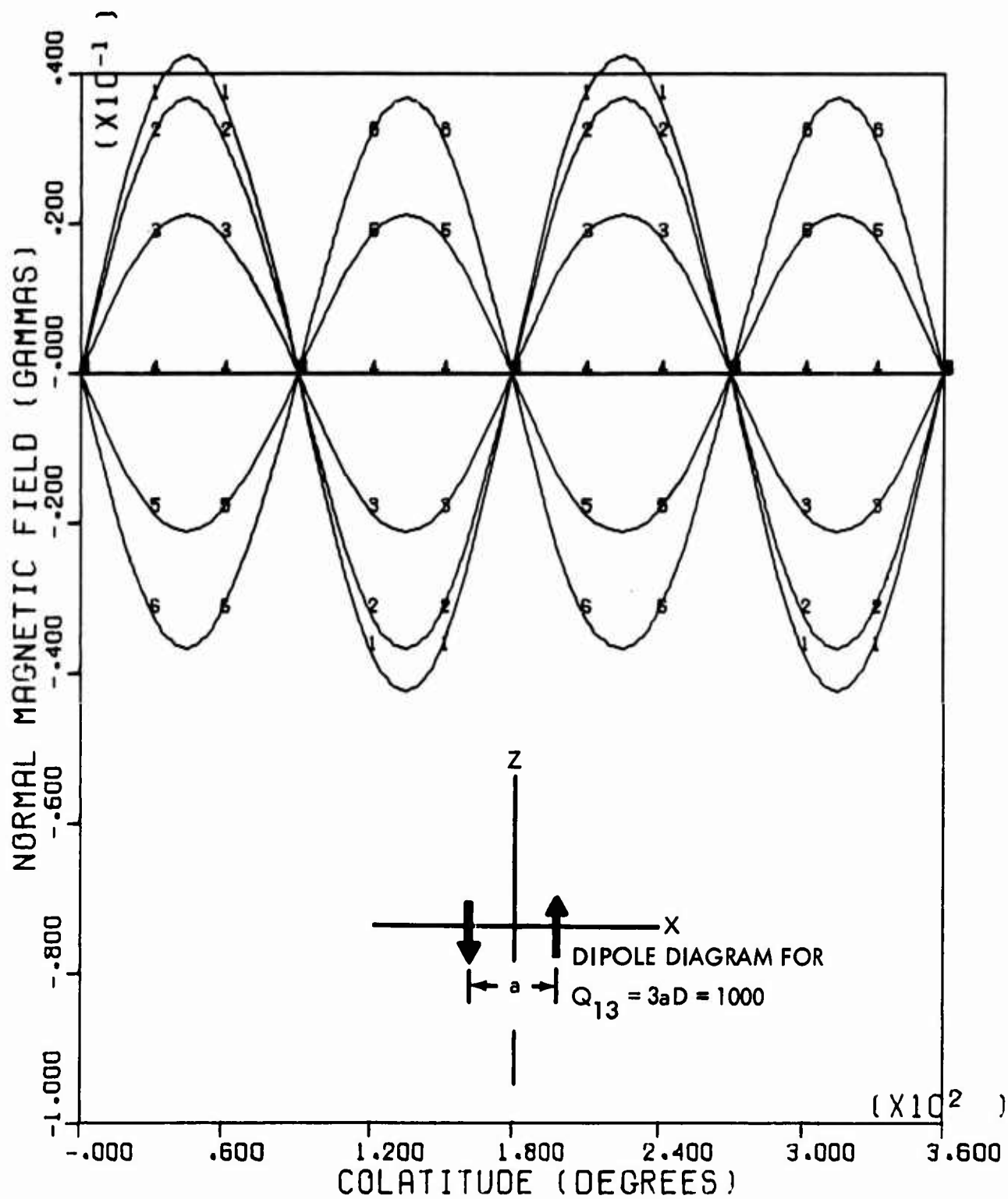
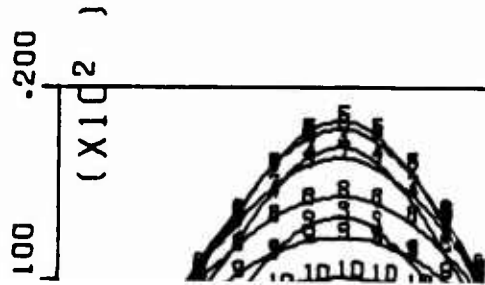


FIG. 23 QUADRUPOLE CURVES FOR $Q_{13} = 1000 \text{ GAUSS-CM}^4$

NOLTR 73-191



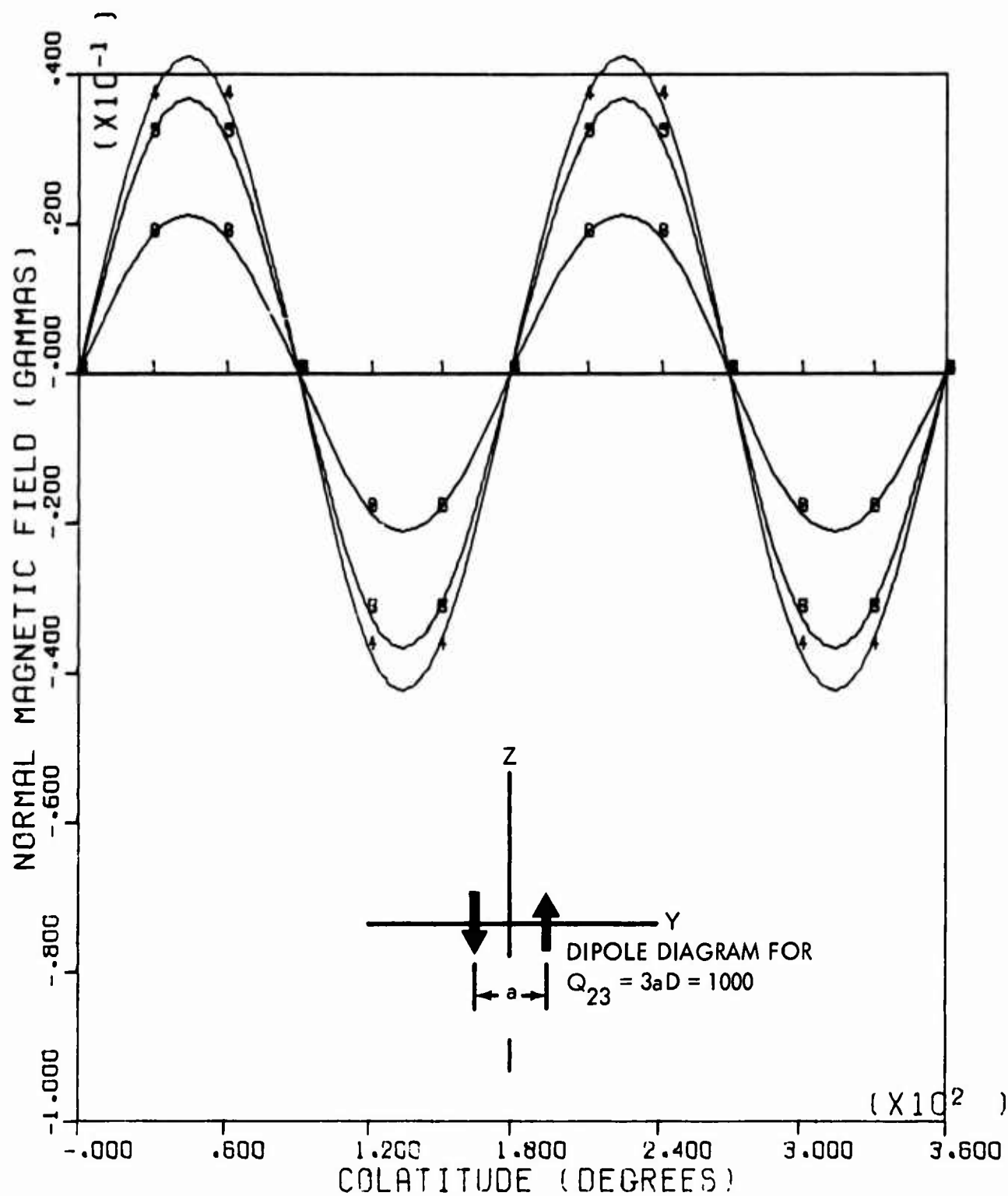


FIG. 24 QUADRUPOLE CURVES FOR $Q_{23} = 1000 \text{ GAUSS-CM}^4$

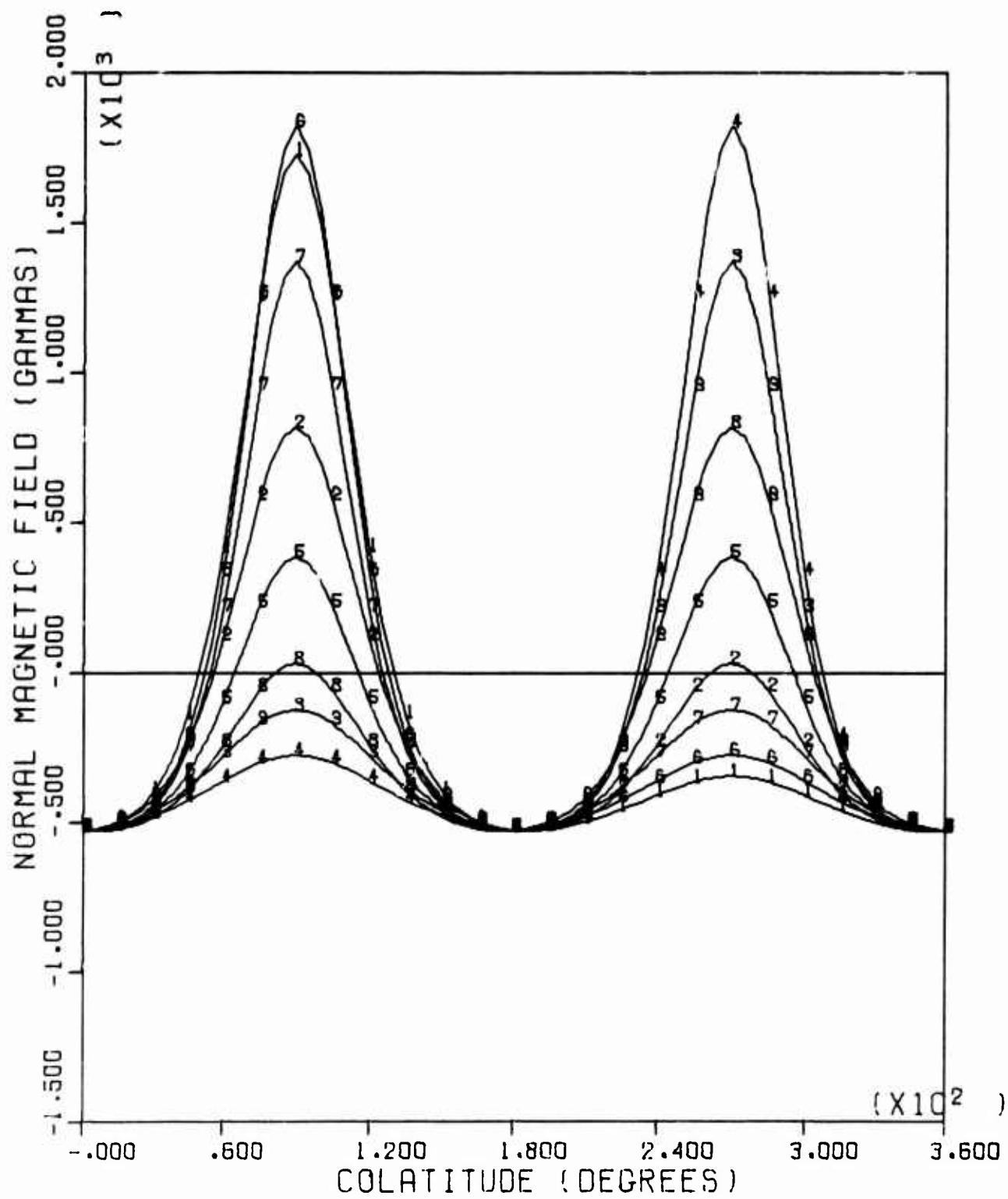


FIG. 25 DATA CURVES FROM SA5024 FOR SAMPLE PROBLEM

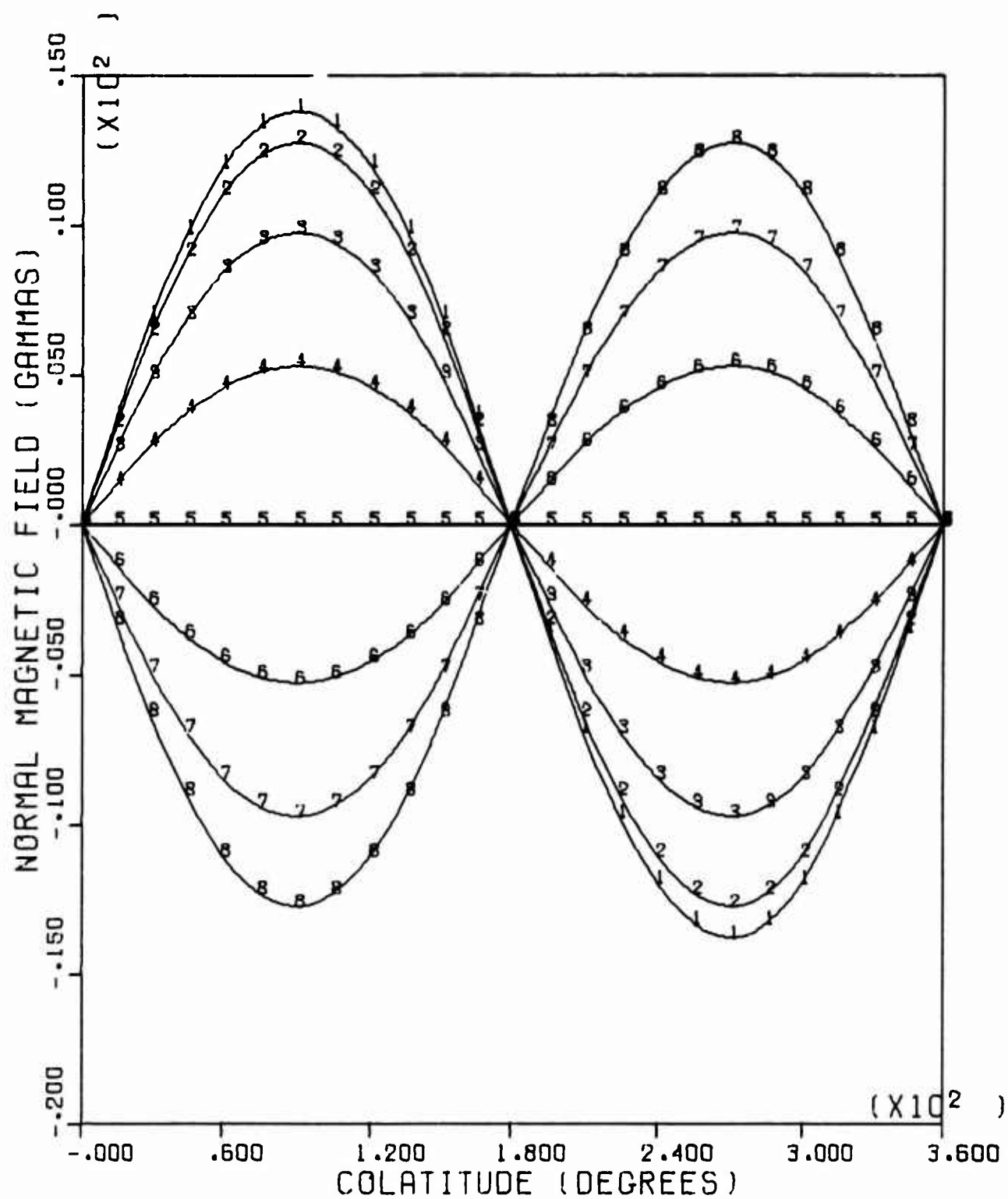


FIG. 26 MULTIPOLE COMPONENT OF DEGREE 1 (DIPOLE COMPONENT)

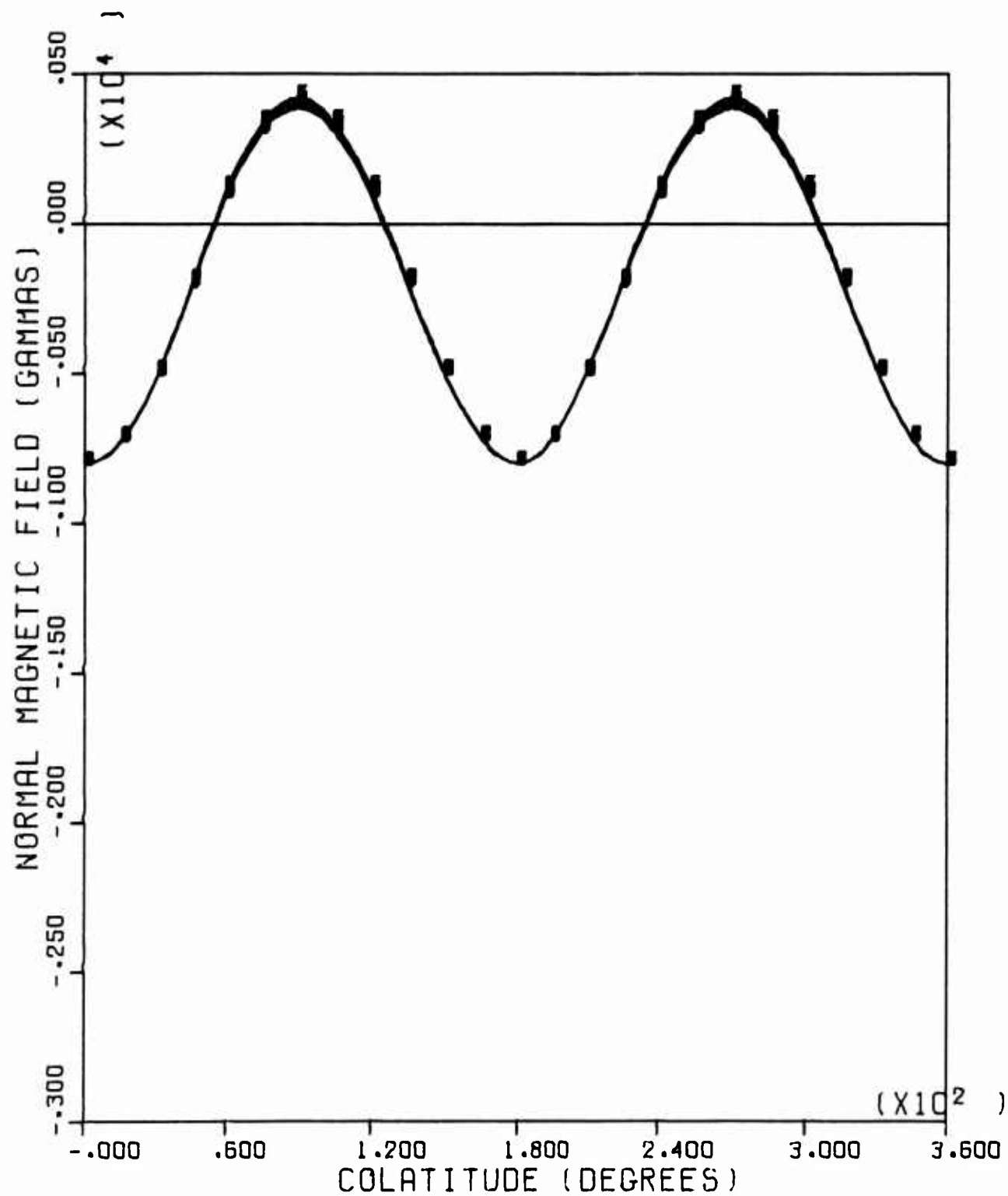


FIG. 27 MULTIPOLE COMPONENT OF DEGREE 2 (QUADRUPOLE COMPONENT)

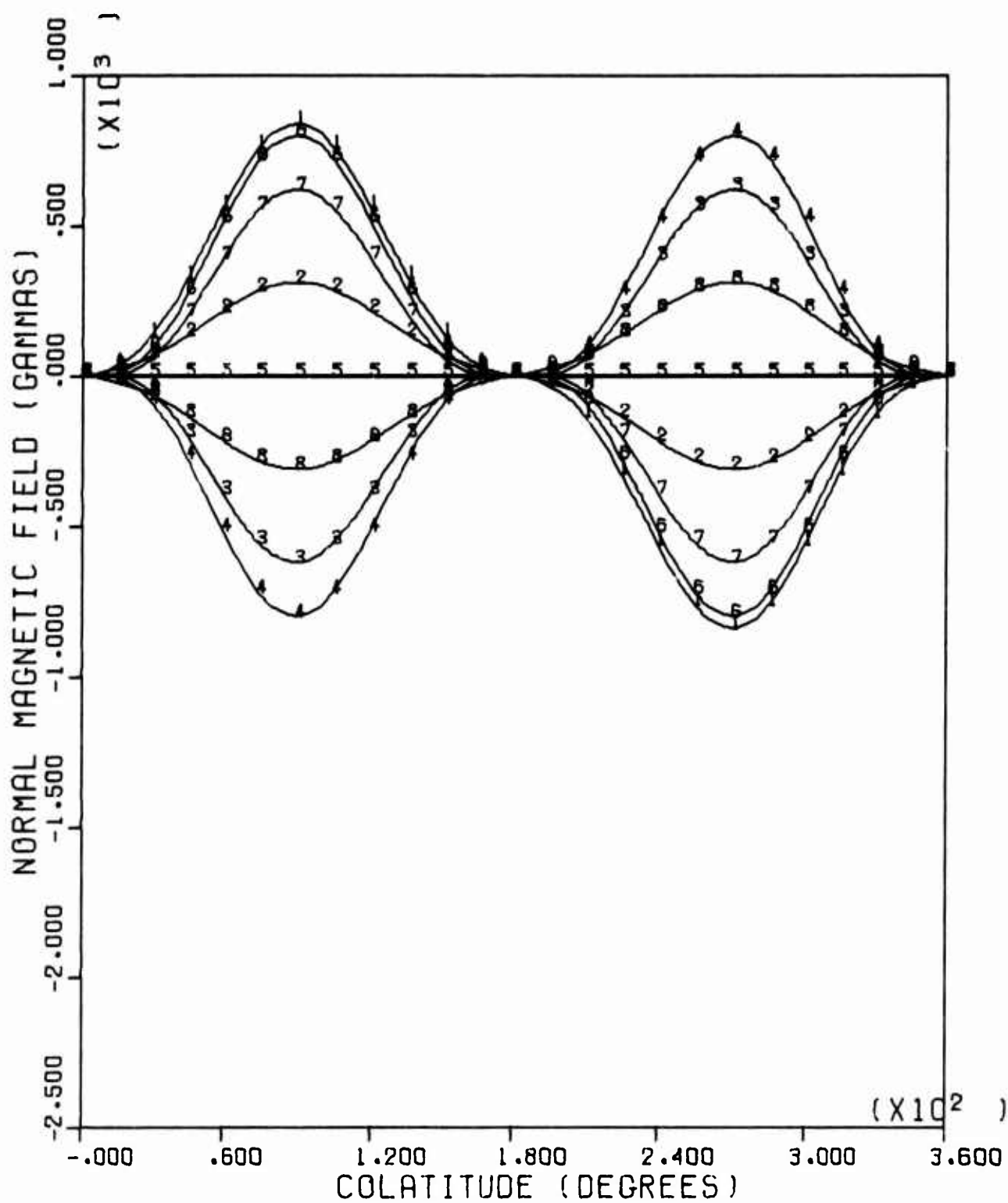


FIG. 28 MULTIPOLE COMPONENT OF DEGREE 3

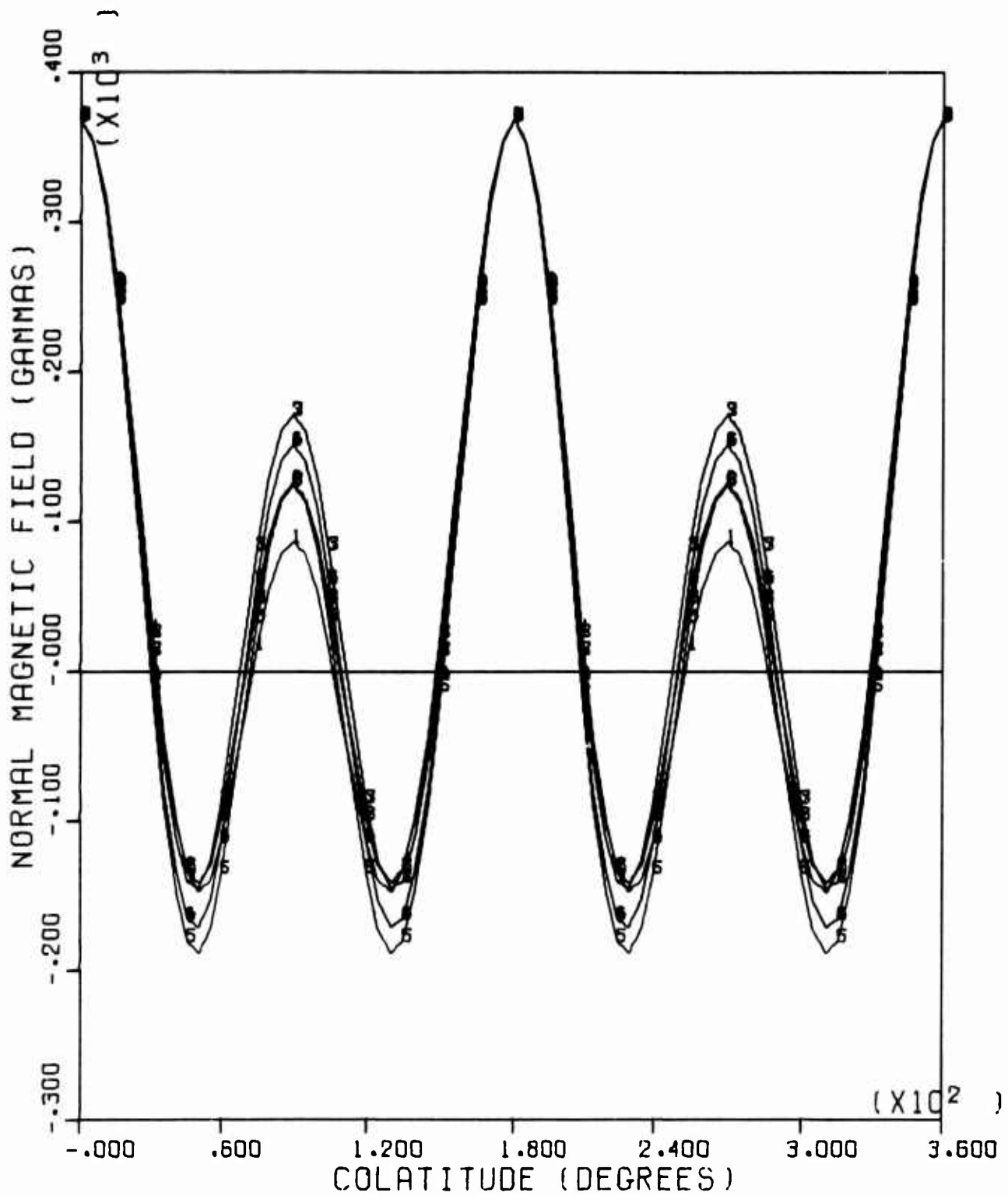


FIG. 29 MULTIPOLE COMPONENT OF DEGREE 4

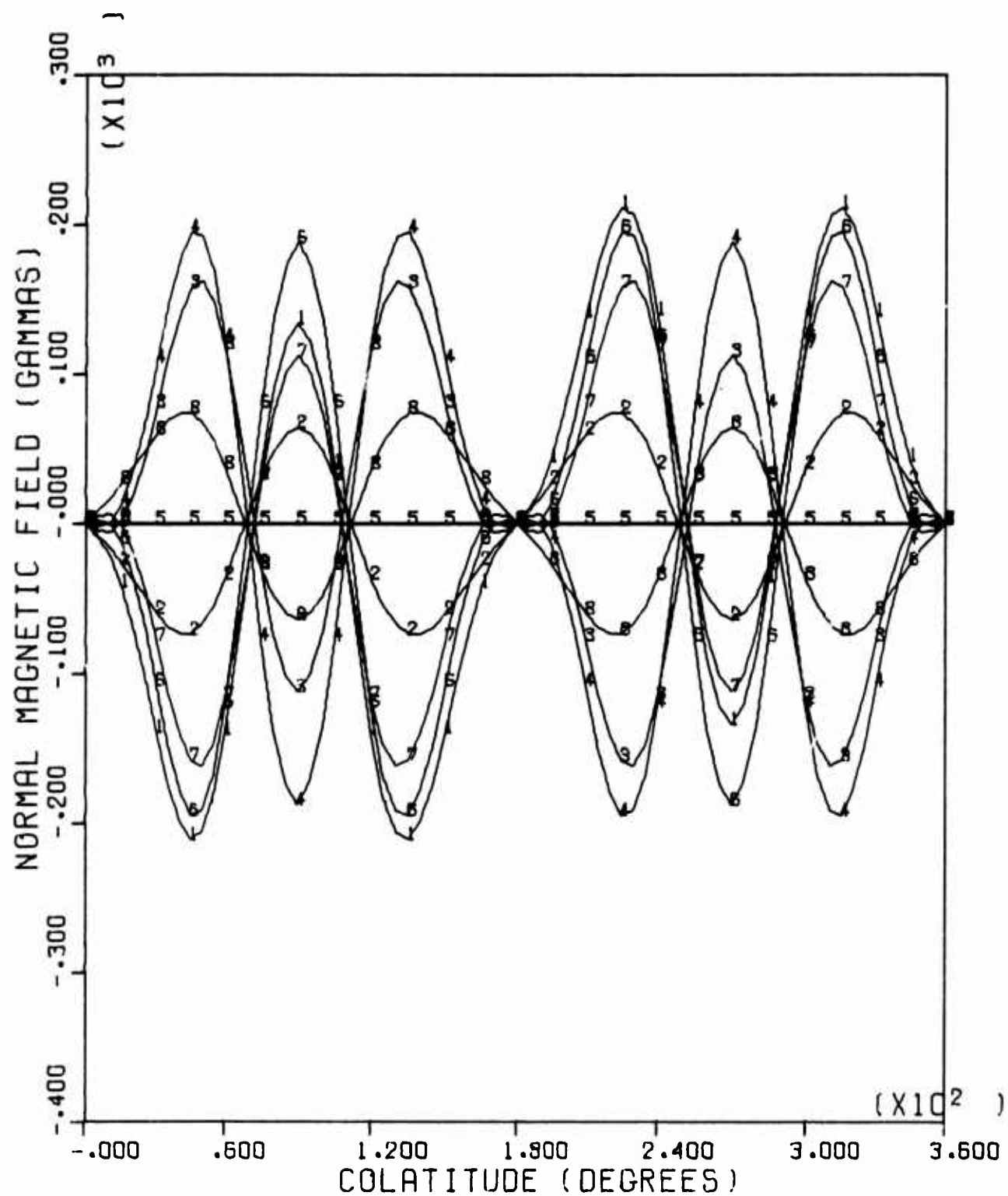


FIG. 30 MULTIPOLE COMPONENT OF DEGREE 5

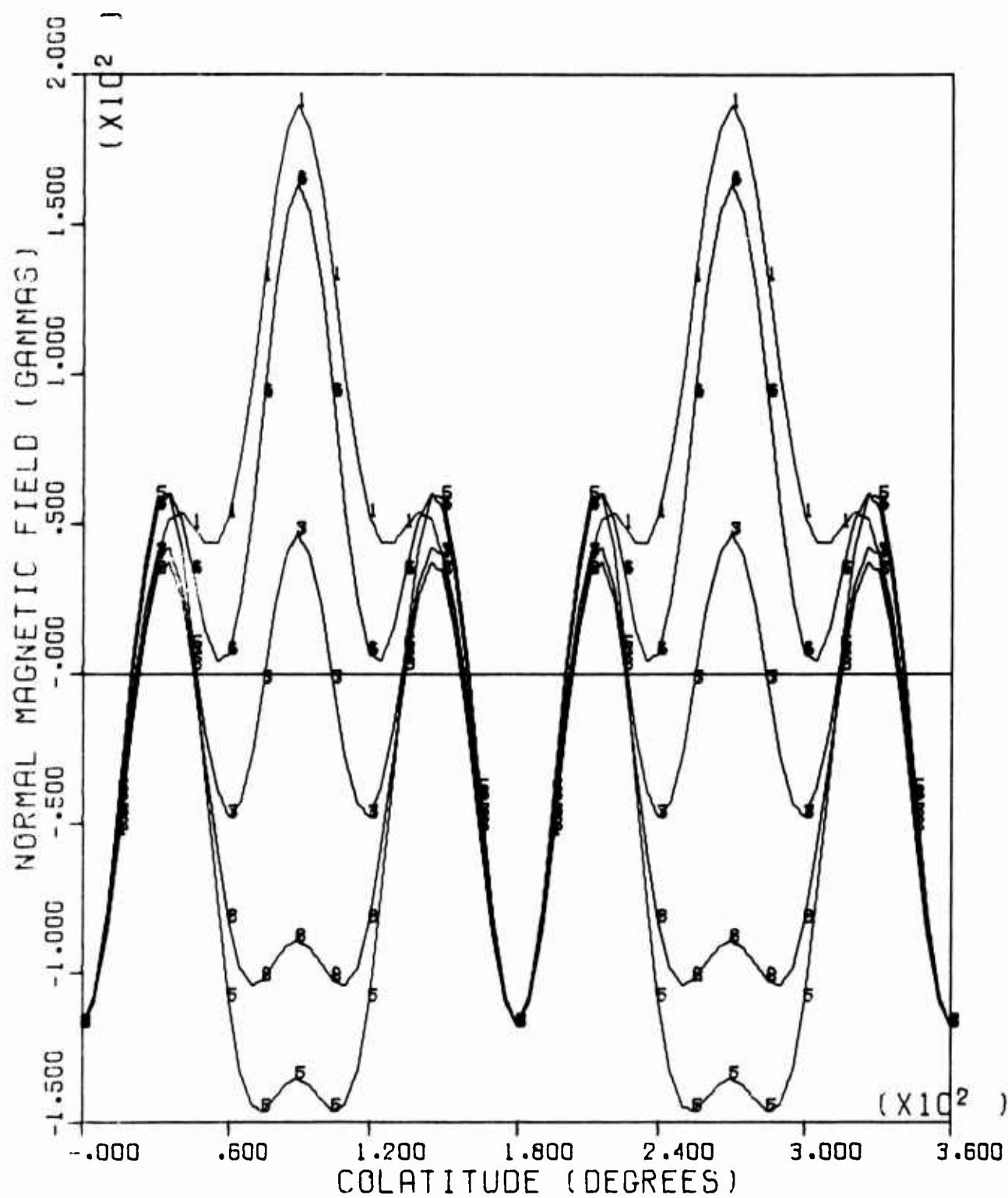


FIG. 31 MULTIPOLE COMPONENT OF DEGREE 6

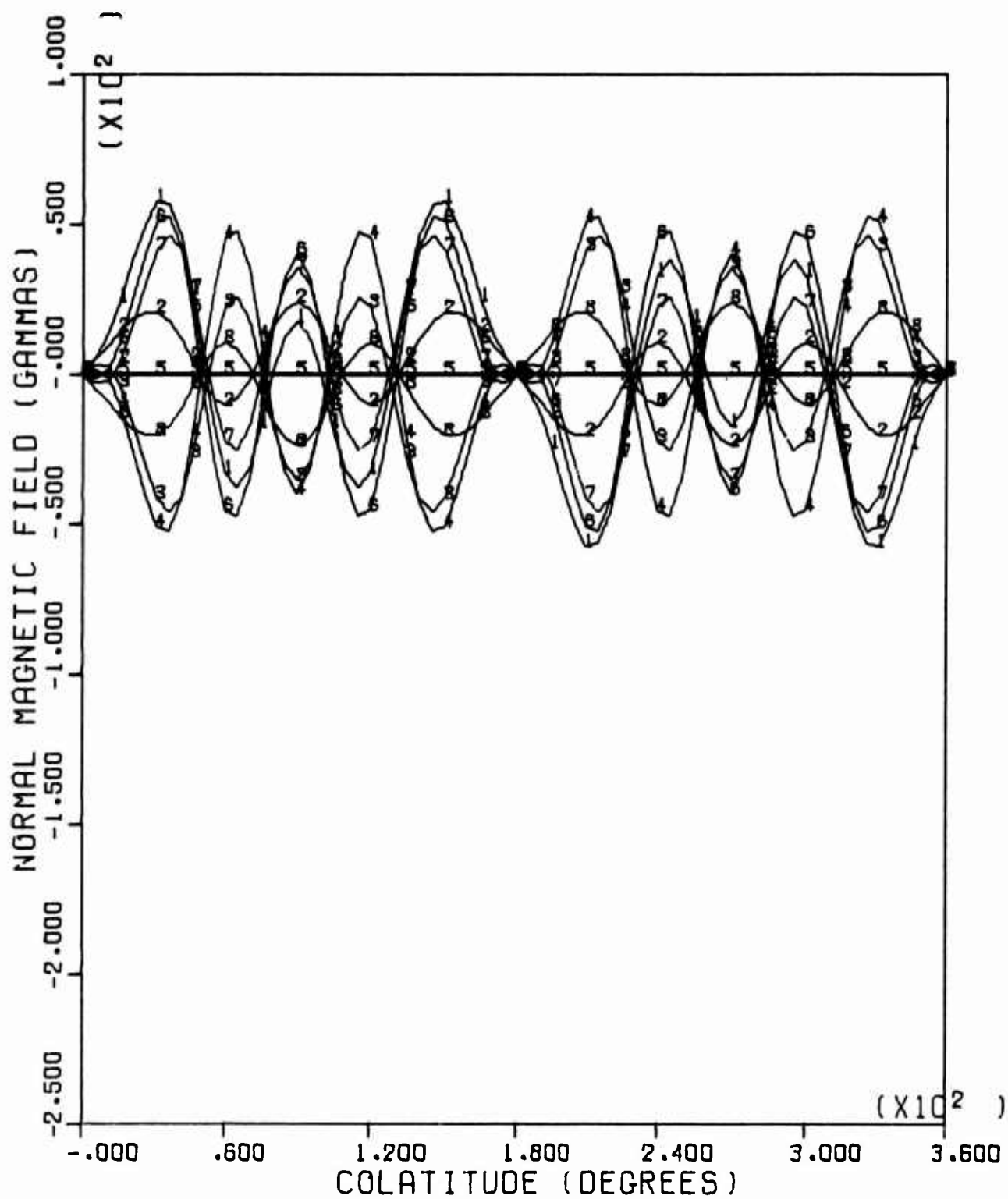


FIG. 32 MULTIPOLE COMPONENT OF DEGREE 7

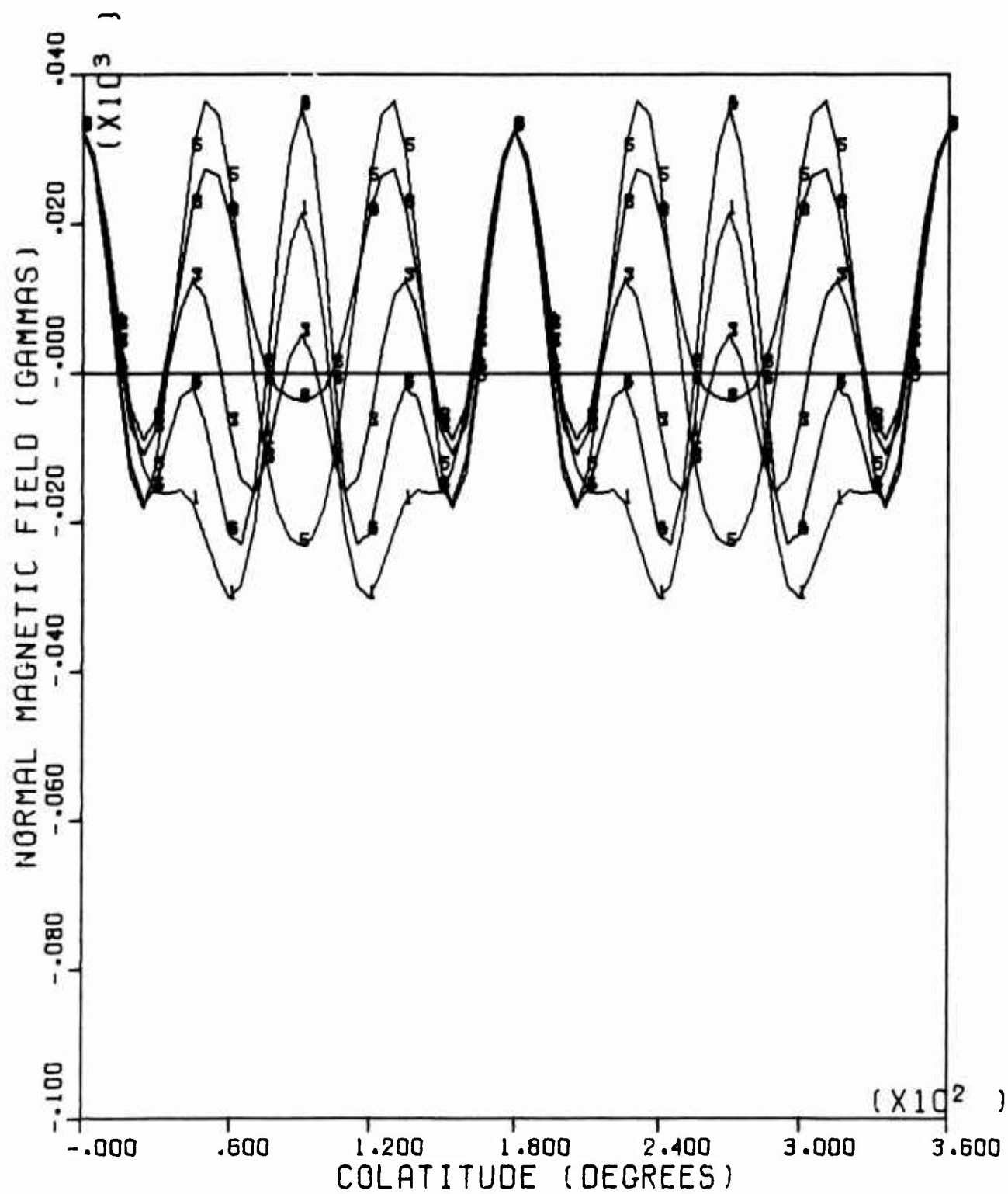


FIG. 33 MULTIPOLE COMPONENT OF DEGREE 8

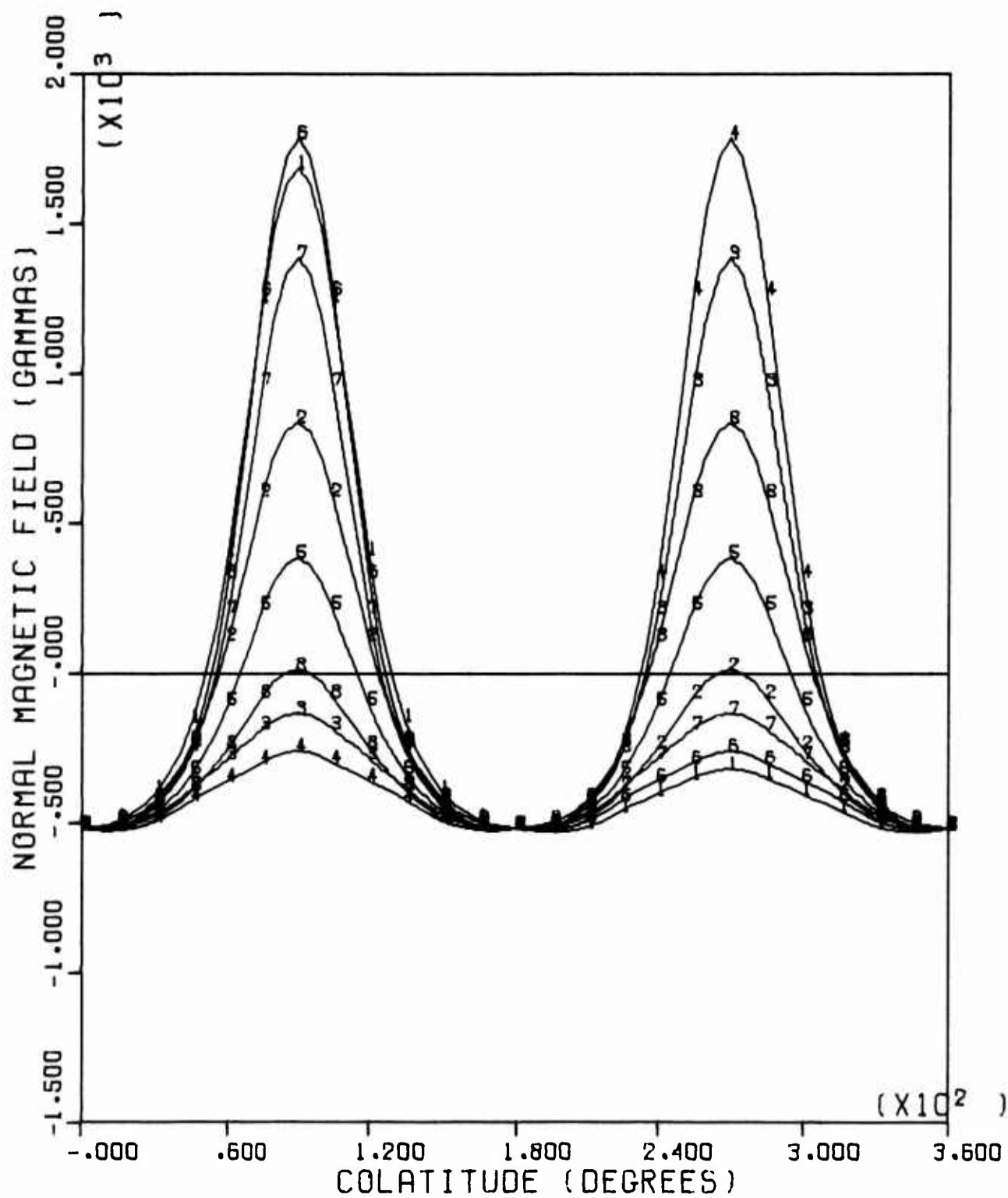


FIG. 34 SUMMATION OF FIRST EIGHT MULTIPOLE COMPONENTS

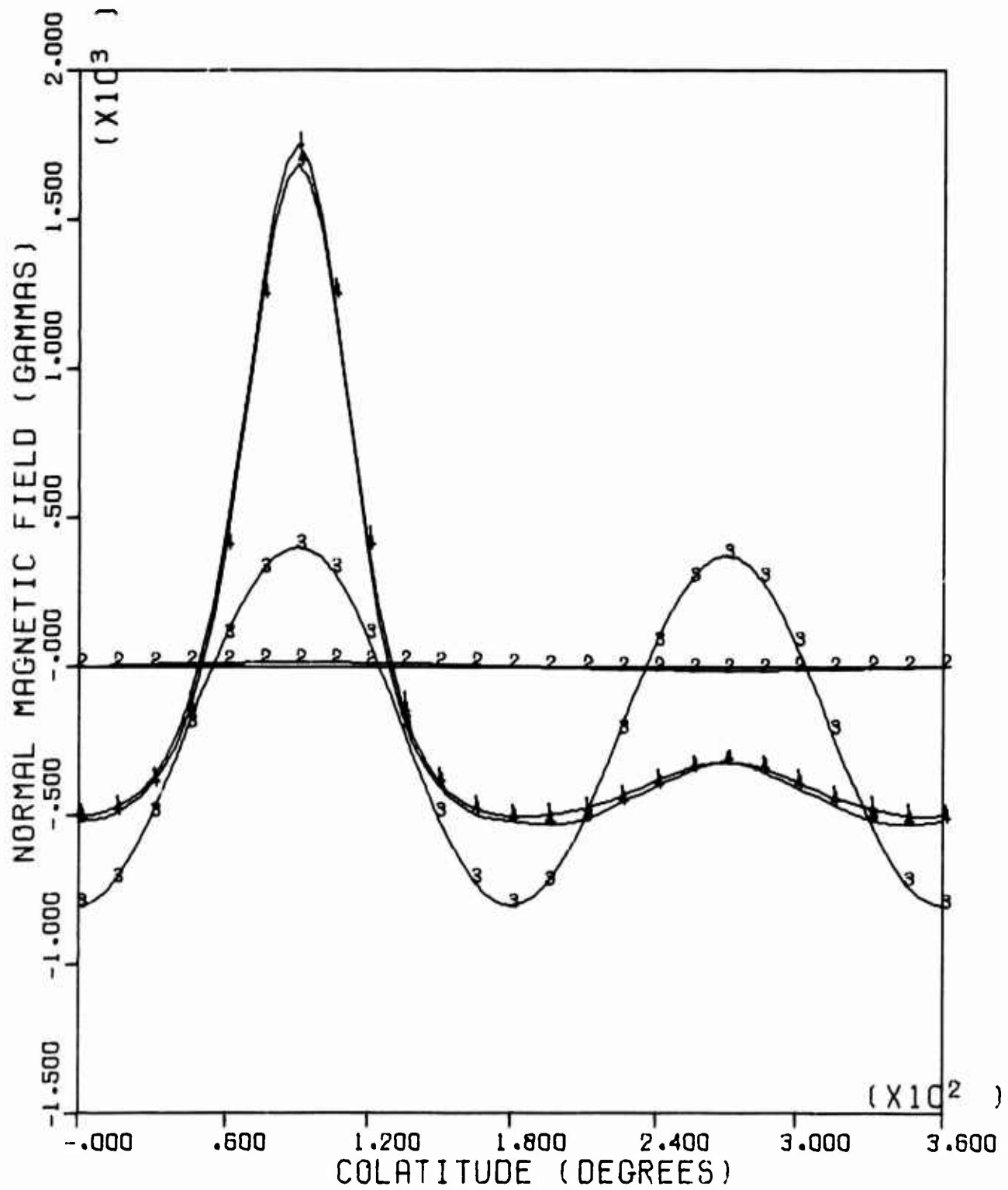


FIG. 35 INDIVIDUAL CURVES DEMONSTRATING APPROXIMATION ACCURACY

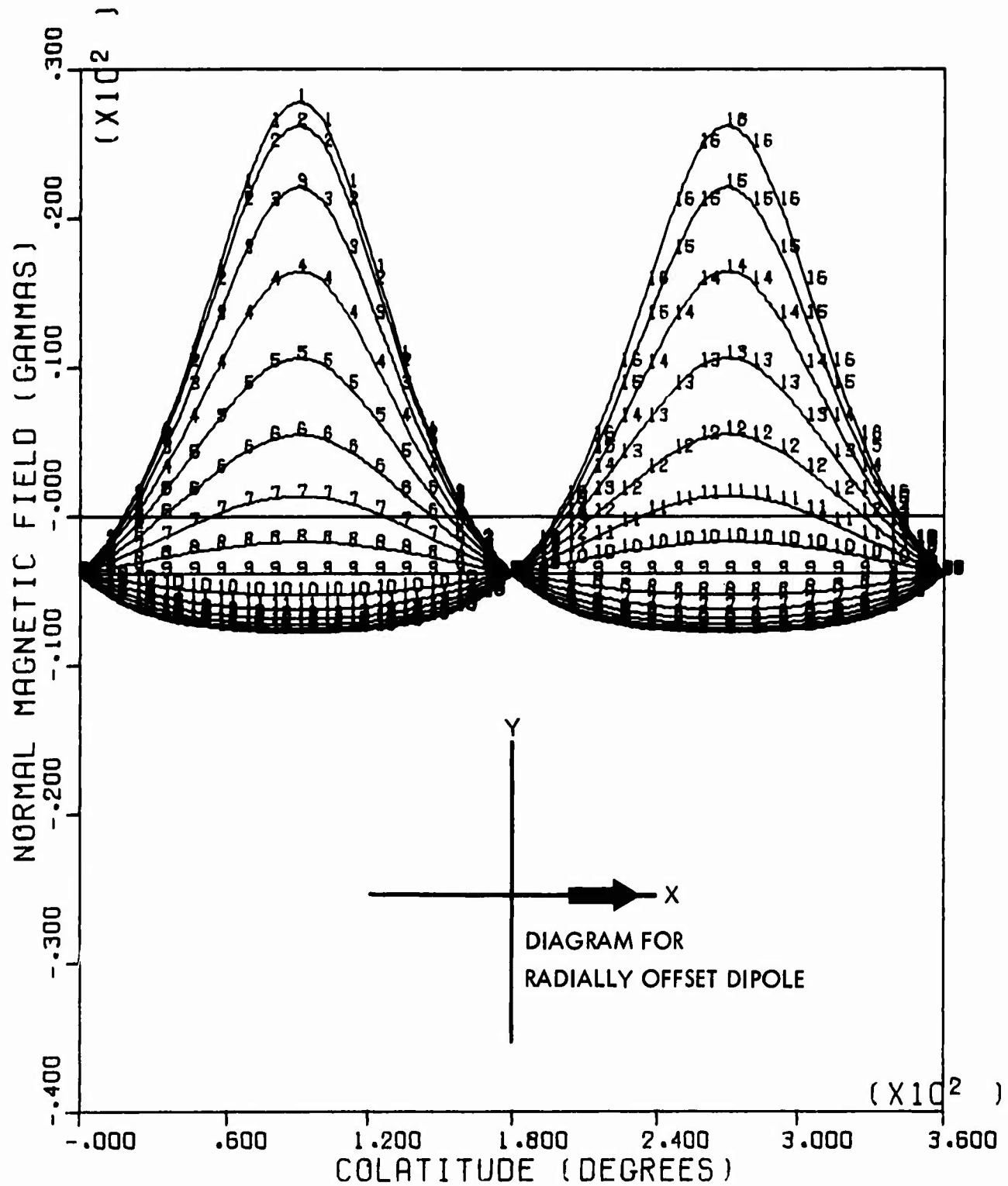


FIG. 36 DATA CURVES FOR RADIALLY OFFSET DIPOLE

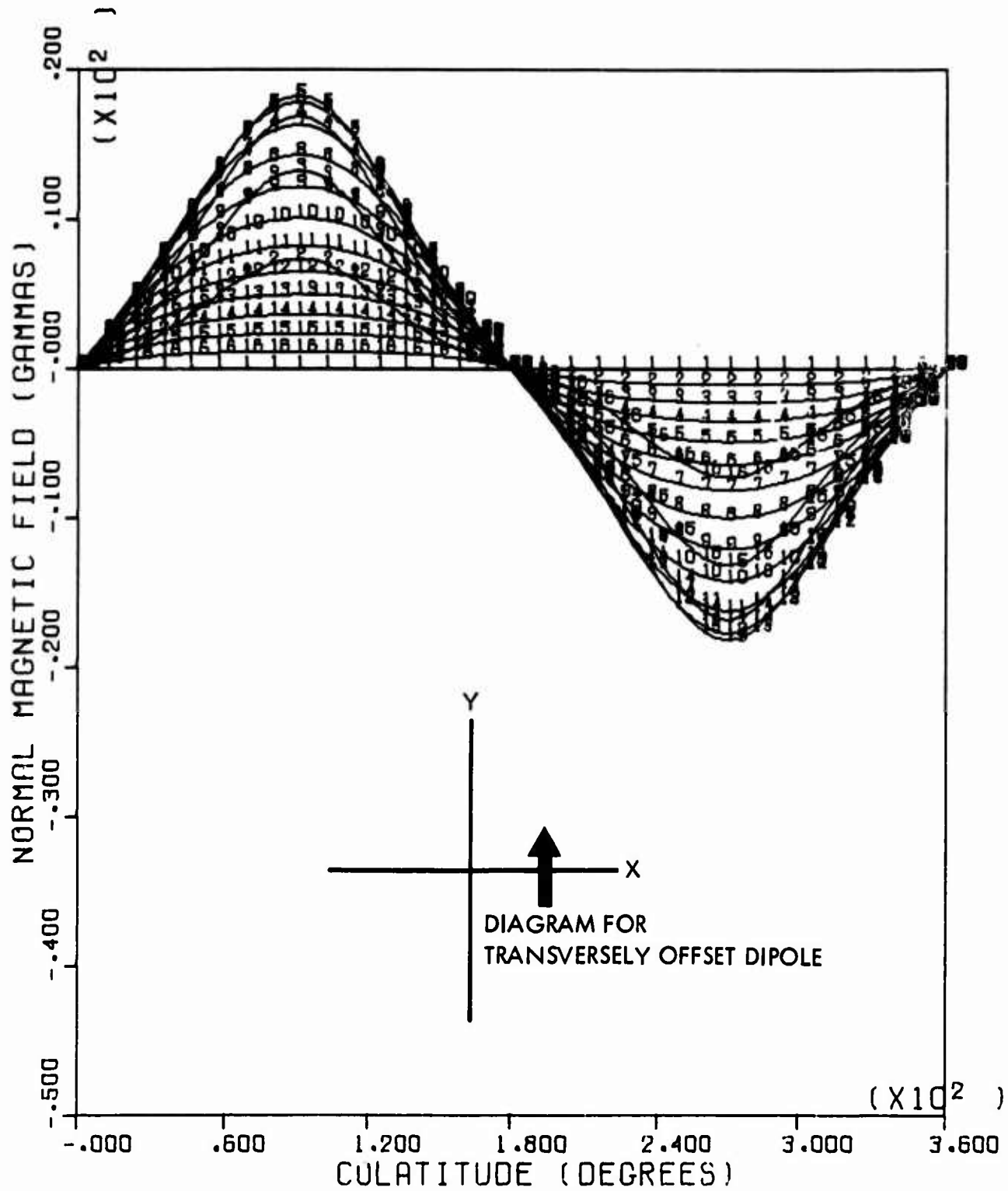


FIG. 37 DATA CURVES FOR TRANSVERSELY OFFSET DIPOLE

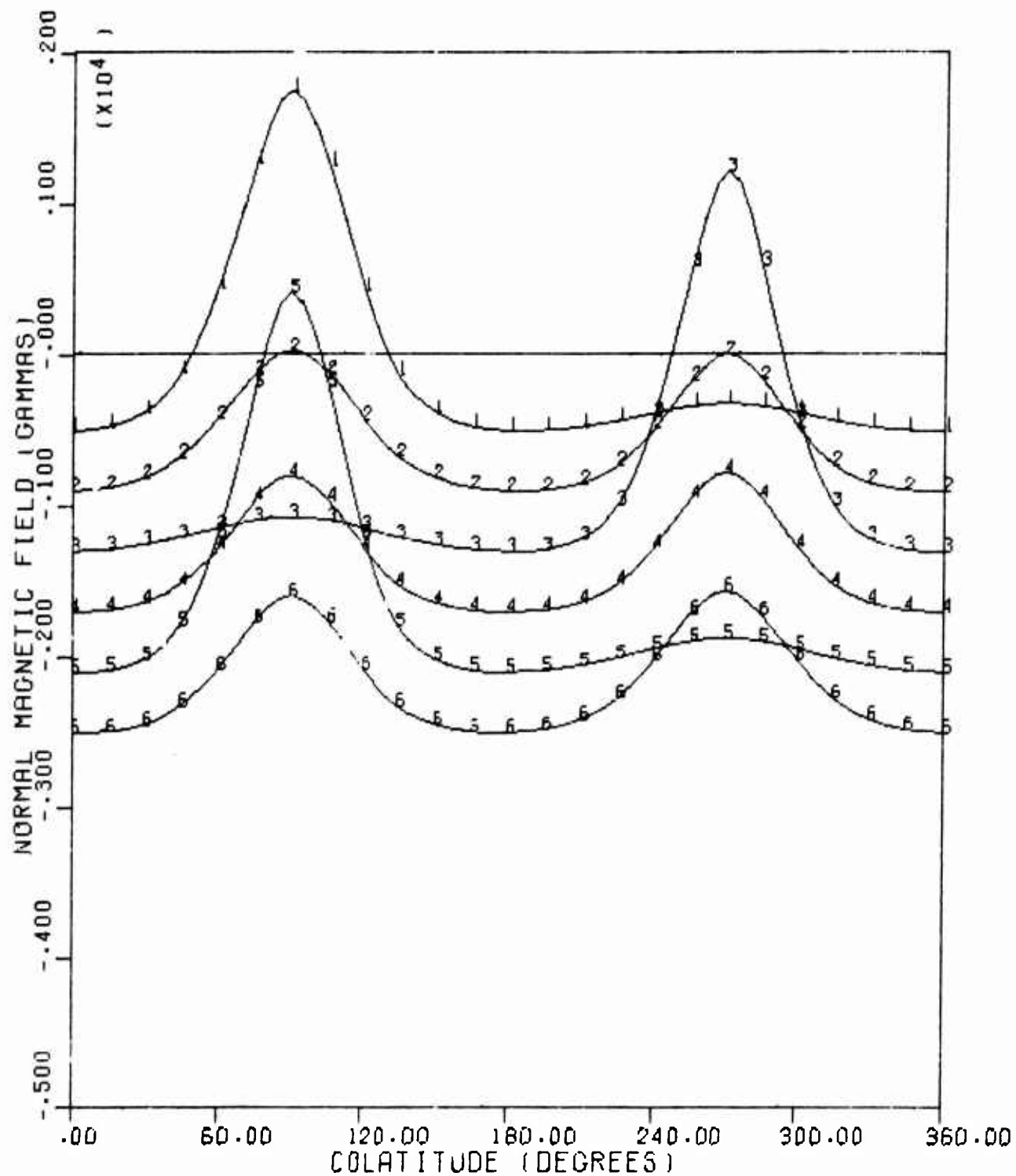


FIG. 38 SAMPLE GRAPH FROM ELECTROSTATIC PLOTTER

APPENDIX A

GLOSSARY OF SYMBOLS AND TERMS

$A(I)$ - is the spherical harmonic coefficient A_n^m in the form used by the computer programs. Eqs. (5) and (8) define the relationships.

$A(n,m)$ - is the spherical harmonic coefficient A_n^m as an element of a two-dimensional array.

A_n^m - is the spherical harmonic coefficient as defined by Eq. (5).

$B(I)$ - is the spherical harmonic coefficient B_{nn}^I in the form used by the computer programs. Eqs. (5) and (8) define the relationships.

$B(n,m)$ - is the spherical harmonic coefficient B_n^m as an element of a two-dimensional array.

B_n^m - is the spherical harmonic coefficient as defined by Eq. (5).

C - is the constant factor in the set of weighting factors $\{Y(I)\}$. C has the value (π/NO) representing the interval of longitude. (see Eqs. (14) and (15)).

$C(I)$ - is the Fourier coefficient representing the amplitude of the I th term in the expansion of the data curves. Eqs. (32) and (34) define the relationships.

C_n^m - is the coefficient used in Eq. (26b). It originates in the identity

$$\frac{dP_n^m(\cos \theta)}{d\theta} = m \cos \theta P_n^m(\cos \theta) / \sin \theta - C_n^m P_{n-1}^{m+1}(\cos \theta)$$

where

$$C_n^m = \left[(n-m)(n+m+1) \right]^{\frac{1}{2}} \text{ for } (m > 0) \\ = \left[n(n+1)/2 \right]^{\frac{1}{2}} \text{ for } (m = 0)$$

- CS - is the parameter in the program SA3024 that represents the data value resulting from a static measurement by the magnetic sensor with a calibration signal turned on.
- CV - is the value (in gammas) of the calibration signal. The values of CS, CV, and CZ are used to convert the measured data into gammas, i.e., each data value is multiplied by $CV/(CS-CZ)$.
- CZ - is the parameter in the program SA3024 that represents the data value resulting from a static measurement by the magnetic sensor with the calibration signal turned off.
- CO - ($= C_0$) is the Fourier coefficient representing the constant component of the data curves. It is defined in Eq. (30).
- Cl_j - is the Fourier coefficient for the j th cosine term in the expansion of the data curves. It is defined in Eq. (31b.).
- C6 - is the parameter in the BASIC program SA1024 which corresponds to CV.
- C7 - is the parameter in the BASIC program SA1024 which corresponds to CS.
- C8 - is the parameter in the BASIC program SA1024 which corresponds to CZ.
- D - is the magnitude of the dipole moment \bar{D} . It is equal to $(D_1^2 + D_2^2 + D_3^2)^{\frac{1}{2}}$.
- \bar{D} - is the dipole moment vector (D_1, D_2, D_3) defined by Eq. (1).
- D(I) - is the I th factor of the set of weighting factors $\{Y(I)\}$. It is a measure of the interval of colatitude assigned to the data points $\{F(I,J), J = 1, 2, \dots, NO\}$. (See Eqs. (14) through (18).)
- D_1, D_2, D_3 - are the x, y, and z-components, respectively, of the dipole moment \bar{D} .
- DEV - is the maximum difference between data points of a measured curve and the corresponding points of the curve defined by the approximated Fourier coefficients.

- EA - is the parameter in programs SA4024 and SA5024 that represents the angular error (in degrees) to be randomly inserted into the simulated measurement data. It represents the measurement position errors.
- ED - is the parameter in programs SA4024 and SA5024 that represents a constant offset (in gammas) of the measurement instrumentation. The analysis procedure handles this value as the magnetic field from a monopole moment.
- EG - is the parameter in programs SA4024 and SA5024 that represents the error (in gammas) to be randomly inserted into the simulated measurement data. It represents instrumentation inaccuracies.
- F(I) - is the argument representing the x, y, and z-component of the magnetic field vector computed by the subroutine AMPFLD. F(I), for $I = 1, 2, \dots, N$ is also the array of data for the subroutine FNCTON.
- F(IJ) - is the single-dimensioned array of data values representing the normal component of the magnetic field on the surface of the measurement sphere surrounding the satellite. The relationship is $F(IJ) = f(\theta(I), \varphi(J))$ for $IJ = I + (J - 1) \cdot N1$, $I = 1, 2, \dots, N1$, and $J = 1, 2, \dots, N0$.
- F(I,J) - is the two-dimensional array of numbers representing discrete values of the normal component of the magnetic field on the surface of the measurement sphere surrounding the satellite. The values are actually stored in the single-dimensioned array $F(IJ) = F(I,J) = f(\theta(I), \varphi(J))$ for $IJ = I + (J - 1) \cdot N1$, $I = 1, 2, \dots, N1$, and $J = 1, 2, \dots, N0$.
- $f(\theta, \varphi)$ - represents the normal component of the total magnetic field on the surface of the measurement sphere of radius $R1$ surrounding the satellite.
- $f_n(\theta, \varphi)$ - represents the normal component of the magnetic field on the surface of the measurement sphere from the multipole moment of degree n .
- $f_{nj}(\theta, \varphi)$ - represents the normal component of the magnetic field on the surface of the measurement sphere from the j th multipole magnet of degree n .
- F9 - is the format for reading and printing the spherical coefficients A(I) and B(I), e.g., F9 = (1H ,7E10.4).

$\vec{H}(\vec{R})$ - represents the total magnetic field vector at the point \vec{R} .

$\vec{H}_n(\vec{R})$ - represents the magnetic field vector at the point \vec{R} from a multipole magnet of degree n .

\vec{i} - represents the unit vector along the x-axis.

IP - is the parameter in the program SA4024 that determines whether or not the simulated measurement data is to be interpolated and plotted. $IP \neq 0$ means that the data will be interpolated and plotted.

IP1 - is the parameter in the programs SA3024 and SA5024 which corresponds to IP.

IP2 - is the parameter in the programs SA3024 and SA5024 that determines whether or not the measurement data is to be printed. $IP2 \neq 0$ means that the data will be printed.

IR - is the parameter in the programs SA2024 and SA3024 that determines whether or not to read the measurement data from a data file. $IR \neq 0$ means that the data will be read from the file DAT024.

IW - is the parameter in programs SA3024 and SA5024 that determines which integrating scheme is to be used. $IW = 0$ means that the exact, algebraic scheme is to be used.

\vec{j} - represents the unit vector along the y-axis.

\vec{k} - represents the unit vector along the z-axis.

m - is the order of the coefficients A_n^m and B_n^m , and the polynomials $P_n^m(\cos \theta)$ and $P_{n,m}(\cos \theta)$.

M - is the program parameter representing the order of the coefficients and polynomials associated with multipole magnets.

n - is the degree of the coefficients A_n^m and B_n^m , and the polynomials $P_n^m(\cos \theta)$ and $P_{n,m}(\cos \theta)$.

- N - is the program parameter representing the degree of the coefficients and polynomials associated with multipole magnets. N is also used to define the number of distinct data points for the subroutine FNCTON.
- NO - is the number of curves of data. It is an even integer (≤ 16) corresponding to the number of great circles of data on the measurement sphere.
- N1 - is the number of equally spaced data points per curve ranging from 0 through 360 degrees in colatitude. N1 is an odd integer (≤ 33).
- N2 - is the parameter in the BASIC program SA1024 which corresponds to IP2. N2 is also used in the subroutine FNCTON to designate the number of Fourier terms that produces the minimum error (DEV) between the data and the computed curve.
- N3 - is the number of distinct factors $\{D(I)\}$ of the weights $\{Y(I)\}$. N3 is the largest integer that is less than or equal to $(N1 + 3)/4$, i.e., $N3 = \left[\max_{1 \leq i} i \right]$ such that $i \leq (N1 + 3)/4$.
- $\bar{n}(\theta, \varphi)$ - is the unit normal vector on (exterior to) the surface of the measurement sphere at the point $(R1, \theta, \varphi)$. It has rectilinear components $(\sin \theta \cos \varphi, \sin \theta \sin \varphi, \cos \theta)$.
- NH - is the parameter in the program SA3024 that represents the highest degree spherical harmonic term to be computed from the data ($NH = 1$ for dipoles, 2 for quadrupoles, etc.). NH is also the parameter in the program SA4024 that represents the total number of different harmonics (degrees) of multipole magnets to be considered.
- NH1 - is the parameter in the program SA5024 that represents the total number of different harmonics (degrees) of multipole magnets to be considered.
- NH2 - is the parameter in the program SA5024 that represents the highest degree spherical harmonic term to be computed from the data ($NH2 = 1$ for dipoles, 2 for quadrupoles, etc.).
- NN - is the parameter in the programs SA4024 and SA5024 that represents the number of multipole magnets with harmonic number (degree) NN.

- NN - is the parameter in the programs SA4024 and SA5024 that represents the harmonic number (degree) for the multipole data being read in (NN = 1 for dipole, 2 for quadrupoles, etc.).
- P - represents the multipole position vector in the programs SA4024 and SA5024. It also represents the weighting factors in the subroutine AMPMNT.
- P4 - is the parameter in programs SA3024, SA4024, and SA5024 that represents the scale factor in gammas/inch for the y-axis if the data is to be plotted. If PY = 0.0, and if the data is to be plotted, a suitable factor will be computed from the data.
- $P_{n,m}(\cos \theta)$ - represents the associated Legendre polynomial of degree n and order m. Reference (b) gives the definition as $P_{n,m}(\cos \theta) \equiv \sin^m \theta \frac{d^m P_n(\cos \theta)}{d(\cos \theta)^m}$ where $P_n(\cos \theta)$ is the regular Legendre polynomial of degree n. (See also Eq. (2) of Appendix K.)
- $P_n^m(\cos \theta)$ - represents the Schmidt polynomial of degree n and order m. They are defined in Eq. (3) of Appendix K.
- PHI(I) - is the Fourier coefficients in the subroutine FNCTON which represent the phase angles defined by Eqs. (31) and (33).
- Q_{11}, \dots, Q_{33} - represent the coefficients of the quadrupole moment term in rectilinear coordinates.
- \bar{r} - represents an element of the orthonormal set of vectors defined by Eq. (23).
- \bar{R} - represents the position vector for a point at which the magnetic field is to be computed by the subroutine AMPFLD. Coordinate transformations are included in Figure 1.
- \bar{R}_1 - represents the vector between a multipole position and the point at which the magnetic field is to be computed by the subroutine AMPFLD.
- R1 - is the parameter in the programs SA1024, ..., SA5024 that represents the radius of the measurement sphere in inches.

Sl_j - is the Fourier coefficient for the j th sine term in the expansion of the data curves. It is defined by Eq. (31a).

$X(I)$ - represents the surface area or weight assigned to the data points $\{F(I,J), J = 1, 2, \dots, NO\}$ when the geometric or approximate integrating scheme is used. It is defined in Eq. (4).

$Y(I)$ - represents the surface area or weight assigned to the data points $\{F(I,J), J = 1, 2, \dots, NO\}$ when either integration scheme is used.

YDIST - is the parameter in the subroutine DATPLT that corresponds to the parameter PY.

θ - represents the spherical coordinate of colatitude defined in terms of rectilinear coordinates as $\theta = \tan^{-1} [(x^2 + y^2)^{1/2} / z]$.

$\bar{\theta}$ - represents an element of the orthonormal set of vectors defined by Eq. (23).

$\theta(I)$ - represents the I th element in the set of values of colatitude at which the data $F(I,J)$ is recorded.

φ - represents the spherical coordinate of longitude defined in terms of rectilinear coordinates as $\varphi = \tan^{-1}(x/y)$.

$\bar{\varphi}$ - represents an element of the orthonormal set of vectors defined by Eq. (23).

$\varphi(J)$ - represents the j th element in the set of values of longitude at which the data $F(I,J)$ is recorded.

APPENDIX B
LISTING OF SA1024

00010 REM PROGRAM SATDPL	SA100010
00020 DIM R(3),W(3),F(300)	SA100020
00030 LET P1=3.14159265	SA100030
00040 READ N0,N1	SA100040
00050 IF N0<>0 THEN 00070	SA100050
00060 STOP	SA100060
00070 READ R1,C6	SA100070
00080 LET R1=R1*2.54	SA100080
00090 PRINT "ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT"	SA100090
00100 READ N2	SA100100
00110 LET D0=2.*P1/(N1-1)	SA100110
00120 FOR I=1 TO 3 STEP 1	SA100120
00130 LET W(I)=0	SA100130
00140 LET R(I)=0	SA100140
00150 NEXT I	SA100150
00160 LET N3=N0*N1	SA100160
00170 READ C7,C8	SA100170
00180 FOR I=1 TO N3 STEP 1	SA100180
00190 READ F(I)	SA100190
00200 NEXT I	SA100200
00210 PRINT	SA100210
00220 GOSUB 01160	SA100220
00230 IF N2=0 THEN 00250	SA100230
00240 GOSUB 01000	SA100240
00250 PRINT	SA100250
00260 FOR K0=1 TO N0 STEP 1	SA100260
00270 LET O3=P1*(K0-1)/N0	SA100270
00280 LET S3=SIN(O3)	SA100280
00290 LET C3=COS(O3)	SA100290
00300 FOR L0=1 TO N1 STEP 1	SA100300
00310 LET K1=L0+(K0-1)*N1	SA100310
00320 LET F1=F(K1)*1.E-5	SA100320
00330 LET O4=2.*P1*(L0-1)/(N1-1)	SA100330
00340 LET S4=SIN(O4)	SA100340
00350 LET C4=COS(O4)	SA100350
00360 LET P5=(SIN(D0/4)**2)/(2.*N0)	SA100360
00370 IF L0=1 THEN 00430	SA100370
00380 IF L0=N1 THEN 00430	SA100380
00390 IF L0<>(N1+1)/2 THEN 00420	SA100390
00400 LET P5=P5*2.	SA100400
00410 GO TO 00430	SA100410
00420 LET P5=ABS(SIN(O4)*SIN(D0/2))/(2.*N0)	SA100420
00430 LET R(1)=C3*S4	SA100430
00440 LET R(2)=S3*S4	SA100440
00450 LET R(3)=C4	SA100450
00460 FOR I=1 TO 3 STEP 1	SA100460
00470 W(I)=W(I)+(F1*P5)*R(I)	SA100470
00480 NEXT I	SA100480

NOLTR 73-191

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00490 NEXT L0
00500 NEXT K0
00510 FOR I=1 TO 3 STEP 1
00520 LET W(I)=(1.5*(R1**3))*W(I)
00530 NEXT I
00540 LET R(1)=SQR(W(1)**2+W(2)**2+W(3)**2)
00550 LET S8=W(2)
00560 LET S9=W(1)
00570 GOSUB 01670
00580 LET R(2)=C5*180./P1
00590 LET S8=SQR(W(1)**2+W(2)**2)
00600 LET S9=W(3)
00610 GOSUB 01670
00620 LET R(3)=C5*180./P1
00630 PRINT "(DX,DY,DZ) = ";W(1);W(2);W(3)
00640 PRINT "(D ,02,01) = ";R(1);R(2);R(3)
00650 GOSUB 01320
00660 GO TO 00040
00670
00680 REM DATA N0,N1
00690 DATA 6,25
00700 REM DATA R1,C6
00710 DATA 96.0,1.0
00720 REM DATA N2
00730 DATA 1
00740 REM DATA C7,C8,(F(I),I=1,N3)
00750 DATA 1.0, 0.0,
00760 DATA -504.3, -475.5, -381.5, -123.3, 438.4, 1273.7, 1749.2,SA100760
00770 DATA 1259.6, 431.3, -125.7, -382.1, -476.3, -502.7, -799.8,SA100770
00780 DATA -476.5, -436.5, -386.3, -341.0, -322.6, -339.4, -386.5,SA100780
00790 DATA -435.9, -476.7, -499.0, -503.7, -902.8, -881.2, -814.6,SA100790
00800 DATA -675.4, -420.8, -118.9, 27.2, -119.7, -426.7, -672.7,SA100800
00810 DATA -815.5, -880.6, -904.1, -897.6, -851.0, -728.7, -483.8,SA100810
00820 DATA -164.5, 7.2, -160.4, -480.6, -728.9, -850.4, -897.1,SA100820
00830 DATA -903.7, -1302.5, -1287.7, -1254.4, -1206.6, -1148.0, -1096.8,SA100830
00840 DATA -1077.9, -1097.1, -1148.4, -1206.6, -1255.4, -1288.3, -1303.0,SA100840
00850 DATA -1292.6, -1219.0, -990.5, -394.6, 589.2, 1219.4, 595.2,SA100850
00860 DATA -406.4, -993.0, -1219.6, -1291.0, -1304.2, -1703.5, -1691.6,SA100860
00870 DATA -1642.3, -1519.4, -1281.9, -961.8, -800.9, -966.5, -1277.0,SA100870
00880 DATA -1520.1, -1642.5, -1692.0, -1704.1, -1690.6, -1640.8, -1515.1,SA100880
00890 DATA -1274.8, -946.5, -779.8, -946.2, -1274.5, -1518.1, -1641.6,SA100890
00900 DATA -1691.1, -1703.7, -2102.4, -2092.4, -2019.9, -1788.4, -1202.2,SA100900
00910 DATA -211.0, 418.3, -207.2, -1209.2, -1792.5, -2019.2, -2091.1,SA100910
00920 DATA -2103.6, -2088.4, -2056.8, -2005.6, -1947.4, -1898.5, -1877.5,SA100920
00930 DATA -1897.9, -1947.7, -2006.3, -2054.5, -2088.6, -2103.4, -2502.4,SA100930
00940 DATA -2497.0, -2452.0, -2327.8, -2086.3, -1770.2, -1598.7, -1762.8,SA100940
00950 DATA -2082.9, -2329.9, -2448.4, -2497.0, -2503.6, -2481.7, -2415.6,SA100950
00960 DATA -2274.3, -2022.5, -1713.6, -1566.7, -1720.3, -2020.7, -2271.2,SA100960
00970 DATA -2414.6, -2481.5, -2503.5
00980 DATA 0, 0,
00990
01000 REM SUBROUTINE PRINTF
01010 FOR K0= 1 TO N0 STEP 1
01020 LET K2=1+(K0-1)*(N1)
01030 LET K3=N1+(K0-1)*(N1)
01040 PRINT
01050 PRINT "CURVE NO."; K0
01060 FOR K4= K2 TO K3 STEP 1
01070 IF F(K4)<1E-2 THEN 01090
01080 GO TO 01110

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NOLTR 73-191

01090 IF F(K4)<-1E-2 THEN 01110	SA101090
01100 LET F(K4)=0.0	SA101100
01110 PRINT F(K4),	SA101110
01120 NEXT K4	SA101120
01130 NEXT K0	SA101130
01140 RETURN	SA101140
01150	SA101150
01160 REM SUBROUTINE SHIFT	SA101160
01170 LET Z1=(F(1)+F(N1))/2.	SA101170
01180 LET N3=(N1+1)/2	SA101180
01190 LET Z2=F(N3)	SA101190
01200 FOR K0= 2 TO N0 STEP 1	SA101200
01210 LET K2=1+(K0-1)*N1	SA101210
01220 LET K3=K0*N1	SA101220
01230 LET K4=N3+(K0-1)*N1	SA101230
01240 LET C1=(2.*Z1+Z2-F(K2)-F(K4)-F(K3))/3.	SA101240
01250 FOR L0=1 TO N1 STEP 1	SA101250
01260 LET K4=L0+(K0-1)*N1	SA101260
01270 LET F(K4)=(F(K4)+C1)*C6/(C7-C8)	SA101270
01280 NEXT L0	SA101280
01290 NEXT K0	SA101290
01300 RETURN	SA101300
01310	SA101310
01320 REM SUBROUTINE PRINT1	SA101320
01330 DIM S(3)	SA101330
01340 LET P1=3.14159265	SA101340
01350 FOR I=1 TO 3 STEP 1	SA101350
01360 LET S(I)=-W(I)	SA101360
01370 NEXT I	SA101370
01380 LET R(1)=SQR(S(1)**2+S(2)**2+S(3)**2)	SA101380
01390 LET S8=S(2)	SA101390
01400 LET S9=S(1)	SA101400
01410 GOSUB 01670	SA101410
01420 LET R(2)=C5*180./P1	SA101420
01430 LET S8=SQR(S(1)**2+S(2)**2)	SA101430
01440 LET S9=S(3)	SA101440
01450 GOSUB 01670	SA101450
01460 LET R(3)=C5*180./P1	SA101460
01470 LET W3=ABS(W(3))	SA101470
01480 LET W4=.2*W3	SA101480
01490 LET A1=SQR(S(1)**2+S(2)**2)	SA101490
01500 LET A2=.2*A1	SA101500
01510 PRINT " THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD";	SA101510
01520 PRINT " BE" A1" GAUSS-"	SA101520
01530 PRINT "CENTIMETER-CUBED WITH THE NORTH POLE POINTING"R(2);	SA101530
01540 PRINT "DEGREES FROM +X."	SA101540
01550 PRINT " (THE MAGNET SHOULD READ"A2" GAMMA AT ONE METER.)"	SA101550
01560 PRINT " THE COMPENSATING MAGNET FOR THE Z-AXIS SHOULD BE";	SA101560
01570 PRINT W3" GAUSS-"	SA101570
01580 IF S(3)>0. THEN 01620	SA101580
01590 PRINT "CENTIMETER-CUBED WITH THE NORTH POLE POINTING";	SA101590
01600 PRINT "TOWARDS -Z (THE MAGNET "	SA101600
01610 GO TO 01640	SA101610
01620 PRINT "CENTIMETER-CUBED WITH THE NORTH POLE POINTING ";	SA101620
01630 PRINT "TOWARDS +Z (THE MAGNET) "	SA101630
01640 PRINT "SHOULD READ"W4" GAMMA AT ONE METER."	SA101640
01650 RETURN	SA101650
01660	SA101660
01670 REM SUBROUTINE ARCTAN	SA101670
01680 IF S9=0 THEN 01750	SA101680

NOLTR 73-191

01690 LET C5=S8/S9
01700 IF S970 THEN 01730
01710 LET C5=ATN(C5)+P1
01720 GO TO 01790
01730 LET C5=ATN(C5)
01740 GO TO 01790
01750 IF S8<0 THEN 01780
01760 LET C5=P1/2
01770 GO TO 01790
01780 LET C5=3*P1/2
01790 RETURN
01800 END

SA101690
SA101700
SA101710
SA101720
SA101730
SA101740
SA101750
SA101760
SA101770
SA101780
SA101790
SA101800

APPENDIX C
LISTING OF SA2024

PROGRAM SATDPL(INPUT,OUTPUT,DAT024,TAPE7=DAT024)	SA200010
DIMENSION R(3),W(3),F(800)	SA200020
P1=3.14159265358979	SA200030
REWIND 7	SA200040
10 READ 01, NO,N1,IR	SA200050
IF(NO.EQ.0) STOP	SA200060
READ 03, R1,C6	SA200070
R1=R1*2.54	SA200080
PRINT 02	SA200090
READ 01, N2	SA200100
D0=2.*P1/FLOAT(N1-1)	SA200110
DO 20 I=1,3	SA200120
W(I)=0	SA200130
R(I)=0	SA200140
20 CONTINUE	SA200150
N3=NO*N1	SA200160
IF(IR.EQ.0) READ 03, C7,C8,(F(I),I=1,N3)	SA200170
IF(IR.EQ.0) WRITE(7,03) C7,C8,(F(I),I=1,N3)	SA200180
IF(IR.NE.0) READ(7,03) C7,C8,(F(I),I=1,N3)	SA200190
CALL SHIFT(F,NO,N1,C6,C7,C8)	SA200200
IF(N2.NE.0) CALL PRINTF(F,NO,N1)	SA200210
DO 60 K0=1,NO	SA200220
O3=P1*(FLOAT(K0-1)/FLOAT(NO))	SA200230
S3=SIN(O3)	SA200240
C3=COS(O3)	SA200250
DO 60 L0=1,N1	SA200260
KL=L0+(K0-1)*N1	SA200270
F1=F(KL)*1.E-5	SA200280
O4=2.*P1*(FLOAT(L0-1)/FLOAT(N1-1))	SA200290
S4=SIN(O4)	SA200300
C4=COS(O4)	SA200310
P5=(SIN(D0/4)**2)/(2.*FLOAT(NO))	SA200320
IF(L0.EQ.1.OR.L0.EQ.N1) GO TO 40	SA200330
IF(L0.NE.(N1+1)/2) GO TO 30	SA200340
P5=P5*2.	SA200350
GO TO 40	SA200360
30 P5=ABS(SIN(O4)*SIN(D0/2))/(2.*FLOAT(NO))	SA200370
40 R(1)=C3*S4	SA200380
R(2)=S3*S4	SA200390
R(3)=C4	SA200400
DO 50 I=1,3	SA200410
50 W(I)=W(I)+(F1*P5)*R(I)	SA200420
60 CONTINUE	SA200430
DO 70 I=1,3	SA200440
W(I)=(1.5*(R1**3))*W(I)	SA200450
70 CONTINUE	SA200460
CALL SPCOOR(W,R)	SA200470
PRINT 04, W(1),W(2),W(3)	SA200480

```

PRINT 05, R(1), R(2), R(3)
CALL PRINT: 5
GO TO 10
01 FORMAT(14I5)
02 FORMAT(44HOENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT9)
03 FORMAT(9F6.1)
04 FORMAT(14H (DX,DY,DZ) = ,3F10.2)
05 FORMAT(14H (D ,O2,O1) = ,3F10.2)
END

SUBROUTINE PRINTF(F,NO,N1)
DIMENSION F(800)
DO 30 KO=1,N0
KS=1+(KO-1)*(N1)
KE=N1+(KO-1)*(N1)
PRINT 11,KO
30 PRINT 12,(F(KK),KK=KS,KE)
RETURN
11 FORMAT(11HO CURVE NO. ,I2)
12 FORMAT(1H ,8F8.1)
END

SUBROUTINE SHIFT(F,NO,N1,CV,CS,CZ)
DIMENSION F(800)
P1=(F(1)+F(N1))/2.
N3=(N1+1)/2
P2=F(N3)
DO 30 KO=1,N0
KS=1+(KO-1)*N1
KE=KO*N1
KK=N3+(KO-1)*N1
COR=(2.*P1+P2-F(KS)-F(KK)-F(KE))/3.
DO 30 LO=1,N1
KK=LO+(KO-1)*N1
30 F(KK)=(F(KK)+COR)*CV/(CS-CZ)
RETURN
END

SUBROUTINE SPCOOR(D,R)
DIMENSION D(3),R(3)
DATA PI/3.14159265358979/
R(1)=SQRT(DOT(D,D))
R(2)=0.
R(3)=0.
IF(R(1).EQ.0.) RETURN
IF(D(1)**2+D(2)**2.NE.0.) R(2)=ATAN2(D(2),D(1))*180./PI
R(3)=ATAN2(SQRT(D(1)**2+D(2)**2),D(3))*180./PI
RETURN
END

FUNCTION DOT(X,Y)
DIMENSION X(3),Y(3)
DOT=X(1)*Y(1)+X(2)*Y(2)+X(3)*Y(3)
RETURN
END

SUBROUTINE PRINTI(W)
DIMENSION R(3),S(3),W(3)
DATA ST,GT/1H-,1H+/
PI=3.14159265

```

SA200490
SA200500
SA200510
SA200520
SA200530
SA200540
SA200550
SA200560
SA200570
SA200580
SA200590
SA200600
SA200610
SA200620
SA200630
SA200640
SA200650
SA200660
SA200670
SA200680
SA200690
SA200700
SA200710
SA200720
SA200730
SA200740
SA200750
SA200760
SA200770
SA200780
SA200790
SA200800
SA200810
SA200820
SA200830
SA200840
SA200850
SA200860
SA200870
SA200880
SA200890
SA200900
SA200910
SA200920
SA200930
SA200940
SA200950
SA200960
SA200970
SA200980
SA200990
SA201000
SA201010
SA201020
SA201030
SA201040
SA201050
SA201060
SA201070
SA201080

DO 10 I=1,3	SA201090
10 S(I)=-W(I)	SA201100
CALL SPCOOR(S,R)	SA201110
W3=ABS(W(3))	SA201120
W32=.2*W3	SA201130
AM1=SQRT(S(1)**2+S(2)**2)	SA201140
AM12=.2*AM1	SA201150
PRINT 11, AM1	SA201160
PRINT 12, R(2)	SA201170
PRINT 13, AM12	SA201180
PRINT 14, W3	SA201190
IF(S(3).LT.0.) PRINT 15, ST	SA201200
IF(S(3).GE.0.) PRINT 15, GT	SA201210
PRINT 16, W32	SA201220
11 FORMAT(54H THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD B	SA201230
+ ,2HE ,F7.1,7H GAUSS-)	SA201240
12 FORMAT(47H CENTIMETER-CUBED WITH THE NORTH POLE POINTING ,F6.1,	SA201250
+17H DEGREES FROM +X.)	SA201260
13 FORMAT(25H (THE MAGNET SHOULD READ ,F7.1,21H GAMMA AT ONE METER.))	SA201270
14 FORMAT(43H THE COMPENSATING MAGNET FOR THE Z-AXIS,	SA201280
+11H SHOULD BE ,F7.1,7H GAUSS-)	SA201290
15 FORMAT(46H CENTIMETER-CUBED WITH THE NORTH POLE POINTING,	SA201300
+9H TOWARDS ,A1,14HZ. (THE MAGNET)	SA201310
16 FORMAT(13H SHOULD READ ,F7.1,22H GAMMA AT ONE METER.))	SA201320
RETURN	SA201330
END	SA201340

APPENDIX D
SAMPLE PROBLEM FOR SA2024

NØL INTERCØM
TYPE "LØGIN."
LØGIN(S)
024533LACK/ /4

09/07/73 09.46.40. BD/42/34
C- SETUP.FØRTRAN

ØN AT 09.46.56. 09/07/73
**FØRTRAN
**NEW ØR ØLD FILE- ATTACH(BN2024,BN2024)*ATTACH(DAT024,DAT024)*TAPE(ØN)

09.47.39.ATTACH(BN2024,BN2024)
09.47.50.ATTACH(DAT024,DAT024)
**READY.

BN2024.

6 25 1
96.00 1.000

0
6 25 0
96.00 1.000

1
10 0 -5043 -4755 -3815 -1233 4384 12737 17492
12596 4313 -1257 -3821 -4763 -5027 -4998 -4765 -4365
-3863 -3410 -3226 -3394 -3865 -4359 -4767 -4990 -5037
-9028 -8812 -8146 -6754 -4208 -1189 272 -1197 -4267
-6727 -8155 -8806 -9041 -8976 -8510 -7287 -4838 -1645
72 -1604 -4806 -7289 -8504 -8971 -9037 -13025 -12877
-12544 -12066 -11480 -10968 -10779 -10971 -11484 -12066 -12554
-12883 -13030 -12926 -12190 -9905 -3946 5892 12194 5952
-4064 -9930 -12196 -12910 -13042 -17035 -16916 -16423 -15194
-12819 -9618 -8009 -9665 -12770 -15201 -16425 -16920 -17041
-16906 -16408 -15151 -12748 -9465 -7798 -9462 -12745 -15181
-16416 -16911 -17037 -21024 -20924 -20199 -17884 -12022 -2110
4183 -2072 -12092 -17925 -20192 -20911 -21036 -20884 -20568
-20056 -19474 -18985 -18775 -18979 -19477 -20063 -20545 -20886
-21034 -25024 -24970 -24520 -23278 -20863 -17702 -15987 -17628
-20829 -23299 -24484 -24970 -25036 -24817 -24156 -22743 -20225
-17136 -15667 -17203 -20207 -22712 -24146 -24815 -25035

Data Tape

0 0 0
TAPE(ØFF)

ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT:

(DX,DY,DZ) = 825.54 -211.43 20.71

(D ,02,01) = 852.44 -14.37 88.61

THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD BE 852.2 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING 165.6 DEGREES FROM +X.

(THE MAGNET SHOULD READ 170.4 GAMMA AT ONE METER.)

THE COMPENSATING MAGNET FOR THE Z-AXIS SHOULD BE 20.7 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING TOWARDS -Z. (THE MAGNET SHOULD READ 4.1 GAMMA AT ONE METER.)

ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT:

CURVE NO. 1

-504.3	-475.5	-381.5	-123.3	438.4	1273.7	1749.2	1259.6
431.3	-125.7	-382.1	-476.3	-502.7	-499.8	-476.5	-436.5
-386.3	-341.0	-322.6	-339.4	-386.5	-435.9	-476.7	-499.0
-503.7							

CURVE NO. 2

-502.8	-481.2	-414.6	-275.4	-20.8	281.1	427.2	280.3
-26.7	-272.7	-415.5	-480.6	-504.1	-497.6	-451.0	-328.7
-83.8	235.5	407.2	239.6	-80.6	-328.9	-450.4	-497.1
-503.7							

CURVE NO. 3

-502.8	-488.0	-454.7	-406.9	-348.3	-297.1	-278.2	-297.4
-348.7	-406.9	-455.7	-488.6	-503.3	-492.9	-419.3	-190.8
405.1	1388.9	2019.1	1394.9	393.3	-193.3	-419.9	-491.3
-504.5							

CURVE NO. 4

-503.3	-491.4	-442.1	-319.2	-81.7	238.4	399.3	233.7
-76.8	-319.9	-442.3	-491.8	-503.9	-490.4	-440.6	-314.9
-74.6	253.7	420.4	254.0	-74.3	-317.9	-441.4	-490.9
-503.5							

CURVE NO. 5

-502.8	-492.8	-420.3	-188.8	397.4	1388.6	2017.9	1392.4
390.4	-192.9	-419.6	-491.5	-504.0	-488.8	-457.2	-406.0
-347.8	-298.9	-277.9	-298.3	-348.1	-406.7	-454.9	-489.0
-503.8							

CURVE NO. 6

-502.8	-497.4	-452.4	-328.2	-86.7	229.4	400.9	236.8
-83.3	-330.3	-448.8	-497.4	-504.0	-482.1	-416.0	-274.7
-22.9	286.0	432.9	279.3	-21.1	-271.6	-415.0	-481.9
-503.9							

(DX,DY,DZ) = 825.54 -211.43 20.71

(D ,02,01) = 852.44 -14.37 88.61

THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD BE 852.2 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING 165.6 DEGREES FROM +X.

(THE MAGNET SHOULD READ 170.4 GAMMA AT ONE METER.)

THE COMPENSATING MAGNET FOR THE Z-AXIS SHOULD BE 20.7 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING TOWARDS -Z. (THE MAGNET SHOULD READ 4.1 GAMMA AT ONE METER.)

09.53.46.STOP

**READY.

LOGOUT.

CP TIME 1.187

PP TIME 48.498

CONNECT TIME 0 HR 9 MIN 10 SEC

TOTAL COST OF SESSION = \$ 2.28

09/07/73 LOGGED OUT AT 09.55.50.<

Notes:

1. The file BN2024 is the binary version of SA2024. It consists of six binary records (subprograms).
2. The information typed in by the user has been underlined.

APPENDIX E

LISTING OF SA3024

```

PROGRAM DIPANL(INPUT=65,OUTPUT=65,DAT024=55,TAPE5=INPUT,
+ TAPE6=OUTPUT,TAPE7=DAT024,TAPE99)
C
C          SATELLITE ANALYSIS PROGRAM
C
C  THIS PROGRAM ANALYZES MAGNETIC DATA REPRESENTING THE NORMAL COMPON-
C  ENT OF THE MAGNETIC FIELD FROM A SATELLITE. THE DATA IS ENTERED IN
C  THE FOLLOWING ORDER --
C
C  N0 - THE NUMBER OF GREAT CIRCLES OF DATA. (N0 IS USUALLY EVEN, E.G.,
C      N0=(N1-1)/2. THE PROGRAM STOPS IF N0=0.)
C
C  N1 - THE NUMBER OF DATA POINTS PER GREAT CIRCLE. (N1 IS ALWAYS ODD.
C      THE FIRST DATA POINT IS THE SAME AS THE LAST FOR EACH GREAT CIR-
C      LE.)
C
C  NH - THE HARMONIC NUMBER (DEGREE) REPRESENTING THE HIGHEST DEGREE
C      SPHERICAL HARMONIC TERM TO BE COMPUTED FROM THE DATA. (NH=1 FOR
C      DIPOLES, 2 FOR QUADRUPOLES, ETC.)
C
C  IR - DETERMINES WHETHER OR NOT TO READ THE DATA FROM THE FILE
C      DAT024. (IR=0 MEANS THAT THE DATA WILL NOT BE READ FROM
C      THE FILE.)
C
C  IP1 - DETERMINES WHETHER OR NOT THE MAGNETIC DATA IS TO BE PLOTTED.
C      (IP1=0 MEANS THAT THE DATA WILL NOT BE PLOTTED.)
C
C  IW - DETERMINES THE TYPE OF INTEGRATING SCHEME TO BE USED. (IW=0
C      MEANS THAT THE EXACT, ALGEBRAIC SCHEME IS TO BE USED.)
C
C  R1 - THE RADIUS OF THE MEASUREMENT SPHERE IN INCHES.
C
C  CV - THE VALUE OF THE CAL SIGNAL (CS-CZ) IN GAMMAS.
C
C  PY - THE SCALE FACTOR (GAMMAS/INCH) FOR THE Y-AXIS IF THE DATA IS TO
C      BE PLOTTED. (IF PY=0.0 A FACTOR WILL BE COMPUTED FROM THE DATA.)
C
C  IP2 - DETERMINES WHETHER OR NOT THE MAGNETIC DATA IS TO BE PRINTED.
C      (IP=0 MEANS THAT THE DATA WILL NOT BE PRINTED.)
C
C  CS - THE STATIC MEASUREMENT WITH THE CAL SIGNAL.
C
C  CZ - THE STATIC MEASUREMENT WITHOUT THE CAL SIGNAL.
C
C  F(1) - THE DATA REPRESENTING THE NORMAL COMPONENT OF THE MAGNETIC
C          FIELD ALONG GREAT CIRCLES ON THE MEASUREMENT SPHERE.

```

```

SA300018
SA300020
SA300030
SA300040
SA300050
SA300060
SA300070
SA300080
SA300090
SA300100
SA300110
SA300120
SA300130
SA300140
SA300150
SA300160
SA300170
SA300180
SA300190
SA300200
SA300210
SA300220
SA300230
SA300240
SA300250
SA300260
SA300270
SA300280
SA300290
SA300300
SA300310
SA300320
SA300330
SA300340
SA300350
SA300360
SA300370
SA300380
SA300390
SA300400
SA300410
SA300420
SA300430
SA300440
SA300450
SA300460
SA300470
SA300480

```


		SA300490
		SA300500
C	*****	SA300510
	DIMENSION F(1000),PV(200),A(11),B(10),R(3),D(3)	SA300520
	NAMelist/NAM/NO,N1,NH,IR,IP1,IW,R1,CV,PY	SA300530
	DATA P1/3.14159265358979/	SA300540
	REWIND 7	SA300550
	IPT=0	SA300560
10	READ(5,01) NO,N1,NH,IR,IP1,IW	SA300570
	IF(IPT.EQ.0.AND.IP1.NE.0) CALL CALCM1(0,10H024 LACKEY,-10.)	SA300580
	IF(NO.EQ.0) GO TO 90	SA300590
	IF(IP1.NE.0) IPT=1	SA300600
	READ(5,02) R1,CV,PY	SA300610
	WRITE(6,NAM)	SA300620
	WRITE(6,03)	SA300630
	READ(5,01) IP2	SA300640
	N3=NO*N1	SA300650
	IF(IR.EQ.0) READ(5,04) CS,CZ,(F(I),I=1,N3)	SA300660
	IF(IR.EQ.0) WRITE(7,04) CS,CZ,(F(I),I=1,N3)	SA300670
	IF(IR.NE.0) READ(7,04) CS,CZ,(F(I),I=1,N3)	SA300680
	CALL SHIFT(F,NO,N1,CV,CS,CZ)	SA300690
	IF(IW.EQ.0) CALL WGT1(NO,N1,PV)	SA300700
	IF(IW.NE.0) CALL WGT3(NO,N1,PV)	SA300710
	CALL AMPMNT(NO,N1,0,R1,F,PV,A,B)	SA300720
	A0=A(1)*1.E-5	SA300730
	F0=A0/((R1*2.54)**2)	SA300740
	DO 20 I=1,N3	SA300750
20	F(I)=F(I)*1.E-5-F0	SA300760
	WRITE(6,05) A0	SA300770
	IF(IP2.NE.0) CALL PRINTF(F,NO,N1)	SA300780
	IF(IP1.NE.0) CALL DATPLT(NO,N1,F,PY)	SA300790
	DO 80 NI=1,NH	SA300800
	CALL AMPMNT(NO,N1,NI,R1,F,PV,A,B)	SA300810
	IF(NI.GT.1) GO TO 60	SA300820
	D(1)=A(1)	SA300830
	D(2)=B(1)	SA300840
	D(3)=A(1)	SA300850
	CALL SPCOOR(D,R)	SA300860
	WRITE(6,06) D	SA300870
	WRITE(6,07) R	SA300880
	CALL PRINTI(D)	SA300890
	GO TO 70	SA300900
60	IF(NI.GT.2) GO TO 70	SA300910
	S3=SQRT(3.)	SA300920
	Q11=S3*A(3)-A(1)	SA300930
	Q22=-S3*A(3)-A(1)	SA300940
	Q33=-Q11-Q22	SA300950
	Q12=S3*B(2)	SA300960
	Q13=S3*A(2)	SA300970
	Q23=S3*B(1)	SA300980
	WRITE(6,09) Q11,Q22,Q33,Q12,Q13,Q23	SA300990
70	WRITE(6,11) NI	SA301000
	NI1=NI+1	SA301010
	WRITE(6,12) (A(I),I=1,NI1)	SA301020
	WRITE(6,12) (B(I),I=1,NI)	SA301030
80	CONTINUE	SA301040
	GO TO 10	SA301050
90	IF(IPT.NE.0) CALL CALCM1(0,10H024 LACKEY,+10.)	SA301060
	STOP	SA301070
01	FORMAT(14I5)	SA301080

```

02 FORMAT(9F8.4)
03 FORMAT(44H0ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT9)
04 FORMAT(9F6.1)
06 FORMAT(14H (DX,DY,DZ) = ,3F10.2)
05 FORMAT(26H0THE MONOPOLE MOMENT IS --,F20.6//)
07 FORMAT(14H (D ,Q2,Q1) = ,3F10.2)
09 FORMAT(54H0THE QUADRUPOLE MOMENTS Q11,Q22,Q33,Q12,Q13,Q23 ARE --/
+ (3E20.12))
11 FORMAT(27H0THE A ( B COEFFS. FOR THE ,12,16TH HARMONIC ARE9)
12 FORMAT(1H ,5E20.12)
END

SUBROUTINE AMPMNT(N0,N1,N,R1,F,P,A,B)
DIMENSION F(1000),P(200),PN(10,200),A(11),B(10)
DATA PI/3.14159265358979323846/
CALL POLVAL(N1,N,PN)
KL=0
N2=(N1+1)/2
D1=PI/FLOAT(N0)
D2=2.*PI/FLOAT(N1-1)
DO 10 K=1,N0
A1=D1*FLOAT(K-1)
LP=+1
LI=0
DO 10 L=1,N1
KL=KL+1
IF(L.GT.N2) LP=-1
LI=LI+LP
A2=A1+FLOAT(1-LP)*PI/2.
F1=F(KL)
A2N=FLOAT(2*N+1)
A1N=FLOAT(N+1)
AR=(R1*2.54)**(N+2)
NN=N+1
DO 10 MM=1,NN
M=MM-1
AM=FLOAT(M)
IF(KL.EQ.1) A(MM)=0.
IF(KL.EQ.1.AND.M.GT.0) B(M)=0.
PNM1=PN(MM,LI)*COS(AM*A2)
PNM2=PN(MM,LI)*SIN(AM*A2)
AC=(A2N*AR/A1N)*F1*P(LI)
A(MM)=A(MM)+AC*PNM1
IF(M.GT.0) B(M)=B(M)+AC*PNM2
10 CONTINUE
RETURN
END

SUBROUTINE WGT1(N0,N1,P)
DIMENSION D(50),P(200)
N4=(N1+1)/4
AN=FLOAT(N0)
CALL WGT2(N1,D)
DO 10 J1=1,N4
J3=N1/2+2-J1
DJ=D(J1)/(4.*AN)
P(J1)=DJ
10 P(J3)=DJ
IF(INT(FLOAT(N1-1)/4.+0.1)*4.NE.(N1-1)) GO TO 20
J1=N4+1

```

SA301090
SA301100
SA301110
SA301120
SA301130
SA301140
SA301150
SA301160
SA301170
SA301180
SA301190
SA301200
SA301210
SA301220
SA301230
SA301240
SA301250
SA301260
SA301270
SA301280
SA301290
SA301300
SA301310
SA301320
SA301330
SA301340
SA301350
SA301360
SA301370
SA301380
SA301390
SA301400
SA301410
SA301420
SA301430
SA301440
SA301450
SA301460
SA301470
SA301480
SA301490
SA301500
SA301510
SA301520
SA301530
SA301540
SA301550
SA301560
SA301570
SA301580
SA301590
SA301600
SA301610
SA301620
SA301630
SA301640
SA301650
SA301660
SA301670
SA301680

```

P(J1)=D(J1)/(4.*AN)
20 N4=(N1+1)/2
P(N4)=P(N4)*2.
RETURN
END

SUBROUTINE WGT2(N,D)
DIMENSION A(50,50),D(50),C(50)
DOUBLE PRECISION A,C,PI,AN
DATA PI/3.141592653589793238462643D0/
N3=(N+1)/4
N4=N3+1
DO 10 I=2,N4
C(I)=1./FLOAT(2*I-1)
A(I,1)=1.
A(I,N4)=0.
DO 10 J=2,N3
AN=2.*PI*DBLE(FLOAT(J-1))/DBLE(FLOAT(N-1))
10 A(I,J)=(DCOS(AN))*((2*I-2)
C(I)=2.
A(I,N4)=1.
DO 20 J=1,N3
20 A(I,J)=2.
N3=N4
IF(INT(FLOAT(N-1)/4.+1)*4.NE.(N-1)) N3=N3-1
CALL GAUSEL(A,C,N3,N3,1)
DO 30 I=1,N3
30 D(I)=C(I)
RETURN
END

SUBROUTINE WGT3(N0,N1,P)
DIMENSION P(200)
DATA PI/3.14159265358979323846/
D2=2.*PI/FLOAT(N1-1)
E1=(SIN(D2/4.))*2
E2=SIN(D2/2.)
P(1)=E1/(2.*FLOAT(N0))
N2=(N1+1)/2
P(N2)=P(1)*2.
IF(N1.LE.3) RETURN
N3=N2-1
DO 10 L=2,N3
S2=SIN(D2*FLOAT(L-1))
P(L)=A*S(S2*E2)/(2.*FLOAT(N0))
10 CONTINUE
RETURN
END

SUBROUTINE GAUSEL(A,C,M,N,IT)
DIMENSION A(50,50),C(50),IB(50)
DOUBLE PRECISION A,C,D
C WRITE(6,03)
DO 20 I=1,M
C 20 WRITE(6,04) (A(I,J),J=1,N),C(I)
DO 70 I=1,N
70 IB(I)=I
DO 60 I=1,M
IF(IT.NE.0.AND.M.NE.N) GO TO 35
D=0.

```

SA301690
SA301700
SA301710
SA301720
SA301730
SA301740
SA301750
SA301760
SA301770
SA301780
SA301790
SA301800
SA301810
SA301820
SA301830
SA301840
SA301850
SA301860
SA301870
SA301880
SA301890
SA301900
SA301910
SA301920
SA301930
SA301940
SA301950
SA301960
SA301970
SA301980
SA301990
SA302000
SA302010
SA302020
SA302030
SA302040
SA302050
SA302060
SA302070
SA302080
SA302090
SA302100
SA302110
SA302120
SA302130
SA302140
SA302150
SA302160
SA302170
SA302180
SA302190
SA302200
SA302210
SA302220
SA302230
SA302240
SA302250
SA302260
SA302270
SA302280

```

JJ=0
DO 10 J=1,N
IF(DABS(A(I,J)).LE.D) GO TO 10
JJ=J
D=DABS(A(I,J))
10 CONTINUE
IF(JJ.EQ.1) GO TO 35
IF(JJ.NE.0) GO TO 140
WRITE(6,01) I
GO TO 110
140 DO 160 J=1,M
D=A(J,JJ)
A(J,JJ)=A(J,I)
160 A(J,I)=D
ID=IB(JJ)
IB(JJ)=IB(I)
IB(I)=ID
35 D=A(I,1)
DO 30 J=1,M
30 A(I,J)=A(I,J)/D
C(I)=C(I)/D
DO 50 J=1,M
IF(J.EQ.1) GO TO 50
D=A(J,1)
IF(D.EQ.0.) GO TO 50
DO 40 K=1,N
40 A(J,K)=A(J,K)-D*A(I,K)
C(J)=C(J)-D*C(I)
50 CONTINUE
60 CONTINUE
110 IF(1T.EQ.0.OR.M.NE.N) GO TO 100
DO 120 I=1,N
120 A(I,N)=C(I)
DO 130 I=1,N
II=IB(I)
130 C(II)=A(I,N)
100 CONTINUE
C IF(1T.EQ.0) WRITE(6,02) (IB(I),I=1,N)
C WRITE(6,05)
C DO 80 I=1,M
C 80 WRITE(6,04) (A(I,J),J=1,N),C(I)
RETURN
01 FORMAT(5H1ROW ,I2,14H IS ALL ZEROS.)
02 FORMAT(24H0SINGLE PERMUTATION IS.,20I3//)
03 FORMAT(21H0THE INPUT MATRIX IS9//)
04 FORMAT(1H ,7D16.8)
05 FORMAT(22H0THE OUTPUT MATRIX IS9//)
END

SUBROUTINE POLVAL(N1,N,PN)
DIMENSION PN(10,200)
DATA PI/3.14159265258979323846/
D2=2.*PI/FLOAT(N1-1)
N2=(N1+1)/2
DO 10 I=1,N2
A2=D2*FLOAT(I-1)
C2=COS(A2)
NN=N+1
DO 10 MM=1,NN
M=MM-1

```

SA302290
SA302300
SA302310
SA302320
SA302330
SA302340
SA302350
SA302360
SA302370
SA302380
SA302390
SA302400
SA302410
SA302420
SA302430
SA302440
SA302450
SA302460
SA302470
SA302480
SA302490
SA302500
SA302510
SA302520
SA302530
SA302540
SA302550
SA302560
SA302570
SA302580
SA302590
SA302600
SA302610
SA302620
SA302630
SA302640
SA302650
SA302660
SA302670
SA302680
SA302690
SA302700
SA302710
SA302720
SA302730
SA302740
SA302750
SA302760
SA302770
SA302780
SA302790
SA302800
SA302810
SA302820
SA302830
SA302840
SA302850
SA302860
SA302870
SA302880

```

      PN(MM,I)=SPNM(N,M,CZ)
10 CONTINUE
      RETURN
      END

      SUBROUTINE SHIFT(F,N0,N1,CV,CS,CZ)
      DIMENSION F(800)
      P1=(F(1)+F(N1))/2.
      N3=(N1+1)/2
      P2=F(N3)
      DO 10 K0=1,N0
      KS=1+(K0-1)*N1
      KE=K0*N1
      KK=N3+(K0-1)*N1
      COR=(2.*P1+P2-F(KS)-F(KK)-F(KE))/3.
      DO 10 L0=1,N1
      KK=L0+(K0-1)*N1
10 F(KK)=(F(KK)+COR)*CV/(CS-CZ)
      RETURN
      END

      SUBROUTINE PRINTI(W)
      DIMENSION R(3),S(3),W(3)
      DATA ST,GT/1H-,1H+/
      P1=3.14159265
      DO 10 I=1,3
10 S(I)=-W(I)
      R(1)=SQRT(S(1)**2+S(2)**2+S(3)**2)
      R(2)=ATAN2(S(2),S(1))*180./P1
      R(3)=ATAN2(SQRT(S(1)**2+S(2)**2),S(3))*180./P1
      W3=ABS(W(3))
      W32=.2*W3
      AM1=SQRT(S(1)**2+S(2)**2)
      AM12=.2*AM1
      PRINT 11, AM1
      PRINT 12, R(2)
      PRINT 13, AM12
      PRINT 14, W3
      IF(S(3).LT.0.) PRINT 15, ST
      IF(S(3).GE.0.) PRINT 15, GT
      PRINT 16, W32
11 FORMAT(54H      THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD B
      +,2HE ,F7.1,7H GAUSS-)
12 FORMAT(47H CENTIMETER-CUBED WITH THE NORTH POLE POINTING ,F6.1,
      +17H DEGREES FROM +X.)
13 FORMAT(25H (THE MAGNET SHOULD READ ,F7.1,21H GAMMA AT ONE METER.))
14 FORMAT(43H      THE COMPENSATING MAGNET FOR THE Z-AXIS,
      +11H SHOULD BE ,F7.1,7H GAUSS-)
15 FORMAT(46H CENTIMETER-CUBED WITH THE NORTH POLE POINTING,
      +9H TOWARDS ,A1,14HZ. (THE MAGNET)
16 FORMAT(13H SHOULD READ ,F7.1,22H GAMMA AT ONE METER.))
      RETURN
      END

      FUNCTION SPNM(N,M,X)
      SPNM=PNM(N,M,X)
      IF(M.EQ.0.OR.M.GT.N) RETURN
      SPNM=SQRT(2.*ANF(N-M)/ANF(N+M))*SPNM
      RETURN
      END

```

SA302890
 SA302900
 SA302910
 SA302920
 SA302930
 SA302940
 SA302950
 SA302960
 SA302970
 SA302980
 SA302990
 SA303000
 SA303010
 SA303020
 SA303030
 SA303040
 SA303050
 SA303060
 SA303070
 SA303080
 SA303090
 SA303100
 SA303110
 SA303120
 SA303130
 SA303140
 SA303150
 SA303160
 SA303170
 SA303180
 SA303190
 SA303200
 SA303210
 SA303220
 SA303230
 SA303240
 SA303250
 SA303260
 SA303270
 SA303280
 SA303290
 SA303300
 SA303310
 SA303320
 SA303330
 SA303340
 SA303350
 SA303360
 SA303370
 SA303380
 SA303390
 SA303400
 SA303410
 SA303420
 SA303430
 SA303440
 SA303450
 SA303460
 SA303470
 SA303480

```

FUNCTION PNM(N,M,X)
PNM=0.
IF(M.GT.N) RETURN
IF((ABS(1.-ABS(X))).GT.1.E-9) GO TO 20
IF(M.NE.0) RETURN
PNM=-1.
IF(X.GT.0..OR.N.EQ.INT(FLOAT(N)/2.+1)*2) PNM=1.
RETURN
20 CNM=(2.**N)*ANF(N)*ANF(N-M)
PNM=1.
IF(M.NE.0) PNM=SQRT(1.-X*X)**M
CNM=ANF(2*N)*PNM/CNM
PNM=1.
IF(M.NE.N) PNM=X**(N-M)
IF(N-M.LE.1) GO TO 40
PRD1=1.
NT=(N-M)/2
DO 30 I=1,NT
AN1=FLOAT(N-M-2*I+2)
AN2=FLOAT(2*I)
AN3=FLOAT(2*N-2*I+1)
PRD1=-PRD1*AN1*(AN1-1)/(AN2*AN3)
NE=N-M-2*I
AN1=1.
IF(NE.GT.0) AN1=X**NE
PNM=PNM+PRD1*AN1
30 CONTINUE
40 PNM=CNM*PNM
RETURN
END

FUNCTION ANF(N)
DOUBLE PRECISION AN
ANF=1.
AN=1.DO
IF(N.LT.0) PRINT 01
IF(N.LT.2) RETURN
DO 10 I=2,N
10 AN=AN*DBLE(FLOAT(I))
ANF=SNGL(AN)
RETURN
01 FORMAT(37H1FACTORIAL INTEGER IS LESS THAN ZERO.//)
END

SUBROUTINE SPCOOR(D,R)
DIMENSION D(3),R(3)
DATA PI/3.14159265358979/
R(1)=SQRT(DOT(D,D))
R(2)=0.
R(3)=0.
IF(R(1).EQ.0.) RETURN
IF(D(1)**2+D(2)**2.NE.0.) R(2)=ATAN2(D(2),D(1))*180./PI
R(3)=ATAN2(SQRT(D(1)**2+D(2)**2),D(3))*180./PI
RETURN
END

SUBROUTINE SUM(A,X,B,Y,Z)
DIMENSION X(3),Y(3),Z(3)
DO 10 I=1,3

```

SA303490
SA303500
SA303510
SA303520
SA303530
SA303540
SA303550
SA303560
SA303570
SA303580
SA303590
SA303600
SA303610
SA303620
SA303630
SA303640
SA303650
SA303660
SA303670
SA303680
SA303690
SA303700
SA303710
SA303720
SA303730
SA303740
SA303750
SA303760
SA303770
SA303780
SA303790
SA303800
SA303810
SA303820
SA303830
SA303840
SA303850
SA303860
SA303870
SA303880
SA303890
SA303900
SA303910
SA303920
SA303930
SA303940
SA303950
SA303960
SA303970
SA303980
SA303990
SA304000
SA304010
SA304020
SA304030
SA304040
SA304050
SA304060
SA304070
SA304080

```

10 Z(I)=A*X(I)+B*Y(I)
RETURN
END

FUNCTION DOT(X,Y)
DIMENSION X(3),Y(3)
DOT=X(1)*Y(1)+X(2)*Y(2)+X(3)*Y(3)
RETURN
END

SUBROUTINE PRINTF(F,N0,N1)
DIMENSION F(800),F1(100)
DO 20 K0=1,N0
DO 10 I=1,N1
II=I+(K0-1)*N1
10 F1(I)=F(II)*1.E5
WRITE(6,01) K0
20 WRITE(6,02) (F1(KK),KK=1,N1)
01 FORMAT(11HOCURVE NO. ,I2)
02 FORMAT(1H ,8F8.1)
RETURN
END

SUBROUTINE DATPLT(N0,N1,F,YDIST)
DIMENSION F(1000),F1(33),F2(73),X(33),X1(73),C(1),P(1),CH(16)
DATA PI/3.14159265358979/
DATA CH/1H1,1H2,1H3,1H4,1H5,1H6,1H7,1H8,1H9,2H10,2H11,2H12,2H13,
+ 2H14,2H15,2H16/
N3=N0*N1
CALL YSCALE(N3,F,YDIST,YMAX,YMIN)
DX=360./FLOAT(N1-1)
F1(1)=F1(2)=YMAX $ X(1)=0. $ X(2)=360.
CALL CALCM1(2,X,F1,0,0.,360.,YMIN,YMAX,6.,7.,TITLE,0.
+ 20HCOLATITUDE (DEGREES),-20,30HNORMAL MAGNETIC FIELD (GAMMAS),
+ 30,0.,18)
F1(2)=YMIN $ X(1)=X(2)=360.
CALL CALCM1(-2,X,F1,0)
F1(1)=F1(2)=0. $ X(1)=360. $ X(2)=0.
IF(YMAX.GT.0.0.AND.YMIN.LT.0.0) CALL CALCM1(-2,X,F1,0)
DO 10 I=1,N1
10 X(I)=DX*FLOAT(I-1)
DO 20 I=1,73
20 X1(I)=5.*FLOAT(I-1)
IJ=0
DO 50 I=1,N0
DO 30 J=1,N1
IJ=IJ+1
30 F1(IJ)=F(IJ)*1.E+5
CALL FNCTON(F1,N1-1,C,C0,P,N2,DEV)
DO 40 N=1,N2
EN=N
DO 40 J=1,73
XX=X1(J)*PI/180.
IF(N.EQ.1) F2(J)=C0
40 F2(J)=F2(J)+C(N)*SIN(EN*XX+P(N))
CALL CALCM1(-73,X1,F2,0)
DO 50 J=1,N1
NC=1+I/10
CALL SYMBL4(X(J)/60.,(F1(J)-YMIN)/YDIST,.08,CH(I),0.,NC)
50 CONTINUE

```

SA304090
SA304100
SA304110
SA304120
SA304130
SA304140
SA304150
SA304160
SA304170
SA304180
SA304190
SA304200
SA304210
SA304220
SA304230
SA304240
SA304250
SA304260
SA304270
SA304280
SA304290
SA304300
SA304310
SA304320
SA304330
SA304340
SA304350
SA304360
SA304370
SA304380
SA304390
SA304400
SA304410
SA304420
SA304430
SA304440
SA304450
SA304460
SA304470
SA304480
SA304490
SA304500
SA304510
SA304520
SA304530
SA304540
SA304550
SA304560
SA304570
SA304580
SA304590
SA304600
SA304610
SA304620
SA304630
SA304640
SA304650
SA304660
SA304670
SA304680

```

RETURN
END

SUBROUTINE FNCTON(F,NSCANS,C,CO,PHI,N2,DEV)
DIMENSION F(1),G(33),C(1),PHI(1)
DATA PI/3.14159265358979/
DX=2.*PI/FLOAT(NSCANS)
NE=NSCANS/2+1
CO=0.
DO 10 J=1,NSCANS
10 CO=CO+F(J)/FLOAT(NSCANS)
DO 30 N=1,NE
EN=N
S1=0.
C1=0.
X=-DX
DO 20 J=1,NSCANS
X=X+DX
S1=S1+(1./PI)*F(J)*SIN(EN*X)*DX
20 C1=C1+(1./PI)*F(J)*COS(EN*X)*DX
C(N)=SQRT(S1**2+C1**2)
IF((FLOAT(N)).EQ.(FLOAT(NSCANS)/2.)) C(N)=C(N)/2.
PHI(N)=0.0
30 IF((S1**2+C1**2).NE.0.) PHI(N)=ATAN2(C1,S1)
DO 50 N=1,NE
EN=N
X=-DX
DEV1=0.
DO 40 J=1,NSCANS
X=X+DX
IF(N.EQ.1) G(J)=CO
G(J)=G(J)+C(N)*SIN(EN*X+PHI(N))
40 IF(ABS(G(J)-F(J)).GT.DEV1) DEV1=ABS(G(J)-F(J))
IF(N.EQ.1) DEV=DEV1
IF(N.EQ.1) N2=1
IF(DEV1.LT.DEV) N2=N
IF(DEV1.LT.DEV) DEV=DEV1
50 CONTINUE
RETURN
END

```

```

SUBROUTINE YSCALE(N,F,YD,YX,YM)
DIMENSION F(800)
PP=F(1)
PN=PP
DO 10 I=1,N
IF(F(I).GT.PP) PP=F(I)
IF(F(I).LT.PN) PN=F(I)
10 CONTINUE
PP=PP*1.E+5
PN=PN*1.E+5
P=PP-PN
SN=0.
IF(PP.GT.0.) SN=1.
IF(YD.NE.0.) GO TO 50
DO 20 I=1,21
FA=(1.E-9)*(10.**I)
FA1=FA*.1
IF(P/FA.GE.1.) GO TO 20
GO TO 30

```

```

SA304690
SA304700
SA304710
SA304720
SA304730
SA304740
SA304750
SA304760
SA304770
SA304780
SA304790
SA304800
SA304810
SA304820
SA304830
SA304840
SA304850
SA304860
SA304870
SA304880
SA304890
SA304900
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SA304920
SA304930
SA304940
SA304950
SA304960
SA304970
SA304980
SA304990
SA305000
SA305010
SA305020
SA305030
SA305040
SA305050
SA305060
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SA305080
SA305090
SA305100
SA305110
SA305120
SA305130
SA305140
SA305150
SA305160
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SA305180
SA305190
SA305200
SA305210
SA305220
SA305230
SA305240
SA305250
SA305260
SA305270
SA305280

```


NOLTR 73-191

```

20 CONTINUE
30 YD=5.*FA1
   YX=FLOAT(INT(PP/YD+SN))*YD
   YM=YX-7.*YD
   CN=4.
   DO 40 I=1,3
     CN=CN*.5
     YD1=CN*FA1
     YX1=FLOAT(INT(PP/YD1+SN))*YD1
     YM1=YX1-7.*YD1
     IF((7.*YD1-P).LT.0.0.OR.YX1.LT.PP.OR.YM1.GT.PN) RETURN
     YD=YD1
     YX=YX1
     YM=YM1
40 CONTINUE
50 YX=FLOAT(INT(PP/YD+SN))*YD
   YM=YX-7.*YD
   RETURN
   END

```

SA305290
 SA305300
 SA305310
 SA305320
 SA305330
 SA305340
 SA305350
 SA305360
 SA305370
 SA305380
 SA305390
 SA305400
 SA305410
 SA305420
 SA305430
 SA305440
 SA305450
 SA305460
 SA305470

APPENDIX F
SAMPLE PROBLEMS FOR SA3024

I. PROBLEM EXECUTED ON INTERCOM (TIME-SHARING SYSTEM)

NØL INTERCØM
TYPE "LØGIN."
LØGIN(S)
024533LACK/ /4

09/07/73 14.35.39. BD/42/31
C- SETUP.FØRTRAN

ØN AT 14.36.03. 09/07/73
**FØRTRAN
**NEW ØR ØLD FILE- ATTACH(BN3024,BN3024)*ATTACH(DAT024,DAT024)

14.36.38.ATTACH(BN3024,BN3024)
14.36.42.ATTACH(DAT024,DAT024)
**READY.
REWIND(BN3024)*CØPYBR(BN3024,FIL,16)*RETURN(BN3024)*TAPE(ØN)

**READY.

FIL.

6	25	3	1	0	1
96.00	1.000	0.000			

6	25	3	0	0	0
96.00	1.000	0.000			

0									
10	0	-5043	-4755	-3815	-1233	4384	12737	17492	
12596	4313	-1257	-3821	-4763	-5027	-4998	-4765	-4365	
-3863	-3410	-3226	-3394	-3865	-4359	-4767	-4990	-5037	
-9028	-8812	-8146	-6754	-4208	-1189	272	-1197	-4267	
-6727	-8155	-8806	-9041	-8976	-8510	-7287	-4838	-1645	
72	-1604	-4806	-7289	-8504	-8971	-9037	-13025	-12877	
-12544	-12066	-11480	-10968	-10779	-10971	-11484	-12066	-12554	
-12883	-13030	-12926	-12190	-9905	-3946	5892	12194	5952	
-4064	-9930	-12196	-12910	-13042	-17035	-16916	-16423	-15194	
-12819	-9618	-8009	-9665	-12770	-15201	-16425	-16920	-17041	
-16906	-16408	-15151	-12748	-9465	-7798	-9462	-12745	-15181	
-16416	-16911	-17037	-21024	-20924	-20199	-17884	-12022	-2110	
4183	-2072	-12092	-17925	-20192	-20911	-21036	-20884	-20568	
-20056	-19474	-18985	-18775	-18979	-19477	-20063	-20545	-20886	
-21034	-25024	-24970	-24520	-23276	-20863	-17702	-15987	-17628	

Data Tape

- 20829-23299-24484-24970-25036-24817-24156-22743-20225
 - 17136-15667-17203-20207-22712-24146-24815-25035.
 0 0 0 0 0 0

Data Tape

TAFE(OFF)

\$NAM

NO = 6,
 NI = 25,
 NH = 3,
 IR = 1,
 IPI = 0,
 IW = 1,
 RI = 0.96E+02,
 CV = 0.1F+01,
 PY = 0.0,

\$END

ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT:

THE MONOFOLF MOMENT IS -- 14.339536

CURVE NO. 1

-528.4	-499.6	-405.6	-147.4	414.3	1249.6	1725.1	1235.5
407.2	-149.8	-406.2	-500.4	-526.8	-523.9	-500.6	-460.6
-410.4	-365.1	-346.7	-363.5	-410.6	-460.0	-500.8	-523.1
-527.8							

CURVE NO. 2

-527.0	-505.4	-438.8	-299.6	-45.0	256.9	403.0	256.1
-50.9	-296.9	-439.7	-504.8	-528.3	-521.8	-475.2	-352.9
-108.0	211.3	383.0	215.4	-104.8	-353.1	-474.6	-521.3
-527.9							

CURVE NO. 3

-527.0	-512.2	-478.9	-431.1	-372.5	-321.3	-302.4	-321.6
-372.9	-431.1	-479.9	-512.8	-527.5	-517.1	-443.5	-215.0
380.9	1364.7	1994.9	1370.7	369.1	-217.5	-444.1	-515.5
-528.7							

CURVE NO. 4

-527.4	-515.5	-466.2	-343.3	-105.8	214.3	375.2	209.6
-100.9	-344.0	-466.4	-515.9	-528.0	-514.5	-464.7	-339.0
-98.7	229.6	396.3	229.9	-96.4	-342.0	-465.5	-515.0
-527.6							

CURVE NO. 5

-527.0	-517.0	-444.5	-213.0	373.2	1364.4	1993.7	1368.2
366.2	-217.1	-443.8	-515.7	-528.2	-513.0	-481.4	-430.2
-372.0	-323.1	-302.1	-322.5	-372.3	-430.9	-479.1	-513.2
-528.0							

CURVE NO. 6

-526.9	-521.5	-476.5	-352.3	-110.8	205.3	376.8	212.7
-107.4	-354.4	-472.9	-521.5	-528.1	-506.2	-440.1	-298.8
-47.0	261.9	408.8	255.2	-45.2	-295.7	-439.1	-506.0
-528.0							

(DX,DY,DZ) = 825.54 211.43 20.71

(D,02,01) = 852.44 -14.37 88.61

THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD BE 852.2 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING 165.6 DEGREES FROM +X. (THE MAGNET SHOULD READ 170.4 GAMMA AT ONE METER.)

THE COMPENSATING MAGNET FOR THE Z-AXIS SHOULD BE 20.7 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING TOWARDS -Z. (THE MAGNET SHOULD READ 4.1 GAMMA AT ONE METER.)

THE A & B COEFFS. FOR THE 1TH HARMONIC ARE:

.207106285746E+02 .825542024056E+03

-.211426879295E+03

THE QUADRUPOLE MOMENTS Q11,Q22,Q33,Q12,Q13,Q23 ARE --

.913801584473E+07 .996620668908E+07 -.191042225338E+08

.141965058911E+05 .189439835558E+05 .902315837020E+04

THE A & B COEFFS. FOR THE 2TH HARMONIC ARE:

-.955211126691E+07 .109373140054E+05 -.239078103461E+06

.520952291397E+04 .819635649778E+04

THE A & B COEFFS. FOR THE 3TH HARMONIC ARE:

-.108665709727E+07 .734751632135E+08 .261862613936E+07 .23785770

1231E+10

.446316961324E+07 .306933172460E+06 .393914212069E+07

\$NAM

NO = 6,
N1 = 25,
NH = 3,
IR = 0,
IP1 = 0,
IW = 0,
R1 = 0.96E+02,
CV = 0.1E+01,
PY = 0.0,

\$END

ENTER 1 OR 0 TO INDICATE DATA PRINT OR NOT:

THE MONOPOLE MOMENT IS -- 15.233732

(DX,DY,DZ) = 827.69 -212.04 20.70
(D,02,01) = 854.67 -14.37 88.61

THE COMPENSATING MAGNET FOR THE XY-PLANE SHOULD BE 854.4 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING 165.6 DEGREES FROM +X. (THE MAGNET SHOULD READ 170.9 GAMMA AT ONE METER.)

THE COMPENSATING MAGNET FOR THE Z-AXIS SHOULD BE 20.7 GAUSS-CENTIMETER-CUBED WITH THE NORTH POLE POINTING TOWARDS -Z. (THE MAGNET SHOULD READ 4.1 GAMMA AT ONE METER.)

THE A & B COEFFS. FOR THE 1TH HARMONIC ARE:

.206988483703E+02 .827692992075E+03
-.212041720726E+03

THE QUADRUPOLE MOMENTS 011,022,033,012,013,023 ARE --

.907751557573E+07 .990809608352E+07 -.189856116592E+08
.142413793441E+05 .190259234406E+05 .904881389053E+04

THE A & B COEFFS. FOR THE 2TH HARMONIC ARE:

-.949280582962E+07 .109846220200E+05 -.239767939880E+06
.522433513554E+04 .822226419797E+04

THE A & B COEFFS. FOR THE 3TH HARMONIC ARE:

-.109329743425E+07 .736637847141E+08 .262663903488E+07 .23853972
2966E+10

.447595915074E+07 .307504565712E+06 .394909710659E+07

14.47.58.STOP

**READY.

LOGOUT.

CP TIME 2.824
PP TIME 80.985
CONNECT TIME 0 HR 14 MIN 12 SEC
TOTAL COST OF SESSION = \$ 3.72
09/07/73 LOGGED OUT AT 14.49.51.<

Notes:

1. The file BN3024 is the binary version of SA3024. It consists of 19 binary records (subprograms). The last three are plotting routines and are not used when executing problems on INTERCOM.
2. The information typed in by the user has been underlined.

APPENDIX (CONT.)
II. PROBLEM SUBMITTED TO BATCH

NØL INTERCOM
TYPE "LOGIN."
LOGIN(S)
024533LACK/ /4

09/07/73 14.50.40. BD/42/31
C- SETUP.GENERAL

ØN AT 14.51.17. 09/07/73
**GENERAL
**NEW CR ØLD FILE- NEW/IECC024*TAPE(ØN)

**READY.
1 IECC3ST,F1,T060,CM060000.55302435,024,LACKEY.
2 ATTACH(ABC,NØLBIN)
3 CØPYN(Ø,DEF,ABC)
4 RETURN(ABC)
5 ATTACH(MHL,DAT024)
6 REWIND(MHL)
7 CØPYBF(MHL,DAT024)
8 RETURN(MHL)
9 ATTACH(BN3024,BN3024)
10 LØAD(BN3024)
11 DEF.
12*WEØR
14 REWIND(ABC)
15 GØULD1,14,ABC
16*WEØR

5490	6	25	3	1	1	1			
5500	96.00	1.000	0.000						
5510	1								
5520	6	25	3	0	1	0			
5530	96.00	1.000	0.000						
5540	0								
5550	10	0	-5043	-4755	-3815	-1233	4384	12737	17492
5560	12596	4313	-1257	-3821	-4763	-5027	-4998	-4765	-4365
5570	-3863	-3410	-3226	-3394	-3865	-4359	-4767	-4990	-5037
5580	-9028	-8812	-8146	-6754	-4208	-1189	272	-1197	-4267
5590	-6727	-8155	-8806	-9041	-8976	-8510	-7287	-4838	-1645
5600	72	-1604	-4806	-7289	-8504	-8971	-9037	-13025	-12877
5610	-12544	-12066	-11480	-10968	-10779	-10971	-11464	-12066	-12554

Tape prepared
beforehand in
LOCAL mode

5620-12883-13030-12926-12190 -9905 -3946 5892 12194 5952
 5630 -4064 -9930-12196-12910-13042-17035-16916-16423-15194
 5640-12819 -9618 -8009 -9665-12770-15201-16425-16920-17041
 5650-16906-16408-15151-12748 -9465 -7798 -9462-12745-15181
 5660-16416-16911-17037-21024-20924-20199-17884-12022 -2110
 5670 4183 -2072-12092-17925-20192-20911-21036-20884-20568
 5680-20056-19474-18985-18775-18979-19477-20063-20545-20886
 5690-21034-25024-24970-24520-23278-20863-17702-15987-17628
 5700-20829-23299-24484-24970-25036-24817-24156-22743-20225
 5710-17136-15667-17203-20207-22712-24146-24815-25035
 5720 0 0 0 0 0 0
 TAPE(OFF)

Tape prepared
 beforehand in
 LOCAL mode

SAVE*PURGE(IECC024)*BATCH.*QUEUES.

**SAVED IECC024
 14.58.13.PURGE(IECC024)
 TYPE FILE NAME-IECC024

TYPE DISPOSITION-INPUT

TYPE FILE NAME-END

QUEUES 15.00.35. I= 19, O= 3, P=v' 1, C= 2.
 INPUT = 19
 CBCGU81-5 ECAAG8Y-3 HHJF18B-2 SAFCS8N-2 DCCXX8Z-3 SAFCS81-2
 BCAJH8E-2 HHJ1J74-2 HHJF18A-2 AJFG084-3 SAFHF71-2 IECC386-1
 ICCBG8X-4 IGBJE7T-2 HHJMV8M-2 HHJ1J8L-2 DBCJG8W-2 IAJFS83-5
 DCAEV8U-5
 OUTPUT= 3
 BDASH80-0 IAJF03F-0 DAYFI7G-0
 PUNCH = 1
 BDASH80-0
 COMMON= 2
 SSSSSSU-0 SSSSSST-0
 CONTROL PTS.
 AJFVD78-2 HHJFI79-5 GRID68P-5 AJFG060-4 AJFXF54-2 GRIDF80-5
 IAJF082-5
 15.00.35.STOP
 **READY.
 LOGOUT.

CP TIME .949
 PP TIME 144.517
 CONNECT TIME 0 HR 11 MIN 2 SEC
 TOTAL COST OF SESSION = \$ 3.75
 09/07/73 LOGGED OUT AT 15.01.42.<

Note:

1. The information typed in by the user has been underlined.

APPENDIX G
LISTING OF SA4024

```

PROGRAM DIPDAT(INPUT=65,OUTPUT=65,TAPE5=INPUT,TAPE6=OUTPUT,
+ DAT024=65,TAPE7=DAT024,TAPE99)
C
C          SATELLITE DATA PROGRAM
C
C  THIS PROGRAM GENERATES AND PLOTS DATA REPRESENTING THE NORMAL COM-
C  PONENT OF THE MAGNETIC FIELD FROM A SATELLITE. THE DATA IS ENTERED
C  IN THE FOLLOWING ORDER --
C
C  N0 - THE NUMBER OF GREAT CIRCLES OF DATA. (N0 IS USUALLY EVEN, E.G.,
C       N0=(N1-1)/2. THE PROGRAM STOPS IF N0=0.)
C
C  N1 - THE NUMBER OF DATA POINTS PER GREAT CIRCLE. (N1 IS ALWAYS ODD.
C       THE FIRST DATA POINT IS THE SAME AS THE LAST FOR EACH GREAT CIR-
C       LE.)
C
C  NH - THE TOTAL NUMBER OF DIFFERENT HARMONICS (DEGREES) OF MULTIPOLE
C       MAGNETS TO BE CONSIDERED.
C
C  IP - DETERMINES WHETHER OR NOT THE MAGNETIC DATA IS TO BE PLOTTED.
C       (IP=0 MEANS THAT THE DATA WILL NOT BE PLOTTED.)
C
C  R1 - THE RADIUS OF THE MEASUREMENT SPHERE IN INCHES.
C
C  EG - THE ERROR (IN GAMMAS) TO BE RANDOMLY INSERTED INTO THE DATA TO
C       REPRESENT INSTRUMENTATION INACCURACIES.
C
C  EA - THE ERROR (IN DEGREES) TO BE RANDOMLY INSERTED INTO THE DATA TO
C       REPRESENT MEASUREMENT POSITION ERRORS.
C
C  ED - THE CONSTANT ERROR (IN GAMMAS) TO BE INSERTED INTO THE DATA.
C       (THIS WILL BE ANALYZED AS MONOPOLE MOMENT.)
C
C  PY - THE SCALE FACTOR (GAMMAS/INCH) FOR THE Y-AXIS IF THE DATA IS TO
C       BE PLOTTED. (IF PY=0.0 A FACTOR WILL BE COMPUTED FROM THE DATA.)
C
C  F9 - THE FORMAT FOR READING AND PRINTING THE SPHERICAL COEFFICIENTS
C       A(I) AND B(I).
C
C  DO ** NI=1,NH
C
C  NN - THE HARMONIC NUMBER (DEGREE) OF THE MULTIPOLE DATA BEING READ
C       IN. (NN=1 FOR DIPOLES, 2 FOR QUADRUPOLES, ETC.)

```

SA400018
SA400020
SA400030
SA400040
SA400050
SA400060
SA400070
SA400080
SA400090
SA400100
SA400110
SA400120
SA400130
SA400140
SA400150
SA400160
SA400170
SA400180
SA400190
SA400200
SA400210
SA400220
SA400230
SA400240
SA400250
SA400260
SA400270
SA400280
SA400290
SA400300
SA400310
SA400320
SA400330
SA400340
SA400350
SA400360
SA400370
SA400380
SA400390
SA400400
SA400410
SA400420
SA400430
SA400440
SA400450
SA400460
SA400470
SA400480

NOLTR 73-191

```

C      NM - THE NUMBER OF MULTIPLES WITH HARMONIC NUMBER NN. SA400490
C      DO ** NJ=1,NM SA400500
C      P - THE POSITION VECTOR (IN INCHES) OF THE MULTIPOLE IN RECTANG- SA400510
C      ULAR COORDINATES. SA400520
C      A(I) - THE SPHERICAL COEFFICIENTS FOR THE MULTIPOLE OF DEGREE NN SA400530
C      WHERE I=1,2,--,NN+1. (I-1 IS THE ORDER OF THE ITH COEFF.) SA400540
C      B(I) - THE SPHERICAL COEFFICIENTS FOR THE MULTIPOLE OF DEGREE NN SA400550
C      WHERE I=1,2,--,NN. (I IS THE ORDER OF THE ITH COEFF.) SA400560
C      **NOTE** THE COEFFICIENTS FOR A DIPOLE OF MOMENT (DX,DY,DZ) ARE SA400570
C      AS FOLLOWS -- SA400580
C      A(1)=DZ, A(2)=DX, AND B(1)=DY. SA400590
C      ** CONTINUE SA400600
C      ***** SA400610
C      DIMENSION F(1000),P(3),A(11),B(10),R(3),DM(3),F1(3),F9(7) SA400620
C      +,A1(33),A2(1000) SA400630
C      NAMELIST/NAM/NO,N1,NH,IP,R1,EG,EA,ED,PY SA400640
C      DATA PI/3.14159265358979/ SA400650
C      IPT=0 SA400660
C      10 READ(5,01) NO,N1,NH,IP SA400670
C      IF(IPT.EQ.0.AND.IP.NE.0) CALL CALCM1(0,10H024 LACKEY,-10.) SA400680
C      IF(NO.EQ.0) GO TO 70 SA400690
C      IF(IP.NE.0) IPT=1 SA400700
C      READ(5,02) R1,EG,EA,ED,PY SA400710
C      READ(5,04) F9 SA400720
C      WRITE(6,NAM) SA400730
C      WRITE(6,14) F9 SA400740
C      D1=PI/FLOAT(NO) SA400750
C      D2=2.*PI/FLOAT(N1-1) SA400760
C      IZ=1 SA400770
C      DO 50 NI=1,NH SA400780
C      READ(5,01) NN,NM SA400790
C      WRITE(6,03) NN,NM SA400800
C      NN1=NN+1 SA400810
C      DO 40 NJ=1,NM SA400820
C      READ(5,02) P SA400830
C      READ(5,F9) (A(I),I=1,NN1) SA400840
C      READ(5,F9) (B(I),I=1,NN) SA400850
C      WRITE(6,02) P SA400860
C      WRITE(6,F9) (A(I),I=1,NN1) SA400870
C      WRITE(6,F9) (B(I),I=1,NN) SA400880
C      IF(NJ.EQ.1.AND.IZ.EQ.1) CALL SUM(0.,P,0.,P,DM) SA400890
C      IF(NN.NE.1) GO TO 20 SA400900
C      DM(1)=DM(1)+A(2) SA400910
C      DM(2)=DM(2)+B(1) SA400920
C      DM(3)=DM(3)+A(1) SA400930
C      20 DO 30 K=1,N0 SA400940
C      IF(IZ.EQ.1) A1(K)=FLOAT(K-1)*D1+(RANF(1.1)-.5)*2.*EA*PI/180. SA400950
C      S1=SIN(A1(K)) SA400960
C      C1=COS(A1(K)) SA400970
C      DO 30 L=1,N1 SA400980
C      KL=L+(K-1)*N1 SA400990

```

```

IF (IZ.EQ.1) A2(KL)=FLOAT(L-1)*D2+(RANF(1,1)-.5)*2.*EA*PI/180.
S2=SIN(A2(KL))
C2=COS(A2(KL))
IF (IZ.EQ.1) F(KL)=0.
R(1)=R1*C1*S2
R(2)=R1*S1*S2
R(3)=R1*C2
CALL AMPFLD(R,P,NN,A,B,F1)
30 F(KL)=F(KL)+DOT(F1,R)/SQRT(DOT(R,R))
IZ=0
40 CONTINUE
50 CONTINUE
N3=NO*N1
DO 60 I=1,N3
60 F(I)=F(I)+ED*1.E-5+(RANF(1,1)-.5)*2.E-5*EG
CALL PRNPUF(F,NO,N1)
CALL SPCOOR(DM,F1)
WRITE(6,06) DM,F1
IF (IP.NE.0) CALL DATPLT(NO,N1,F,PY)
GO TO 10
70 IF (IPT.NE.0) CALL CALCM1(0,10H024 LACKEY,+10.)
STOP
01 FORMAT(14I5)
02 FORMAT(9F8.4)
03 FORMAT(48H1 POSITION VECTOR AND COEFFICIENTS FOR NN ( NM = ,2I3,
+ 7H ARE --)
04 FORMAT(7A10)
06 FORMAT(20H1 ACTUAL MOMENT IS --/3F20.12/3F20.12)
14 FORMAT(6H0F9 = ,7A10)
END

SUBROUTINE AMPFLD(R,P,N,A,B,F)
DIMENSION R(3),P(3),A(11),B(10),F(3),U(3),V(3),W(3)
DATA PI/3.14159265358979/
CALL SUM(1.,R,-1.,P,U)
CALL SPCOOR(U,V)
R1=V(1)*2.54
IF (R1.NE.0.) GO TO 10
PRINT 01
RETURN
10 O1=V(2)*PI/180.
O2=V(3)*PI/180.
S1=SIN(O1) $ S2=SIN(O2)
C1=COS(O1) $ C2=COS(O2)
U(1)=C1*S2 $ U(2)=S1*S2 $ U(3)=C2
V(1)=C1*C2 $ V(2)=S1*C2 $ V(3)=-S2
W(1)=-S1 $ W(2)=C1 $ W(3)=0.
H1=H2=H3=0.
NN=N+1
DO 30 MM=1,NN
M=MM-1
P1=FLOAT(M)*O1
SM1=SIN(P1) $ CM1=COS(P1)
P1=SPNM(N,M,C2)
P2=SPNM(N,M+1,C2)
P3=P4=0. $ CP=SQRT(FLOAT(N*(N+1)))/2.)
IF (M.EQ.0) GO TO 20
P4=FLOAT((N-M+1)*(N-M+2))*PNM(N+1,M-1,C2)+PNM(N+1,M+1,C2)
P4=.5*SQRT(2.*ANF(N-M)/ANF(N+M))*P4
P3=C2*P4 $ CP=SQRT(FLOAT((N-M)*(N+M+1)))

```

SA401090
SA401100
SA401110
SA401120
SA401130
SA401140
SA401150
SA401160
SA401170
SA401180
SA401190
SA401200
SA401210
SA401220
SA401230
SA401240
SA401250
SA401260
SA401270
SA401280
SA401290
SA401300
SA401310
SA401320
SA401330
SA401340
SA401350
SA401360
SA401370
SA401380
SA401390
SA401400
SA401410
SA401420
SA401430
SA401440
SA401450
SA401460
SA401470
SA401480
SA401490
SA401500
SA401510
SA401520
SA401530
SA401540
SA401550
SA401560
SA401570
SA401580
SA401590
SA401600
SA401610
SA401620
SA401630
SA401640
SA401650
SA401660
SA401670
SA401680

```

20 B1=0.
   IF(M.GT.0) B1=B(M)
   A1=A(MM)*CM1+B1*SM1
   H1=H1+A1*P1
   H2=H2+A1*(CP*P2-P3)
   H3=H3+(A(MM)*SM1-B1*CM1)*P4
30 CONTINUE
   R1=R1**(N+2)
   H1=FLOAT(N+1)*H1/R1 $ H2=H2/R1 $ H3=H3/R1
   CALL SUM(H1,U,H2,V,F) $ CALL SUM(1.,F,H3,W,F)
   RETURN
01 FORMAT(50HOFIELD VECTOR CANNOT BE COMPUTED AT POLE POSITION.//)
   END

FUNCTION SPNM(N,M,X)
  SPNM=PNM(N,M,X)
  IF(M.EQ.0.OR.M.GT.N) RETURN
  SPNM=SQRT(2.*ANF(N-M)/ANF(N+M))*SPNM
  RETURN
END

FUNCTION PNM(N,M,X)
  PNM=0.
  IF(M.GT.N) RETURN
  IF((ABS(1.-ABS(X))).GT.1.E-9) GO TO 20
  IF(M.EQ.0) RETURN
  PNM=-1.
  IF(X.GT.0..OR.N.EQ.INT(FLOAT(N)/2.+1)*2) PNM=1.
  RETURN
20 CNM=(2.**N)*ANF(N)*ANF(N-M)
  PNM=1.
  IF(M.EQ.0) PNM=SQRT(1.-X*X)**M
  CNM=ANF(2*N)*PNM/CNM
  PNM=1.
  IF(M.EQ.N) PNM=X**(N-M)
  IF(N-M.LE.1) GO TO 40
  PRD1=1.
  NT=(N-M)/2
  DO 30 I=1,NT
    AN1=FLOAT(N-M-2*I+2)
    AN2=FLOAT(2*I)
    AN3=FLOAT(2*N-2*I+1)
    PRD1=-PRD1*AN1*(AN1-1)/(AN2*AN3)
    NE=N-M-2*I
    AN1=1.
    IF(NE.GT.0) AN1=X**NE
    PNM=PNM+PRD1*AN1
30 CONTINUE
40 PNM=CNM*PNM
   RETURN
   END

FUNCTION ANF(N)
  DOUBLE PRECISION AN
  ANF=1.
  AN=1.D0
  IF(N.LT.0) PRINT 01
  IF(N.LT.2) RETURN
  DO 10 I=2,N
10 AN=AN*DBLE(FLOAT(I))

```

SA401690
 SA401700
 SA401710
 SA401720
 SA401730
 SA401740
 SA401750
 SA401760
 SA401770
 SA401780
 SA401790
 SA401800
 SA401810
 SA401820
 SA401830
 SA401840
 SA401850
 SA401860
 SA401870
 SA401880
 SA401890
 SA401900
 SA401910
 SA401920
 SA401930
 SA401940
 SA401950
 SA401960
 SA401970
 SA401980
 SA401990
 SA402000
 SA402010
 SA402020
 SA402030
 SA402040
 SA402050
 SA402060
 SA402070
 SA402080
 SA402090
 SA402100
 SA402110
 SA402120
 SA402130
 SA402140
 SA402150
 SA402160
 SA402170
 SA402180
 SA402190
 SA402200
 SA402210
 SA402220
 SA402230
 SA402240
 SA402250
 SA402260
 SA402270
 SA402280

ANF=SNGL(AN)	SA402290
RETURN	SA402300
01 FORMAT(37H1FACTORIAL INTEGER IS LESS THAN ZERO.//)	SA402310
END	SA402320
	SA402330
SUBROUTINE SPCOOR(D,R)	SA402340
DIMENSION D(3),R(3)	SA402350
DATA PI/3.14159265358979/	SA402360
R(1)=SQRT(DOT(D,D))	SA402370
R(2)=0.	SA402380
R(3)=0.	SA402390
IF(R(1).EQ.0.) RETURN	SA402400
IF(D(1)**2+D(2)**2.NE.0.) R(2)=ATAN2(D(2),D(1))*180./PI	SA402410
R(3)=ATAN2(SQRT(D(1)**2+D(2)**2),D(3))*180./PI	SA402420
RETURN	SA402430
END	SA402440
	SA402450
SUBROUTINE SUM(A,X,B,Y,Z)	SA402460
DIMENSION X(3),Y(3),Z(3)	SA402470
DO 10 I=1,3	SA402480
10 Z(I)=A*X(I)+B*Y(I)	SA402490
RETURN	SA402500
END	SA402510
	SA402520
FUNCTION DOT(X,Y)	SA402530
DIMENSION X(3),Y(3)	SA402540
DOT=X(1)*Y(1)+X(2)*Y(2)+X(3)*Y(3)	SA402550
RETURN	SA402560
END	SA402570
	SA402580
SUBROUTINE PRNPUF(F,N0,N1)	SA402590
DIMENSION F(800),F1(33),IF(800)	SA402600
N3=N0*N1	SA402610
ICS=10	SA402620
ICZ=0	SA402630
II=0	SA402640
DO 20 K0=1,N0	SA402650
KS=1+(K0-1)*N1	SA402660
KE=KS+N1-1	SA402670
DO 10 I=1,N1	SA402680
II=II+1	SA402690
F1(I)=F(II)*1.E5	SA402700
10 IF(II)=INT(F(II)*1.E+6)	SA402710
WRITE(6,01) K0	SA402720
20 WRITE(6,02) (F1(I),I=1,N1)	SA402730
WRITE(7,03) ICS,ICZ,(IF(I),I=1,N3)	SA402740
01 FORMAT(11HOCURVE NO. ,I2)	SA402750
02 FORMAT(1H ,8F8.1)	SA402760
03 FORMAT(9I6)	SA402770
RETURN	SA402780
END	SA402790
	SA402800
SUBROUTINE DATPLT(N0,N1,F,YDIST)	SA402810
DIMENSION F(1000),F1(33),F2(73),X(33),X1(73),C(1),P(1),CH(16)	SA402820
DATA PI/3.14159265358979/	SA402830
DATA CH/1H1,1H2,1H3,1H4,1H5,1H6,1H7,1H8,1H9,2H10,2H11,2H12,2H13,	SA402840
+ 2H14,2H15,2H16/	SA402850
N3=N0*N1	SA402860
CALL YSCALE(N3,F,YDIST,YMAX,YMIN)	SA402870
DX=360./FLOAT(N1-1)	SA402880

```

F1(1)=F1(2)=YMAX $ X(1)=0. $ X(2)=360.
CALL CALCM1(2,X,F1,0,0.,360.,YMIN,YMAX,6.,7.,TITLE,0,
+ 20HCOLATITUDE (DEGREES),-20,30HNORMAL MAGNETIC FIELD (GAMMAS),
+ 30,0.,18)
F1(2)=YMIN $ X(1)=X(2)=360.
CALL CALCM1(-2,X,F1,0)
F1(1)=F1(2)=0. $ X(1)=360. $ X(2)=0.
IF(YMAX.GT.0.0.AND.YMIN.LT.0.0) CALL CALCM1(-2,X,F1,0)
DO 10 I=1,N1
10 X(I)=DX*FLOAT(I-1)
DO 20 I=1,73
20 X1(I)=5.*FLOAT(I-1)
IJ=0
DO 50 I=1,N0
DO 30 J=1,N1
IJ=IJ+1
30 F1(J)=F(IJ)*1.E+5
CALL FNCTON(F1,N1-1,C,CO,P,N2,DEV)
DO 40 N=1,N2
EN=N
DO 40 J=1,73
XX=X1(J)*PI/180.
IF(N.EQ.1) F2(J)=C0
40 F2(J)=F2(J)+C(N)*SIN(EN*XX+P(N))
CALL CALCM1(-73,X1,F2,0)
DO 50 J=1,N1
NC=1+I/10
CALL SYMBL4(X(J)/60.,(F1(J)-YMIN)/YDIST,.08,CH(I),0.,NC)
50 CONTINUE
RETURN
END

SUBROUTINE FNCTON(F,NSCANS,C,CO,PHI,N2,DEV)
DIMENSION F(1),G(33),C(1),PHI(1)
DATA PI/3.14159265358979/
DX=2.*PI/FLOAT(NSCANS)
NE=NSCANS/2+1
C0=0.
DO 10 J=1,NSCANS
10 C0=C0+F(J)/FLOAT(NSCANS)
DO 30 N=1,NE
EN=N
S1=0.
C1=0.
X=-DX
DO 20 J=1,NSCANS
X=X+DX
S1=S1+(1./PI)*F(J)*SIN(EN*X)*DX
20 C1=C1+(1./PI)*F(J)*COS(EN*X)*DX
C(N)=SQRT(S1**2+C1**2)
IF((FLOAT(N)).EQ.(FLOAT(NSCANS)/2.)) C(N)=C(N)/2.
PHI(N)=0.0
30 IF((S1**2+C1**2).NE.0.) PHI(N)=ATAN2(C1,S1)
DO 50 N=1,NE
EN=N
X=-DX
DEV1=0.
DO 40 J=1,NSCANS
X=X+DX
IF(N.EQ.1) G(J)=C0

```

SA402890
SA402900
SA402910
SA402920
SA402930
SA402940
SA402950
SA402960
SA402970
SA402980
SA402990
SA403000
SA403010
SA403020
SA403030
SA403040
SA403050
SA403060
SA403070
SA403080
SA403090
SA403100
SA403110
SA403120
SA403130
SA403140
SA403150
SA403160
SA403170
SA403180
SA403190
SA403200
SA403210
SA403220
SA403230
SA403240
SA403250
SA403260
SA403270
SA403280
SA403290
SA403300
SA403310
SA403320
SA403330
SA403340
SA403350
SA403360
SA403370
SA403380
SA403390
SA403400
SA403410
SA403420
SA403430
SA403440
SA403450
SA403460
SA403470
SA403480

G(J)=G(J)+C(N)*SIN(EN*X+PHI(N))	SA403490
40 IF(ABS(G(J)-F(J)).GT.DEV1) DEV1=ABS(G(J)-F(J))	SA403500
IF(N.EQ.1) DEV=DEV1	SA403510
IF(N.EQ.1) N2=1	SA403520
IF(DEV1.LT.DEV) N2=N	SA403530
IF(DEV1.LT.DEV) DEV=DEV1	SA403540
50 CONTINUE	SA403550
RETURN	SA403560
END	SA403570
	SA403580
SUBROUTINE YSCALE(N,F,YD,YX,YM)	SA403590
DIMENSION F(800)	SA403600
PP=F(1)	SA403610
PN=PP	SA403620
DO 10 I=1,N	SA403630
IF(F(I).GT.PP) PP=F(I)	SA403640
IF(F(I).LT.PN) PN=F(I)	SA403650
10 CONTINUE	SA403660
PP=PP*1.E+5	SA403670
PN=PN*1.E+5	SA403680
P=PP-PN	SA403690
SN=0.	SA403700
IF(PP.GT.0.) SN=1.	SA403710
IF(YD.NE.0.) GO TO 50	SA403720
DO 20 I=1,21	SA403730
FA=(1.E-9)*(10.**I)	SA403740
FA1=FA*.1	SA403750
IF(P/FA.GE.1.) GO TO 20	SA403760
GO TO 30	SA403770
20 CONTINUE	SA403780
30 YD=5.*FA1	SA403790
YX=FLOAT(INT(PP/YD+SN))*YD	SA403800
YM=YX-7.*YD	SA403810
CN=4.	SA403820
DO 40 I=1,3	SA403830
CN=CN*.5	SA403840
YD1=CN*FA1	SA403850
YX1=FLOAT(INT(PP/YD1+SN))*YD1	SA403860
YM1=YX1-7.*YD1	SA403870
IF((7.*YD1-P).LT.0.0.OR.YX1.LT.PP.OR.YM1.GT.PN) RETURN	SA403880
YD=YD1	SA403890
YX=YX1	SA403900
YM=YM1	SA403910
40 CONTINUE	SA403920
50 YX=FLOAT(INT(PP/YD+SN))*YD	SA403930
YM=YX-7.*YD	SA403940
RETURN	SA403950
END	SA403960

APPENDIX H
 SAMPLE PROBLEMS FOR SA4024

I. PROBLEM EXECUTED ON INTERCOM (TIME-SHARING SYSTEM)

NOL INTERCOM
 TYPE "LOGIN."
LOGIN.

TYPE USERNAME/PASSWORD/TTY NO. (OUTSIDE NOL TTY=88)
024533LACK/ /4

09/10/73 15.12.25. BD/42/35
 C- SETUP.FORTRAN

ON AT 15.12.38. 09/10/73
 **FORTRAN
 **NEW OR OLD FILE- ATTACH(BN4024,BN4024)

15.12.56.ATTACH(BN4024,BN4024)
 **READY.
REVIND(BN4024)*COPYBR(BN4024,FIL,9)*RETURN(BN4024)

**READY.

FIL.

6	25	1	0
96.00	1.000	0.200	25.00
(1H,7E10.4)			

SNAM

NO = 6.

NI = 25,
 NH = 1,
 IP = 0,
 RI = 0.96E+02,
 EG = 0.1E+01,
 FA = 0.2E+00,
 FD = 0.25E+02,
 FY = -0.0,

SEND

F9 = (1H, 7E10.4)
1 3

POSITION VECTOR AND COEFFICIENTS FOR NN & NM = 1 3 ARE --

39.00 0.000 0.000
+.0000E+00+.3100E+05
+.0000E+00

39.0000 0.0000 0.0000
 0. .3100E+05
 0.
-21.00 36.373 0.000
+.0000E+00-.1500E+05
+.2598E+05

21.0000 36.3730 0.0000
 0. -.1500E+05
 .2598E+05
-21.00 -36.373 0.000
+.0000E+00-.1500E+05
-.2598E+05

21.0000-36.3730 0.0000
 0. - .1500E+05
 - .2598E+05

CURVE NO. 1
 -504.3 -475.6 -381.5 -123.3 438.5 1273.8 1749.2 1259.7
 431.3 -125.8 -382.2 -476.4 -502.8 -499.8 -476.5 -436.6
 -386.3 -341.0 -322.6 -339.4 -386.6 -436.0 -476.8 -499.1
 -503.8

CURVE NO. 2
 -502.8 -481.3 -414.6 -275.5 -20.8 281.0 427.2 280.3
 -26.8 -272.8 -415.6 -480.7 -504.2 -497.6 -451.0 -328.7
 -83.8 235.4 407.3 239.5 -80.6 -329.0 -450.5 -497.1
 -503.8

CURVE NO. 3
 -502.5 -487.8 -454.5 -406.6 -348.0 -296.8 -278.0 -297.1
 -348.4 -406.6 -455.4 -488.3 -503.1 -492.7 -419.0 -190.5
 405.3 1389.2 2019.5 1395.2 393.6 -193.1 -419.7 -491.0
 -504.3

CURVE NO. 4
 -503.5 -491.6 -442.4 -319.4 -81.9 238.1 399.1 233.5
 -77.1 -320.2 -442.6 -492.0 -504.1 -490.6 -440.9 -315.1
 -74.8 253.5 420.1 253.7 -74.5 -318.2 -441.7 -491.1
 -503.7

CURVE NO. 5
 -502.5 -492.5 -420.0 -188.4 397.7 1389.0 2018.3 1392.8
 390.7 -192.5 -419.2 -491.1 -503.6 -488.4 -456.9 -405.6
 -347.5 -298.5 -277.5 -298.0 -347.8 -406.3 -454.5 -488.6
 -503.4

CURVE NO. 6
 -502.4 -497.0 -452.0 -327.9 -86.4 229.7 401.3 237.1
 -82.9 -329.9 -448.4 -497.0 -503.7 -481.8 -415.6 -274.3
 -22.6 286.3 433.2 279.6 -20.7 -271.3 -414.7 -481.6
 -503.6

ACTUAL MOMENT IS --
 1000.000000000000 0.000000000000 0.000000000000
 1000.000000000000 0.000000000000 90.000000000000

0 0 0 0

15.20.24.STOP

**READY.

LOGOUT.

CP TIME 4.233

PP TIME 56.828

CONNECT TIME 0 HR 8 MIN 11 SEC

TOTAL COST OF SESSION = \$ 2.60

09/10/73 LOGGED OUT AT 15.20.36.<

Notes:

1. The file BN4024 is the binary version of SA4024. It consists of 12 binary records (subprograms). The last three are plotting routines and are not used when executing problems on INTEROOM.
2. The information typed in by the user has been underlined.

APPENDIX H (CONT.)
II. LISTING OF DAT024

10	0	-5043	-4755	-3815	-1233	4384	12737	17492
12596	4313	-1257	-3821	-4763	-5027	-4998	-4765	-4365
-3863	-3410	-3226	-3394	-3865	-4359	-4767	-4990	-5037
-5028	-4812	-4146	-2754	-208	2810	4272	2802	-267
-2727	-4155	-4806	-5041	-4976	-4510	-3287	-838	2354
4072	2395	-806	-3289	-4504	-4971	-5037	-5025	-4877
-4544	-4066	-3480	-2968	-2779	-2971	-3484	-4066	-4554
-4883	-5030	-4926	-4190	-1905	4053	13892	20194	13952
3935	-1930	-4196	-4910	-5042	-5035	-4916	-4423	-3194
-819	2381	3990	2334	-770	-3201	-4425	-4920	-5041
-4906	-4408	-3151	-748	2534	4201	2537	-745	-3181
-4416	-4911	-5037	-5024	-4924	-4199	-1884	3977	13889
20183	13927	3907	-1925	-4192	-4911	-5036	-4884	-4568
-4056	-3474	-2985	-2775	-2979	-3477	-4063	-4545	-4886
-5034	-5024	-4970	-4520	-3278	-863	2297	4012	2371
-829	-3299	-4484	-4970	-5036	-4817	-4156	-2743	-225
2863	4332	2796	-207	-2712	-4146	-4815	-5035	

Notes:

1. The data is stored in DAT024 in a way that allows the file to be read in a (9F6.1) format.
2. The first two data points on the first line are values representing the calibration and zero readings, CS and CZ. The value of the calibration signal CV should be set to 1.0 when using data generated from SA4024.

APPENDIX H (CONT.)
 III. PROBLEM SUBMITTED TO BATCH

NOL INTERCOM
 TYPE "LOGIN."
LOGIN(S)
024533LACK/ /4

09/11/73 09.43.58. BC/42/34
 C- SETUP.GENERAL

ON AT 09.44.23. 09/11/73
 **GENERAL
 **NEW OR OLD FILE- NEW/IECC024*TAPE(ON)
 **READY.

1 IECC4ST,P1,T200,CM060000.55302435,024,LACKEY.
 2 ATTACH(ABC,NOLBIN)
 3 COPYN(0,DEF,ABC)
 4 RETURN(ABC)
 5 ATTACH(BN4024,BN4024)
 6 RFWIND(BN4024)
 7 COPYBF(BN4024,CBA)
 8 RETURN(BN4024)
 9 LOAD(CBA)
 10 DEF.
 11 *WFØR
 12 RFWIND(ABC)
 13 GOULD1,14,ABC
 14 *WFØR
 6100 6 25 1 1
 6110 96.00 1.000 0.200 25.00 0.000
 6120 (1H,7E10.4)
 6130 1 3
 6140 39.00 0.000 0.000
 6150 +.0000E+00+.3100E+05
 6160 +.0000E+00
 6170 -21.00 36.373 0.000
 6180 +.0000E+00-.1500E+05

Tape prepared
 beforehand in
 LOCAL mode

6190	+.2598E+05					
6200	-21.00	-36.373	0.000			
6210	+.0000E+00-	.1500E+05				
6220	-.2598E+05					
6230	6	25	1	1		
6240	96.00	0.000	0.000	0.000	0.000	
6250	(1H ,7E10.4)					
6260	1	1				
6270	0.000	0.000	0.000			
6280	+.0000E+00+	.1000E+04				
6290	+.0000E+00					
6300	6	25	1	1		
6310	96.00	0.000	0.000	0.000	0.000	
6320	(1H ,7E10.4)					
6330	2	1				
6340	0.000	0.000	0.000			
6350	-.9472E+07+	.0000E+00-	.2241E+06			
6360	+.0000E+00+	.0000E+00				
6370	6	25	1	1		
6380	96.00	0.000	0.000	0.000	0.000	
6390	(1H ,7E10.4)					
6400	3	1				
6410	0.000	0.000	0.000			
6420	+.0000E+00+	.6837E+08+	.0000E+00+	.2341E+10		
6430	+.0000E+00+	.0000E+00+	.0000E+00			
6440	6	25	1	1		
6450	96.00	0.000	0.000	0.000	0.000	
6460	(1H ,7E10.4)					
6470	4	1				
6480	0.000	0.000	0.000			
6490	+.1545E+12+	.0000E+00+	.1406E+11+	.0000E+00-	.1860E+11	
6500	+.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00		
6510	6	25	1	1		
6520	96.00	0.000	0.000	0.000	0.000	
6530	(1H ,7E10.4)					
6540	5	1				
6550	0.000	0.000	0.000			
6560	+.0000E+00-	.2180E+13+	.0000E+00-	.2812E+14+	.0000E+00-	.3158E+13
6570	+.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00	
6580	6	25	1	1		
6590	96.00	0.000	0.000	0.000	0.000	
6600	(1H ,7E10.4)					
6610	6	1				
6620	0.000	0.000	0.000			
6630	-.2109E+16+	.0000E+00-	.3228E+15+	.0000E+00+	.3536E+15+	.0000E+00+.4533E+16
6640	+.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00+	.0000E+00
6650	6	25	1	1		
6660	96.00	0.000	0.000	0.000	0.000	
6670	(1H ,7E10.4)					
6680	7	1				
6690	0.000	0.000	0.000			
6700	+.0000E+00+	.4320E+17+	.0000E+00+	.3541E+18+	.0000E+00+	.4962E+17+.0000E+00

Tape prepared
beforehand in
LOCAL Mode

```

6710 -.6764E+17
6720 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E
+00
6730      6      25      1      1
6740      96.00      0.000      0.000      0.000      0.000
6750      (1H ,7E10.4)
6760      8      1
6770      0.000      0.000      0.000
6780 +.2699E+20+.0000E+00+.5698E+19+.0000E+00-.5975E+19+.0000E+00-.4517E
+20
6790 +.0000E+00-.9102E+19
6800 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E
+00
6810 +.0000E+00
6820      6      25      8      1
6830      96.00      0.000      0.000      0.000      0.000
6840      (1H ,7F10.4)
6850      1      1
6860      0.000      0.000      0.000
6870 +.0000E+00+.1000E+04
6880 +.0000E+00
6890      2      1
6900      0.000      0.000      0.000
6910 -.9472E+07+.0000E+00-.2241E+06
6920 +.0000E+00+.0000E+00
6930      3      1
6940      0.000      0.000      0.000
6950 +.0000E+00+.6837E+08+.0000E+00+.2341E+10
6960 +.0000E+00+.0000E+00+.0000E+00
6970      4      1
6980      0.000      0.000      0.000
6990 +.1545E+12+.0000E+00+.1406E+11+.0000E+00-.1860E+11
7000 +.0000E+00+.0000E+00+.0000E+00+.0000E+00
7010      5      1
7020      0.000      0.000      0.000
7030 +.0000E+00-.2180E+13+.0000E+00-.2812E+14+.0000E+00-.3158E+13
7040 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00
7050      6      1
7060      0.000      0.000      0.000
7070 -.2109E+16+.0000E+00-.3228E+15+.0000E+00+.3536E+15+.0000E+00+.4533E
+16
7080 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00
7090      7      1
7100      0.000      0.000      0.000
7110 +.0000E+00+.4320E+17+.0000E+00+.3541E+18+.0000E+00+.4962E+17+.0000E
+00
7120 -.6764E+17
7130 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E
+00
7140      8      1
7150      0.000      0.000      0.000
7160 +.2699E+20+.0000E+00+.5698E+19+.0000E+00-.5975E+19+.0000E+00-.4517E
+20

```

Tape prepared
beforehand in
LOCAL mode

7170 +.0000E+00-.9102E+19
 7180 +.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E+00+.0000E
 +00
 7190 +.0000E+00
 7200 0 0 0 0
 TAPE(0FF)*SAVE*PURGE(IECC024)*BATCH.*QUEUES.

Tape prepared
 beforehand in
 LOCAL mode

**SAVED IECC024
 09.56.18.PURGE(IECC024)
 TYPE FILE NAME-IECC024

TYPE DISPOSITION-INPUT

TYPE FILE NAME-END

QUEUES 09.57.13. I=18, O= 2, P= 0, C= 4.
 INPUT = 18
 IECC43L-1 CABRE3J-4 GFJLK3K-5 GRID73I-5 GRIDN3H-5 CABRE3A-4
 CCBCB3F-5 CABRE3E-4 CABRE3G-4 CABRE3D-4 EAADW3B-5 HHJC027-3
 CABRE25-4 IECC426-1 CABAP3C-5 HHJ1J22-2 CABRE28-4 CESE002-3
 OUTPUT= 2
 AJFLZ1Z-0 HHJDP19-0
 PUNCH = NONE
 COMMON= 4
 GEOPP -0 KEEP7 -0 SSSSSSU-0 SSSSSST-0
 CONTROL PTS.
 DAJRI24-5 DBA1820-1
 09.57.13.STOP
 **READY.
LOGOUT.

CP TIME 1.674
 PP TIME 82.029
 CONNECT TIME 0 HR 14 MIN 9 SEC
 TOTAL COST OF SESSION = \$ 3.58
 09/11/73 LOGGED OUT AT 09.58.07.<

Note:

1. The information typed in by the user has been underlined.

APPENDIX I

LISTING OF SA5024

```

PROGRAM DIPSTD(INPUT=65,OUTPUT=65,TAPE5=INPUT,TAPE6=OUTPUT,TAPE99)  SA500018
                                                                    SA500020
                                                                    SA500030
C          SPHERICAL HARMONIC PROGRAM                                SA500040
                                                                    SA500050
                                                                    SA500060
C  THIS PROGRAM GENERATES AND ANALYZES DATA REPRESENTING THE NORMAL  SA500070
C  COMPONENT OF THE MAGNETIC FIELD FROM A SATELLITE. THE DATA IS ENTERED SA500080
C  IN THE FOLLOWING ORDER --                                         SA500090
                                                                    SA500100
C  N0 - THE NUMBER OF GREAT CIRCLES OF DATA. (N0 IS USUALLY EVEN, E.G., SA500110
C       N0=(N1-1)/2. THE PROGRAM STOPS IF N0=0.)                     SA500120
                                                                    SA500130
C  N1 - THE NUMBER OF DATA POINTS PER GREAT CIRCLE. (N1 IS ALWAYS ODD. SA500140
C       THE FIRST DATA POINT IS THE SAME AS THE LAST FOR EACH GREAT CIR- SA500150
C       LE.)                                                         SA500160
                                                                    SA500170
C  NH1 - THE TOTAL NUMBER OF DIFFERENT HARMONICS (DEGREES) OF MULTIPOLE SA500180
C        MAGNETS TO BE CONSIDERED.                                   SA500190
                                                                    SA500200
C  NH2 - THE HARMONIC NUMBER (DEGREE) REPRESENTING THE HIGHEST DEGREE SA500210
C        SPHERICAL HARMONIC TERM TO BE COMPUTED FROM THE DATA. (NH2=1 SA500220
C        FOR DIPOLES, 2 FOR QUADRUPOLES, ETC.)                     SA500230
                                                                    SA500240
C  IP1 - DETERMINES WHETHER OR NOT THE MAGNETIC DATA IS TO BE PLOTTED. SA500250
C        (IP1=0 MEANS THAT THE DATA WILL NOT BE PLOTTED.)         SA500260
                                                                    SA500270
C  IP2 - DETERMINES WHETHER OR NOT THE MAGNETIC DATA IS TO BE PRINTED. SA500280
C        (IP2=0 MEANS THAT THE DATA WILL NOT BE PRINTED.)         SA500290
                                                                    SA500300
C  IW - DETERMINES THE TYPE OF INTEGRATING SCHEME TO BE USED. (IW=0 SA500310
C        MEANS THAT THE EXACT, ALGEBRAIC SCHEME IS TO BE USED.)     SA500320
                                                                    SA500330
C  R1 - THE RADIUS OF THE MEASUREMENT SPHERE IN INCHES.             SA500340
                                                                    SA500350
C  EG - THE ERROR (IN GAMMAS) TO BE RANDOMLY INSERTED INTO THE DATA TO SA500360
C        REPRESENT INSTRUMENTATION INACCURACIES.                   SA500370
                                                                    SA500380
C  EA - THE ERROR (IN DEGREES) TO BE RANDOMLY INSERTED INTO THE DATA TO SA500390
C        REPRESENT MEASUREMENT POSITION ERRORS.                     SA500400
                                                                    SA500410
C  ED - THE CONSTANT ERROR (IN GAMMAS) TO BE INSERTED INTO THE DATA. SA500420
C        (THIS WILL BE ANALYZED AS MONOPOLE MOMENT.)                SA500430
                                                                    SA500440
C  PY - THE SCALE FACTOR (GAMMAS/INCH) FOR THE Y-AXIS IF THE DATA IS TO SA500450
C        BE PLOTTED. (IF PY=0.0 A FACTOR WILL BE COMPUTED FROM THE DATA.) SA500460
                                                                    SA500470
C  F9 - THE FORMAT FOR READING AND PRINTING THE SPHERICAL COEFFICIENTS SA500480

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NOLTR 73-191

```

C      A(I) AND B(I).
C
C      DO ** NI=1,NH1
C      NN - THE HARMONIC NUMBER (DEGREE) OF THE MULTIPOLE DATA BEING READ
C            IN. (NN=1 FOR DIPOLES, 2 FOR QUADRUPOLES, ETC.)
C      NM - THE NUMBER OF MULTIPOLES WITH HARMONIC NUMBER NN.
C
C      DO ** NJ=1,NM
C      P - THE POSITION VECTOR (IN INCHES) OF THE MULTIPOLE IN RECTANG-
C            ULAR COORDINATES.
C      A(I) - THE SPHERICAL COEFFICIENTS FOR THE MULTIPOLE OF DEGREE NN
C            WHERE I=1,2,--,NN+1. (I-1 IS THE ORDER OF THE ITH COEFF.)
C      B(I) - THE SPHERICAL COEFFICIENTS FOR THE MULTIPOLE OF DEGREE NN
C            WHERE I=1,2,--,NN (I IS THE ORDER OF THE ITH COEFF.)
C      **NOTE** THE COEFFICIENTS FOR A DIPOLE OF MOMENT (DX,DY,DZ) ARE
C            AS FOLLOWS --
C                      A(1)=DZ, A(2)=DX, AND B(1)=DY.
C
C      ** CONTINUE
C
C      *****
C      DIMENSION F(1000),P(3),PV(200),A(11),B(10),R(3),DM(3),F1(3),D(3)
C      +,A1(33),A2(1000),F9(7)
C      NAMELIST/NAM/NO,N1,NH1,NH2,IP1,IP2,IW,R1,EG,EA,ED,PY
C      DATA PI/3.14159265358979/
C      IPT=0
10 READ(5,01) NO,N1,NH1,NH2,IP1,IP2,IW
   IF(IPT.EQ.0.AND.IP1.NE.0) CALL CALCM1(0,10H024 LACKEY,-10.)
   IF(NO.EQ.0) GO TO 110
   IF(IP1.NE.0) IPT=1
   READ(5,02) R1,EG,EA,ED,PY
   READ(5,04) F9
   WRITE(6,NAM)
   WRITE(6,14) F9
   D1=PI/FLOAT(NO)
   D2=2.*PI/FLOAT(N1-1)
   IZ=1
   DO 50 NI=1,NH1
     READ(5,01) NN,NM
     WRITE(6,03) NN,NM
     NN1=NN+1
     DO 40 NJ=1,NM
       READ(5,02) P
       READ(5,F9) (A(I),I=1,NN1)
       READ(5,F9) (B(I),I=1,NN)
       WRITE(6,02) P
       WRITE(6,F9) (A(I),I=1,NN1)
       WRITE(6,F9) (B(I),I=1,NN)
       IF(NJ.EQ.1.AND.IZ.EQ.1) CALL SUM(0.,P,0.,P,DM)
       IF(NN.NE.1) GO TO 20
       DM(1)=DM(1)+A(2)

```

SA500490
SA500500
SA500510
SA500520
SA500530
SA500540
SA500550
SA500560
SA500570
SA500580
SA500590
SA500600
SA500610
SA500620
SA500630
SA500640
SA500650
SA500660
SA500670
SA500680
SA500690
SA500700
SA500710
SA500720
SA500730
SA500740
SA500750
SA500760
SA500770
SA500780
SA500790
SA500800
SA500810
SA500820
SA500830
SA500840
SA500850
SA500860
SA500870
SA500880
SA500890
SA500900
SA500910
SA500920
SA500930
SA500940
SA500950
SA500960
SA500970
SA500980
SA500990
SA501000
SA501010
SA501020
SA501030
SA501040
SA501050
SA501060
SA501070
SA501080

NOLTR 73-191

DM(2)=DM(2)+B(1)	SA501090
DM(3)=DM(3)+A(1)	SA501100
20 DO 30 K=1,N0	SA501110
IF(IZ.EQ.1) A1(K)=FLOAT(K-1)*D1+(RANF(1,1)-.5)*2.*EA*PI/180.	SA501120
S1=SIN(A1(K))	SA501130
C1=COS(A1(K))	SA501140
DO 30 L=1,N1	SA501150
KL=L+(K-1)*N1	SA501160
IF(IZ.EQ.1) A2(KL)=FLOAT(L-1)*D2+(RANF(1,1)-.5)*2.*EA*PI/180.	SA501170
S2=SIN(A2(KL))	SA501180
C2=COS(A2(KL))	SA501190
IF(IZ.EQ.1) F(KL)=0.	SA501200
R(1)=R1*C1*S2	SA501210
R(2)=R1*S1*S2	SA501220
R(3)=R1*C2	SA501230
CALL AMPFLD(R,P,NN,A,B,F1)	SA501240
30 F(KL)=F(KL)+DOT(F1,R)/SQRT(DOT(R,R))	SA501250
IZ=0	SA501260
40 CONTINUE	SA501270
50 CONTINUE	SA501280
N3=N0*N1	SA501290
DO 60 I=1,N3	SA501300
60 F(I)=F(I)+ED*1.E-5+(RANF(1,1)-.5)*2.E-5*EG	SA501310
IF(IW.EQ.0) CALL WGT1(N0,N1,PV)	SA501320
IF(IW.NE.0) CALL WGT3(N0,N1,PV)	SA501330
CALL AMPMNT(N0,N1,0,R1,F,PV,A,B)	SA501340
WRITE(6,05) A(1)	SA501350
IF(IP2.NE.0) CALL PRINTF(F,N0,N1)	SA501360
IF(IP1.NE.0) CALL DATPLT(N0,N1,F,PY)	SA501370
DO 100 NI=1,NH2	SA501380
CALL AMPMNT(N0,N1,NI,R1,F,PV,A,B)	SA501390
IF(NI.GT.1) GO TO 80	SA501400
D(1)=A(2)	SA501410
D(2)=B(1)	SA501420
D(3)=A(1)	SA501430
CALL SPCOOR(D,R)	SA501440
CALL SPCOOR(DM,F1)	SA501450
WRITE(6,06) DM,F1	SA501460
WRITE(6,07) D,R	SA501470
IF(F1(1).EQ.0.) WRITE(6,08)	SA501480
IF(F1(1).EQ.0.) GO TO 90	SA501490
DO 70 I=1,3	SA501500
70 R(I)=(DM(I)-D(I))*100./F1(1)	SA501510
WRITE(6,09) R	SA501520
GO TO 90	SA501530
80 IF(NI.GT.2) GO TO 90	SA501540
S3=SQRT(3.)	SA501550
Q11=S3*A(3)-A(1)	SA501560
Q22=-S3*A(3)-A(1)	SA501570
Q33=-Q11-Q22	SA501580
Q12=S3*B(2)	SA501590
Q13=S3*A(2)	SA501600
Q23=S3*B(1)	SA501610
WRITE(6,11) Q11,Q22,Q33,Q12,Q13,Q23	SA501620
90 WRITE(6,12) NI	SA501630
NI1=NI+1	SA501640
WRITE(6,13) (A(I),I=1,NI1)	SA501650
WRITE(6,13) (B(I),I=1,NI)	SA501660
100 CONTINUE	SA501670
GO TO 10	SA501680

NOLTR 73-191

```

110 IF(IPT.NE.0) CALL CALCM1(0,10H024 LACKEY,+10.)
      STOP
01 FORMAT(14I5)
02 FORMAT(9F8.4)
03 FORMAT(48H1POSITION VECTOR AND COEFFICIENTS FOR NN ( NM = ,213,
      + 7H ARE --)
04 FORMAT(7A10)
05 FORMAT(26H0THE MONOPOLE MOMENT IS --,F20.6//)
06 FORMAT(20H1ACTUAL MOMENT IS --/3F20.6/3F20.6)
07 FORMAT(24H1CALCULATED MOMENT IS --/3F20.6/3F20.6)
08 FORMAT(42H0PERCENT ERROR CANNOT BE COMPUTED - DM = 0)
09 FORMAT(20H0PERCENT ERROR IS --/3E20.12)
11 FORMAT(54H0THE QUADRUPOLE MOMENTS Q11,Q22,Q33,Q12,Q13,Q23 ARE --/
      + (3E20.12))
12 FORMAT(27H0THE A ( B COEFFS. FOR THE ,12,16HTH HARMONIC ARE9)
13 FORMAT(1H ,5E20.12)
14 FORMAT(6H0F9 = ,7A10)
      END

      SUBROUTINE AMPMNT(N0,N1,N,R1,F,P,A,B)
      DIMENSION F(1000),P(200),PN(10,200),A(11),B(10)
      DATA PI/3.14159265358979323846/
      CALL POLVAL(N1,N,PN)
      KL=0
      N2=(N1+1)/2
      D1=PI/FLOAT(N0)
      D2=2.*PI/FLOAT(N1-1)
      DO 10 K=1,N0
      A1=D1*FLOAT(K-1)
      LP=+1
      LI=0
      DO 10 L=1,N1
      KL=KL+1
      IF(L.GT.N2) LP=-1
      LI=LI+LP
      A2=A1+FLOAT(1-LP)*PI/2.
      F1=F(KL)
      A2N=FLOAT(2*N+1)
      A1N=FLOAT(N+1)
      AR=(R1*2.54)**(N+2)
      NN=N+1
      DO 10 MM=1,NN
      M=MM-1
      AM=FLOAT(M)
      IF(KL.EQ.1) A(MM)=0.
      IF(KL.EQ.1.AND.M.GT.0) B(M)=0.
      PNM1=PN(MM,LI)*COS(AM*A2)
      PNM2=PN(MM,LI)*SIN(AM*A2)
      AC=(A2N*AR/A1N)*F1*P(LI)
      A(MM)=A(MM)+AC*PNM1
      IF(M.GT.0) B(M)=B(M)+AC*PNM2
10 CONTINUE
      RETURN
      END

      SUBROUTINE WGT1(N0,N1,P)
      DIMENSION D(50),P(200)
      N4=(N1+1)/4
      AN=FLOAT(N0)
      CALL WGT2(N1,D)

```

SA501690
 SA501700
 SA501710
 SA501720
 SA501730
 SA501740
 SA501750
 SA501760
 SA501770
 SA501780
 SA501790
 SA501800
 SA501810
 SA501820
 SA501830
 SA501840
 SA501850
 SA501860
 SA501870
 SA501880
 SA501890
 SA501900
 SA501910
 SA501920
 SA501930
 SA501940
 SA501950
 SA501960
 SA501970
 SA501980
 SA501990
 SA502000
 SA502010
 SA502020
 SA502030
 SA502040
 SA502050
 SA502060
 SA502070
 SA502080
 SA502090
 SA502100
 SA502110
 SA502120
 SA502130
 SA502140
 SA502150
 SA502160
 SA502170
 SA502180
 SA502190
 SA502200
 SA502210
 SA502220
 SA502230
 SA502240
 SA502250
 SA502260
 SA502270
 SA502280

NOLTR 73-191

```

DO 10 J1=1,N4
J3=N1/2+2-J1
DJ=D(J1)/(4.*AN)
P(J1)=DJ
10 P(J3)=DJ
IF(INT(FLOAT(N1-1)/4.+1)*4.NE.(N1-1)) GO TO 20
J1=N4+1
P(J1)=D(J1)/(4.*AN)
20 N4=(N1+1)/2
P(N4)=P(N4)*2.
RETURN
END

```

```

SUBROUTINE WGT2(N,D)
DIMENSION A(50,50),D(50),C(50)
DOUBLE PRECISION A,C,PI,AN
DATA PI/3.141592653589793238462643D0/
N3=(N+1)/4
N4=N3+1
DO 10 I=2,N4
C(I)=1./FLOAT(2*I-1)
A(I,1)=1.
A(I,N4)=0.
DO 10 J=2,N3
AN=2.*PI*DBLE(FLOAT(J-1))/DBLE(FLOAT(N-1))
10 A(I,J)=(DCOS(AN))**(2*I-2)
C(I)=2.
A(1,N4)=1.
DO 20 J=1,N3
20 A(1,J)=2.
N3=N4
IF(INT(FLOAT(N-1)/4.+1)*4.NE.(N-1)) N3=N3-1
CALL GAUSEL(A,C,N3,N3,1)
DO 30 I=1,N3
30 D(I)=C(I)
RETURN
END

```

```

SUBROUTINE WGT3(N0,N1,P)
DIMENSION P(200)
DATA PI/3.14159265358979323846/
D2=2.*PI/FLOAT(N1-1)
E1=(SIN(D2/4.))**2
E2=SIN(D2/2.)
P(1)=E1/(2.*FLOAT(N0))
N2=(N1+1)/2
P(N2)=P(1)*2.
IF(N1.LE.3) RETURN
N3=N2-1
DO 10 L=2,N3
S2=SIN(D2*FLOAT(L-1))
P(L)=ABS(S2*E2)/(2.*FLOAT(N0))
10 CONTINUE
RETURN
END

```

```

SUBROUTINE GAUSEL(A,C,M,N,IT)
DIMENSION A(50,50),C(50),IB(50)
DOUBLE PRECISION A,C,D
C
WRITE(6,03)

```

SA502290
SA502300
SA502310
SA502320
SA502330
SA502340
SA502350
SA502360
SA502370
SA502380
SA502390
SA502400
SA502410
SA502420
SA502430
SA502440
SA502450
SA502460
SA502470
SA502480
SA502490
SA502500
SA502510
SA502520
SA502530
SA502540
SA502550
SA502560
SA502570
SA502580
SA502590
SA502600
SA502610
SA502620
SA502630
SA502640
SA502650
SA502660
SA502670
SA502680
SA502690
SA502700
SA502710
SA502720
SA502730
SA502740
SA502750
SA502760
SA502770
SA502780
SA502790
SA502800
SA502810
SA502820
SA502830
SA502840
SA502850
SA502860
SA502870
SA502880

C	DO 20 I=1,M	SA502890
C	20 WRITE(6,04) (A(I,J),J=1,N),C(I)	SA502900
	DO 70 I=1,N	SA502910
	70 IB(I)=I	SA502920
	DO 60 I=1,M	SA502930
	IF(IT.NE.0.AND.M.NE.N) GO TO 35	SA502940
	D=0.	SA502950
	JJ=0	SA502960
	DO 10 J=I,N	SA502970
	IF(DABS(A(I,J)).LE.D) GO TO 10	SA502980
	JJ=J	SA502990
	D=DABS(A(I,J))	SA503000
	10 CONTINUE	SA503010
	IF(JJ.EQ.1) GO TO 35	SA503020
	IF(JJ.NE.0) GO TO 140	SA503030
	WRITE(6,01) I	SA503040
	GO TO 110	SA503050
	140 DO 160 J=1,M	SA503060
	D=A(J,JJ)	SA503070
	A(J,JJ)=A(J,I)	SA503080
	160 A(J,I)=D	SA503090
	ID=IB(JJ)	SA503100
	IB(JJ)=IB(I)	SA503110
	IB(I)=ID	SA503120
	35 D=A(I,I)	SA503130
	DO 30 J=I,N	SA503140
	30 A(I,J)=A(I,J)/D	SA503150
	C(I)=C(I)/D	SA503160
	DO 50 J=1,M	SA503170
	IF(J.EQ.1) GO TO 50	SA503180
	D=A(J,I)	SA503190
	IF(D.EQ.0.) GO TO 50	SA503200
	DO 40 K=I,N	SA503210
	40 A(J,K)=A(J,K)-D*A(I,K)	SA503220
	C(J)=C(J)-D*C(I)	SA503230
	50 CONTINUE	SA503240
	60 CONTINUE	SA503250
	110 IF(IT.EQ.0.OR.M.NE.N) GO TO 100	SA503260
	DO 120 I=1,N	SA503270
	120 A(I,N)=C(I)	SA503280
	DO 130 I=1,N	SA503290
	II=IB(I)	SA503300
	130 C(II)=A(I,N)	SA503310
	100 CONTINUE	SA503320
C	IF(IT.EQ.0) WRITE(6,02) (IB(I),I=1,N)	SA503330
C	WRITE(6,05)	SA503340
C	DO 80 I=1,M	SA503350
C	80 WRITE(6,04) (A(I,J),J=1,N),C(I)	SA503360
	RETURN	SA503370
	01 FORMAT(5H1ROW ,I2,14H IS ALL ZEROS.)	SA503380
	02 FORMAT(24HOSINGLE PERMUTATION IS.,20I3//)	SA503390
	03 FORMAT(21H0THE INPUT MATRIX IS9//)	SA503400
	04 FORMAT(1H ,7D16.8)	SA503410
	05 FORMAT(22H0THE OUTPUT MATRIX IS9//)	SA503420
	END	SA503430
		SA503440
	SUBROUTINE POLVAL(N1,N,PN)	SA503450
	DIMENSION PN(10,200)	SA503460
	DATA PI/3.14159265358979323846/	SA503470
	D2=2.*PI/FLOAT(N1-1)	SA503480

```

N2=(N1+1)/2
DO 10 I=1,N2
A2=D2*FLOAT(I-1)
C2=COS(A2)
NN=N+1
DO 10 MM=1,NN
M=MM-1
PN(MM,I)=SPNM(N,M,C2)
10 CONTINUE
RETURN
END

SUBROUTINE AMPFLD(R,P,N,A,B,F)
DIMENSION R(3),P(3),A(11),B(10),F(3),U(3),V(3),W(3)
DATA PI/3.14159265358979/
CALL SUM(1.,R,-1.,P,U)
CALL SPCOOR(U,V)
R1=V(1)*2.54
IF(R1.NE.0.) GO TO 10
PRINT 01
RETURN
10 O1=V(2)*PI/180.
O2=V(3)*PI/180.
S1=SIN(O1) $ S2=SIN(O2)
C1=COS(O1) $ C2=COS(O2)
U(1)=C1*S2 $ U(2)=S1*S2 $ U(3)=C2
V(1)=C1*C2 $ V(2)=S1*C2 $ V(3)=-S2
W(1)=-S1 $ W(2)=C1 $ W(3)=0.
H1=H2=H3=0.
NN=N+1
DO 30 MM=1,NN
M=MM-1
P1=FLOAT(M)*O1
SM1=SIN(P1) $ CM1=COS(P1)
P1=SPNM(N,M,C2)
P2=SPNM(N,M+1,C2)
P3=P4=0. $ CP=SQRT(FLOAT(N*(N+1))/2.)
IF(M.EQ.0) GO TO 20
P4=FLOAT((N-M+1)*(N-M+2))*PNM(N+1,M-1,C2)+PNM(N+1,M+1,C2)
P4=.5*SQRT(2.*ANF(N-M)/ANF(N+M))*P4
P3=C2*P4 $ CP=SQRT(FLOAT((N-M)*(N+M+1)))
20 B1=0.
IF(M.GT.0) B1=B(M)
A1=A(MM)*CM1+B1*SM1
H1=H1+A1*P1
H2=H2+A1*(CP*P2-P3)
H3=H3+(A(MM)*SM1-B1*CM1)*P4
30 CONTINUE
R1=R1**((N+2))
H1=FLOAT(N+1)*H1/R1 $ H2=H2/R1 $ H3=H3/R1
CALL SUM(H1,U,H2,V,F) $ CALL SUM(1.,F,H3,W,F)
RETURN
01 FORMAT(50H0FIELD VECTOR CANNOT BE COMPUTED AT POLE POSITION.//)
END

FUNCTION SPNM(N,M,X)
SPNM=PNM(N,M,X)
IF(M.EQ.0.OR.M.GT.N) RETURN
SPNM=SQRT(2.*ANF(N-M)/ANF(N+M))*SPNM
RETURN

```

SA503490
SA503500
SA503510
SA503520
SA503530
SA503540
SA503550
SA503560
SA503570
SA503580
SA503590
SA503600
SA503610
SA503620
SA503630
SA503640
SA503650
SA503660
SA503670
SA503680
SA503690
SA503700
SA503710
SA503720
SA503730
SA503740
SA503750
SA503760
SA503770
SA503780
SA503790
SA503800
SA503810
SA503820
SA503830
SA503840
SA503850
SA503860
SA503870
SA503880
SA503890
SA503900
SA503910
SA503920
SA503930
SA503940
SA503950
SA503960
SA503970
SA503980
SA503990
SA504000
SA504010
SA504020
SA504030
SA504040
SA504050
SA504060
SA504070
SA504080

```

END
FUNCTION PNM(N,M,X)
PNM=0.
IF(M.GT.N) RETURN
IF((ABS(1.-ABS(X)))>.1.E-9) GO TO 20
IF(M.NE.0) RETURN
PNM=-1.
IF(X.GT.0..OR.N.EQ.INT(FLOAT(N)/2.+1)*2) PNM=1.
RETURN
20 CNM=(2.**N)*ANF(N)*ANF(N-M)
PNM=1.
IF(M.NE.0) PNM=SQRT(1.-X*X)**M
CNM=ANF(2*N)*PNM/CNM
PNM=1.
IF(M.NE.N) PNM=X**(N-M)
IF(N-M.LE.1) GO TO 40
PRD1=1.
NT=(N-M)/2
DO 30 I=1,NT
AN1=FLOAT(N-M-2*I+2)
AN2=FLOAT(2*I)
AN3=FLOAT(2*N-2*I+1)
PRD1=-PRD1*AN1*(AN1-1)/(AN2*AN3)
NE=N-M-2*I
AN1=1.
IF(NE.GT.0) AN1=X**NE
PNM=PNM*PRD1*AN1
30 CONTINUE
40 PNM=CNM*PNM
RETURN
END

FUNCTION ANF(N)
DOUBLE PRECISION AN
ANF=1.
AN=1.D0
IF(N.LT.0) PRINT 01
IF(N.LT.2) RETURN
DO 10 I=2,N
10 AN=AN*DBLE(FLOAT(I))
ANF=SNGL(AN)
RETURN
01 FORMAT(37H1FACTORIAL INTEGER IS LESS THAN ZERO.//)
END

SUBROUTINE SPCOOR(D,R)
DIMENSION D(3),R(3)
DATA PI/3.14159265358979/
R(1)=SQRT(DOT(D,D))
R(2)=0.
R(3)=0.
IF(R(1).EQ.0.) RETURN
IF(D(1)**2+D(2)**2.NE.0.) R(2)=ATAN2(D(2),D(1))*180./PI
R(3)=ATAN2(SQRT(D(1)**2+D(2)**2),D(3))*180./PI
RETURN
END

SUBROUTINE SUM(A,X,B,Y,Z)
DIMENSION X(3),Y(3),Z(3)

```

SA504090
SA504100
SA504110
SA504120
SA504130
SA504140
SA504150
SA504160
SA504170
SA504180
SA504190
SA504200
SA504210
SA504220
SA504230
SA504240
SA504250
SA504260
SA504270
SA504280
SA504290
SA504300
SA504310
SA504320
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SA504370
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SA504660
SA504670
SA504680


```

DO 10 I=1,3
10 Z(I)=A*X(I)+B*Y(I)
RETURN
END

FUNCTION DOT(X,Y)
DIMENSION X(3),Y(3)
DOT=X(1)*Y(1)+X(2)*Y(2)+X(3)*Y(3)
RETURN
END

SUBROUTINE PRINTF(F,NO,N1)
DIMENSION F(800),F1(100)
DO 20 KO=1,NO
DO 10 I=1,N1
II=I+(KO-1)*N1
10 F1(I)=F(II)*1.E5
WRITE(6,01) KO
20 WRITE(6,02) (F1(KK),KK=1,N1)
01 FORMAT(11H0CURVE NO. ,I2)
02 FORMAT(1H ,8F8.1)
RETURN
END

SUBROUTINE DATPLT(NO,N1,F,YDIST)
DIMENSION F(1000),F1(33),F2(73),X(33),X1(73),C(1),P(1),CH(16)
DATA PI/3.14159265358979/
DATA CH/1H1,1H2,1H3,1H4,1H5,1H6,1H7,1H8,1H9,2H10,2H11,2H12,2H13,
+ 2H14,2H15,2H16/
N3=NO*N1
CALL YSCALE(N3,F,YDIST,YMAX,YMIN)
DX=360./FLOAT(N1-1)
F1(1)=F1(2)=YMAX $ X(1)=0. $ X(2)=360.
CALL CALCM1(2,X,F1,0,0.,360.,YMIN,YMAX,6.,7.,TITLE,0,
+ 20HCOLATITUDE (DEGREES),-20,30HNORMAL MAGNETIC FIELD (GAMMAS),
+ 30,0.,18)
F1(2)=YMIN $ X(1)=X(2)=360.
CALL CALCM1(-2,X,F1,0)
F1(1)=F1(2)=0. $ X(1)=360. $ X(2)=0.
IF(YMAX.GT.0.0.AND.YMIN.LT.0.0) CALL CALCM1(-2,X,F1,0)
DO 10 I=1,N1
10 X(I)=DX*FLOAT(I-1)
DO 20 I=1,73
20 X1(I)=5.*FLOAT(I-1)
IJ=0
DO 50 I=1,NO
DO 30 J=1,N1
IJ=IJ+1
30 F1(IJ)=F(IJ)*1.E+5
CALL FNCTON(F1,N1-1,C,CO,P,N2,DEV)
DO 40 N=1,N2
EN=N
DO 40 J=1,73
XX=X1(J)*PI/180.
IF(N.EQ.1) F2(J)=CO
40 F2(J)=F2(J)+C(N)*SIN(EN*XX+P(N))
CALL CALCM1(-73,X1,F2,0)
DO 50 J=1,N1
NC=1+I/10
CALL SYMBL4(X(J)/60.,(F1(J)-YMIN)/YDIST,.08,CH(I),0.,NC)

```

SA504690
SA504700
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SA504970
SA504980
SA504990
SA505000
SA505010
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SA505100
SA505110
SA505120
SA505130
SA505140
SA505150
SA505160
SA505170
SA505180
SA505190
SA505200
SA505210
SA505220
SA505230
SA505240
SA505250
SA505260
SA505270
SA505280

```

50 CONTINUE
RETURN
END

SUBROUTINE FNCTON(F,NSCANS,C,CO,PHI,N2,DEV)
DIMENSION F(1),G(33),C(1),PHI(1)
DATA PI/3.14159265358979/
DX=2.*PI/FLOAT(NSCANS)
NE=NSCANS/2+1
CO=0.
DO 10 J=1,NSCANS
10 CO=CO+F(J)/FLOAT(NSCANS)
DO 30 N=1,NE
EN=N
S1=0.
C1=0.
X=-DX
DO 20 J=1,NSCANS
X=X+DX
S1=S1+(1./PI)*F(J)*SIN(EN*X)*DX
20 C1=C1+(1./PI)*F(J)*COS(EN*X)*DX
C(N)=SQRT(S1**2+C1**2)
IF((FLOAT(N)).EQ.(FLOAT(NSCANS)/2.)) C(N)=C(N)/2.
PHI(N)=0.0
30 IF((S1**2+C1**2).NE.0.) PHI(N)=ATAN2(C1,S1)
DO 50 N=1,NE
EN=N
X=-DX
DEV1=0.
DO 40 J=1,NSCANS
X=X+DX
IF(N.EQ.1) G(J)=CO
G(J)=G(J)+C(N)*SIN(EN*X+PHI(N))
40 IF(ABS(G(J)-F(J)).GT.DEV1) DEV1=ABS(G(J)-F(J))
IF(N.EQ.1) DEV=DEV1
IF(N.EQ.1) N2=1
IF(DEV1.LT.DEV) N2=N
IF(DEV1.LT.DEV) DEV=DEV1
50 CONTINUE
RETURN
END

SUBROUTINE YSCALE(N,F,YD,YX,YM)
DIMENSION F(800)
PP=F(1)
PN=PP
DO 10 I=1,N
IF(F(I).GT.PP) PP=F(I)
IF(F(I).LT.PN) PN=F(I)
10 CONTINUE
PP=PP*1.E+5
PN=PN*1.E+5
P=PP-PN
SN=0.
IF(PP.GT.0.) SN=1.
IF(YD.NE.0.) GO TO 50
DO 20 I=1,21
FA=(1.E-9)*(10.**I)
FA1=FA*.1
IF(P/FA.GE.1.) GO TO 20

```

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SA505790
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SA505860
SA505870
SA505880

GO TO 30	SA505890
20 CONTINUE	SA505900
30 YD=5.*FA1	SA505910
YX=FLOAT(INT(PP/YD+SN))*YD	SA505920
YM=YX-7.*YD	SA505930
CN=4.	SA505940
DO 40 I=1,3	SA505950
CN=CN*.5	SA505960
YD1=CN*FA1	SA505970
YX1=FLOAT(INT(PP/YD1+SN))*YD1	SA505980
YM1=YX1-7.*YD1	SA505990
IF((7.*YD1-P).LT.0.0.OR.YX1.LT.PP.OR.YM1.GT.PN) RETURN	SA506000
YD=YD1	SA506010
YX=YX1	SA506020
YM=YM1	SA506030
40 CONTINUE	SA506040
50 YX=FLOAT(INT(PP/YD+SN))*YD	SA506050
YM=YX-7.*YD	SA506060
RETURN	SA506070
END	SA506080

APPENDIX J
SAMPLE PROBLEMS FOR SA5024

I. PROBLEM EXECUTED ON INTERCOM (TIME-SHARING SYSTEMS)

NØL INTERCØM
TYPE "LØGIN."
LØGIN(S)
024533LACK/ [REDACTED] /4

09/14/73 13.26.48. BC/42/35
C- SETUP.GENERAL

ØN AT 13.27.14. 09/14/73
**GENERAL
**NEW ØR ØLD FILE- ATTACH(AAA,BN5024)*REWIND(AAA)*CØPYBR(AAA,FIL,15)

13.27.44.ATTACH(AAA,BN5024)
**READY.
RETURN(AAA)

**READY.
FIL.

16	33	1	8	0	0	0
96.00	0.000	0.000	0.000			
<u>(7E10.4)</u>						

\$NAM

NO = 16,
N1 = 33,
NH1 = 1,

NH2 = 8,
 IP1 = 0,
 IP2 = 0,
 IW = 0,
 R1 = 0.96E+02,
 EG = 0.0,
 EA = 0.0,
 ED = 0.0,
 PY = -0.0,

SEND

F9 = (7E10.4)
1 3

POSITION VECTOR AND COEFFICIENTS FOR NN & NM = 1 3 ARE --

39.00 0.000 0.000
+.0000E+00+.3100E+05
+.0000E+00

39.0000 0.0000 0.0000

. .3100E+05

.
-21.00 36.373 0.000
+.0000E+00-.1500E+05
+.2598E+05

21.0000 36.3730 0.0000

. -.1500E+05
 .2598E+05

-21.00 -36.373 0.000
+ .0000E+00 - .1500E+05
- .2598E+05

21.0000-36.3730 0.0000

- .1500E+05

.2598E+05

THE MONOPOLE MOMENT IS -- .000005

ACTUAL MOMENT IS --

1000.000000

0.000000

0.000000

1000.000000

0.000000

90.000000

CALCULATED MOMENT IS --

999.999347

.000000

-.000000

999.999347

.000000

90.000000

PERCENT ERROR IS --

.653140392387E-04 -.367553809610E-09 .204352090805E-10

THE A & B COEFFS. FOR THE 1TH HARMONIC ARE:

-.204352090805E-09 .999999346860E+03

.367553809610E-08

THE QUADRUPOLE MOMENTS Q11,Q22,Q33,Q12,Q13,Q23 ARE --

.908333976792E+07 .985968240459E+07 -.189430221725E+08

-.140861894026E-05 .277251221753E-08 -.114908029088E-06

THE A & B COEFFS. FOR THE 2TH HARMONIC ARE:

-.947151108626E+07 .160071067512E-08 -.224110815131E+06
-.663421815261E-07 -.813266524347E-06

THE A & B COEFFS. FOR THE 3TH HARMONIC ARE:

.468939542770E-04 .683659431170E+08 -.223517417908E-07 .23409232
3833E+10
-.814609229565E-04 .884430482984E-05 -.183412397746E-03

THE A & B COEFFS. FOR THE 4TH HARMONIC ARE:

.154465924862E+12 -.776198242188E-03 .140584755795E+11 -.10529816
1507E-01 -.186018026985E+11
.351142883301E-02 .118371725082E+00 -.733375549316E-03 -.24567153
3048E-01

THE A & B COEFFS. FOR THE 5TH HARMONIC ARE:

-.487304687500E+00 -.217954001634E+13 -.256347656250E-01 -.28122218
7225E+14 -.820884704590E-01
-.315797097447E+13
.177050781250E+01 .128906250000E+00 .177172851563E+01 -.77190017
7002E+00 -.294870560169E+02

THE A & B COEFFS. FOR THE 6TH HARMONIC ARE:

-.210884991589E+16 -.195000000000E+02 -.322787259184E+15 .10126562
-.210884991589E+16 -.195000000000E+02 -.322787259184E+15 .10126562
5000E+03 .353625883682E+15
.225479736328E+02 .453272232211E+16
.128687500000E+03 .104384375000E+04 .885156250000E+01 .17665273
4375E+04 -.262702636719E+02
-.114442594910E+04

THE A & B COEFFS. FOR THE 7TH HARMONIC ARE:

-.225424000000E+06 .431994422628E+17 .172800000000E+04 .35406399
7182E+18 -.101790000000E+05
.496241920808E+17 -.293837539062E+05 -.676359128706E+17
-.267840000000E+05 -.154800000000E+05 -.230640000000E+05 -.12074000
0000E+05 .115108750000E+06
-.933500781250E+04 -.660728401367E+06

THE A & B COEFFS. FOR THE 8TH HARMONIC ARE:

.269942772159E+20 .206438400000E+07 .569785815443E+19 -.60129280
0000E+07 -.597539013375E+19
-.156819200000E+07 -.451694469107E+20 .765785875000E+06 -.91015764
9666E+19
-.425574400000E+07 .626974720000E+08 .271462400000E+07 .52427776
0000E+08 .175494400000E+07
.531896320000E+08 -.174570300000E+07 -.105260966187E+09

Q

13.38.54.STOP

**READY.

LOGOUT.

CP TIME 22.322
PP TIME 106.909
CONNECT TIME 0 HR 12 MIN 22 SEC
TOTAL COST OF SESSION = \$ 6.16
09/14/73 LOGGED OUT AT 13.39.10.<

Notes:

1. The file BN5024 is the binary version of SA5024. It consists of 18 binary records (subprograms). The last three are plotting routines and are not used when executing problems on INTERCOM.
2. The information typed in by the user has been underlined.

APPENDIX J (CONT.)
II. PROBLEM SUBMITTED TO BATCH

NØL INTERCØM
TYPE "LØGIN."
LØGIN(S)
024533LACK/ /4

09/14/73 13.40.17. BC/42/35
C- SETUP.GENERAL

ØN AT 13.40.33. 09/14/73
**GENERAL
**NEW ØR ØLD FILE- NEW/IECC024*TAPE(ØN)

**READY.
1 IECC5ST,P1,T200,CM060000.55302435,024,LACKEY.
2 ATTACH(ABC,NØLBIN)
3 CØPYN(O,DEF,ABC)
4 RETURN(ABC)
5 ATTACH(BN5024,BN5024)
6 REWIND(BN5024)
7 CØPYBF(BN5024,CBA)
8 RETURN(BN5024)
9 LØAD(CBA)
1 ODEF.
11 *WEØR
12 REWIND(ABC)
13 GØULD1,14,ABC
14 *WEØR
6100 16 33 1 8 1 1 0
6110 96.00 0.000 0.000 0.000 0.000
6120 (9F8.4)
6130 1 1
6140 20.00 0.000 0.000
6150 0.000 1000.
6160 0.000
6170 16 33 1 8 1 1 0
6180 96.00 0.000 0.000 0.000 0.000
6190 (9F8.4)
6200 1 1
6210 20.00 0.000 0.000
6220 0.000 0.000
6230 1000.
6240 0 0 0 0 0 0 0
TAPE(ØFF)

Tape prepared
beforehand in
LOCAL mode

**READY.
SAVE*PURGE(IECC024)*BATCH.*QUEUES.

**SAVED IECC024
13.43.50.PURGE(IECC024)
TYPE FILE NAME-IECC024

TYPE DISPOSITION-INPUT

TYPE FILE NAME-END

QUEUES 13.44.38. I= 7, O= 0, P= 0, C= 3.
INPUT = 7
CABAP72-2 BDCR07K-1 BDC9R7N-3 BDCSH7U-1 IECC579-1 DCCRW78-5
CESEQ02-3
OUTPUT= NONE
PUNCH = NONE
COMMON= 3
FARE2 -0 SSSSSSU-0 SSSSSST-0
CONTROL PTS.
ICCB63-3 GRIDA7S-5 CABTR7T-5 CBCQ26K-1 AUDIT73-4 HHJL371-2
CCB7870-5
13.44.38.STOP
**READY.
LOGOUT.

CP TIME .785
PP TIME 96.846
CONNECT TIME 0 HR 5 MIN 4 SEC
TOTAL COST OF SESSION = \$ 2.14
09/14/73 LOGGED OUT AT 13.45.21.<

Note:

1. The information typed in by the user has been underlined.

NOLTR 73-191

APPENDIX K

NOL TECHNICAL NOTE 9726
"NUMERICAL COMPUTATION OF LEGENDRE
POLYNOMIALS AND SPHERICAL FUNCTIONS"

TN 9726

FOR INTERNAL USE ONLY

NAVAL ORDINANCE LABORATORY
White Oak, Silver Spring, Maryland

533:MHL:bag

NUMERICAL COMPUTATION OF LEGENDRE POLYNOMIALS AND
SPHERICAL FUNCTIONS

7 November 1972

M. H. Lackey
Magnetic Structures Group

Asgmt: Task No. NOL-786/NPL (Magnetic Calibration for NPL)
TWP No. 530-524

Abst: This report describes computer subroutines for the generation of Legendre polynomials and spherical functions. The subroutines are coded in the FORTRAN IV computer language. A brief description of the functions and some of their properties is also included. A comparison is given of the numerical stability of two different methods of generating the functions.

Ref: (a) S. Chapman & J. Bartels, Geomagnetism (Oxford Press, London, 1940), Vol. II
(b) W. D. Macmillan, The Theory of the Potential (McGraw-Hill Book Co., New York, 1930)
(c) ASM 55 Handbook of Mathematical Functions (U.S. Government Printing Office, Washington, D.C., 1964)
(d) NOLTR 69-60, The Reaction of a Rigid Body to a Uniform Force Field, 17 Mar 1969, M. H. Lackey

Encl: (1) Figure 1, Regular Legendre Polynomials $P_n(\cos\theta)$ For $n = 0$ through 7
(2) Figure 2, Associated Legendre Polynomials $P_{7,m}(\cos\theta)$ For $n = 0$ through 7
(3) Figure 3, Schmidt Polynomials $P_7^m(\cos\theta)$ For $m = 0$ through 7
(4) Figure 4, Unstable Procedure for $P_n^{n-1}(\cos\theta)$ For $n = 15$ through 20
(5) Figure 5, Stable Procedure for $P_n^{n-1}(\cos\theta)$ For $n = 15$ through 20
(6) Appendix A, Listing of Function Subroutines

INTRODUCTION

1. Computer subroutines have been devised to compute values for three special types of polynomials including regular and associated Legendre polynomials, and Schmidt functions. The functions have special orthogonality properties which make them especially useful for interpolation and approximation. Techniques in spherical harmonic analysis can be defined in terms of the associated Legendre polynomials (spherical functions). The computer subroutines can be used to develop numerical methods for the harmonic analysis techniques. A listing of the subroutines in FORTRAN IV is included in Appendix A.

2. The subroutines are based on a finite series definition for the polynomials. It was determined that the series definition produced a more stable method than methods based on recurrence formulas.

POLYNOMIAL DEFINITIONS

3. There are a variety of ways to define the polynomials depending on the use intended for them. The definitions and terminology used in the following definitions conform to the usage in references (a) and (b). Some of the recurrence formulas were obtained from reference (c).

4. The regular Legendre polynomials $P_n(x)$ can be defined on the interval $-1 \leq x \leq 1$ by Rodrique's formula,

$$P_n(x) = \frac{1}{2^n n!} \left(\frac{d}{dx} \right)^n (x^2 - 1)^n \quad (1)$$

for $n \geq 0$ and $|x| \leq 1$. The associated Legendre polynomials or spherical functions $P_{n,m}(x)$ can now be defined as

$$P_{n,m}(x) = (1-x^2)^{m/2} \left(\frac{d}{dx} \right)^m P_n(x) = \frac{(1-x^2)^{m/2}}{2^n n!} \left(\frac{d}{dx} \right)^{m+n} (x^2 - 1)^n \quad (2)$$

for $m \leq 0$, $n \geq 0$, and $|x| \leq 1$. The index n designates the degree of the polynomial and the index m designates the order. It should be noted that some definitions of $P_{n,m}(x)$ contain a factor of $(-1)^m$ depending on whether the polynomials are considered to be functions of x on the interval $-1 \leq x \leq 1$ or as functions of θ (with $x = \cos \theta$) on the interval $0 \leq \theta \leq \pi$. In this report the factor is dropped in keeping with the definitions in reference (a).

5. A comparison of Equations (1) and (2) show that the regular Legendre polynomials $P_n(x)$ are a subset of the associated Legendre polynomials $P_{n,m}(x)$, i.e., $P_n(x) = P_{n,0}(x)$. Also Equation (2) leads to the fact that $P_{n,m}(x) = 0$ for $m > n$.

6. Graphs of typical families of the polynomials are shown in Figures 1 and 2.

The regular Legendre polynomials shown in Figure 1 include P_0, P_1, \dots, P_7 . The polynomial values $P_n(\cos\theta)$ are plotted as functions of θ for $0 \leq \theta \leq 180^\circ$. Figure 2 shows the family $P_{7,m}(\cos\theta)$ for $m = 0, 1, 2, \dots, 7$. Notice the variation in the value of the peaks of each curve. The curves for $m < 4$ appear as straight lines coinciding with the zero axis. This extensive variation in the order of magnitude for different polynomials $P_{n,m}(x)$ has disadvantages, especially in the development of numerical analysis procedures. This led Schmidt (see reference (a)) to develop a new set of polynomials.

7. The Schmidt polynomials $P_n^m(x)$ are defined by scaling the associated Legendre polynomials as follows

$$\left\{ \begin{array}{l} P_n^0(x) - P_{n,0}(x) = P_n(x) \\ P_n^m(x) = \left\{ \frac{2 \cdot (n-m)!}{(n+m)!} \right\}^{1/2} P_{n,m}(x) \end{array} \right\} \quad (3)$$

for $m > 0$ and $|x| \leq 1$. The resulting polynomials all have values of the same order of magnitude. Also, $|P_n^m(x)| \leq 1$ for $|x| \leq 1$. A typical family of Schmidt polynomials $P_7^m(\cos\theta)$ for $m = 0, 1, 2, \dots, 7$ is shown in Figure 3.

NUMERICAL METHODS

8. There are many ways of numerically generating sets of polynomials. One of the most common methods is the use of recurrence formulas. The associated Legendre polynomials satisfy several recurrence formulas including

$$P_{n,m}(x) = [(2n-1) \cdot x \cdot P_{n-1,m}(x) - (n+m-1)P_{n-2,m}(x)] / (n-m) \text{ for } m \neq n \text{ and } n > 1 \quad (4)$$

$$P_{n,m}(x) = [(m-n-1) \cdot x \cdot P_{n,m-1}(x) + (n+m-1)P_{n-1,m-1}(x)] / (1-x^2)^{1/2} \text{ for } |x| < 1, \\ m > 0, \text{ and } n > 0 \quad (5)$$

$$P_{n,m}(x) = 2(m-1) \cdot x \cdot (1-x^2)^{-1/2} P_{n,m-1} - (n+m-1)(n-m+2)P_{n,m-2}(x) \text{ for } |x| < 1 \\ \text{and } m > 1. \quad (6)$$

Equation (4) represents a recurrence formula with varying degree, Equation (5) represents both varying degree and varying order, and Equation (6) represents varying order.

9. Another method for generating the associated Legendre polynomials is given

in reference (a). The polynomials are expressed as a finite (alternating) series in powers of x . The expression is:

$$P_{n,m}(x) = \frac{(2n)! (1-x^2)^{m/2}}{2^n \cdot n! (n-m)!} \left\{ x^{n-m} - \frac{(n-m)(n-m-1)x^{n-m-2}}{2(2n-1)} \right. \\ \left. + \frac{(n-m)(n-m-1)(n-m-2)(n-m-3)x^{n-m-4}}{2 \cdot 4(2n-1)(2n-3)} - \dots \right\} \quad (7)$$

where the last term inside the bracket is constant if $(n-m)$ is even; or is a multiple of x if $(n-m)$ is odd.

NUMERICAL STABILITY

10. Several different methods based on the recurrence formulas were used to test the numerical stability for large values of n and m , and for small values of (x^2-1) . The only recurrence formula with any stability was Equation (4) which has varying degree. This formula can be used by itself only to generate the regular Legendre polynomials. When either Equation (5) or (6) is added the method becomes unstable. For example, Figure 4 shows a family of $P_n^{n-1}(\cos \theta)$ for $n = 15, 16, \dots, 20$. Notice that instability becomes worse as $|x|$ approaches 1 and as n increases.

11. Instead of trying to stabilize the methods using the recurrence formulas it was decided to try the series Equation (7). Although the equation appears to be unstable itself (alternating series can be unstable) the results indicate the opposite. Figure 5 shows the same polynomials displayed in Figure 4 but generated by Equation (7). There seems to be no problem for $|x|$ near one or for large n . It was therefore decided to use Equation (7) for the subroutine. The computing time using Equation (7) was about the same as methods using recurrence formulas.

SUBROUTINE CALL STATEMENTS

12. The polynomial values can be computed as follows: Assume that y is a variable used in a computer program and that y is to be set equal to a polynomial value. Then

$y = P_n(x)$ is written as $Y = PNM(N,0,X)$ for $N \geq 0$ and $|x| \leq 1$.

$y = P_{n,m}(x)$ is written as $Y = PNM(N,M,X)$ for $N \geq 0$, $M \geq 0$, and $|x| \leq 1$.

$y = P_n^m(x)$ is written as $Y = SPNM(N,M,X)$ for $N \geq 0$, $M \geq 0$, and $|x| \leq 1$.

There are three function subprograms in the package. These include:

FUNCTION SPNM(N,M,X)

FUNCTION PNM(N,M,X)

FUNCTION ANF(N)

The subroutine ANF(N) is the factorial subroutine, i.e.,

$y = n!$ is written as $Y = \text{ANF}(N)$ for $N \geq 0$.

The factorial value is return as a floating point number based on the integer N. The subroutine SPNM uses both PNM and ANF as external functions. The subroutine PNM uses ANF as an external function. The subroutine ANF requires no external functions (machine functions only).

USES OF THE FUNCTIONS $P_n^m(x)$

13. It can be shown that the set of polynomials $\{P_n^m(x)\}$ of fixed order m are orthogonal on the interval $-1 \leq x \leq 1$ for all $n \geq m$. In fact:

$$\int_{-1}^{+1} P_n^m(x) P_m^m(x) dx = \delta_{nn}, \frac{1}{2n+1} \quad (8)$$

for $n \geq m$ where δ_{nn} is the Kronecker delta. The polynomials can therefore be normalized and used for interpolation and expansion of arbitrary functions on the interval $-1 \leq x \leq 1$.

14. The polynomials are of primary importance in areas of spherical harmonic analysis and potential theory. If we define two new sets of functions as

$$\{\sin(m\varphi) P_n^m(\cos\theta)\}, \{\cos(m\varphi) P_n^m(\cos\theta)\} \quad (9)$$

and consider φ and θ to be the spherical angles representing longitude and colatitude respectively, then the functions are called spherical surface harmonics, and are orthogonal and complete on the surface of the unit sphere. They satisfy the condition

$$\int_0^\pi \int_0^{2\pi} \left[P_n^m(\cos\theta) \begin{Bmatrix} \cos m\varphi \\ \sin m\varphi \end{Bmatrix} \right] \left[P_{n'}^m(\cos\theta) \begin{Bmatrix} \cos m\varphi \\ \sin m\varphi \end{Bmatrix} \right] \sin\theta d\theta d\varphi = \delta_{nn'} \delta_{mm'} \left\{ \frac{4\pi}{2n+1} \right\} \quad (10)$$

for $n, n', m, m' \geq 0$ (see reference (a)). Then any continuous function $f(\theta, \varphi)$ on the surface of the unit sphere can be expanded into a uniformly convergent series of surface harmonics as

$$f(\theta, \varphi) = \sum_{n=0}^{\infty} S_n(\theta, \varphi) = \sum_{n=0}^{\infty} \left\{ \sum_{m=0}^n (A_n^m \cos m\varphi + B_n^m \sin m\varphi) P_n^m(\cos \theta) \right\} \quad (11)$$

where $S_n(\theta, \varphi)$ represents a general surface harmonic of degree n . The constants A_n^m and B_n^m are determined from the equations

$$\begin{pmatrix} A_n^m \\ B_n^m \end{pmatrix} = \frac{2n+1}{4\pi} \int_0^\pi \int_0^{2\pi} f(\theta, \varphi) P_n^m(\cos \theta) \begin{pmatrix} \cos m\varphi \\ \sin m\varphi \end{pmatrix} \sin \theta d\theta d\varphi \quad (12)$$

for $m \geq 0$ and $n \geq 0$.

15. Numerical integrating schemes for uniformly spaced data $f_{ij} = f(\theta_i, \varphi_j)$ on the surface of the unit sphere are discussed in reference (d).

SUMMARY

16. This report has discussed briefly some of the properties of Legendre polynomials and some of their uses. The subroutines listed in Appendix A can be used in a variety of ways including interpolation, approximation, and spherical harmonic analysis. Plans are being made to develop other subroutines to perform the spherical harmonic analysis and extrapolation of potential functions.

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FIG. 1. REGULAR LEGENDRE POLYNOMIALS $P_n(\cos \theta)$ FOR $n = 0$ THROUGH 7

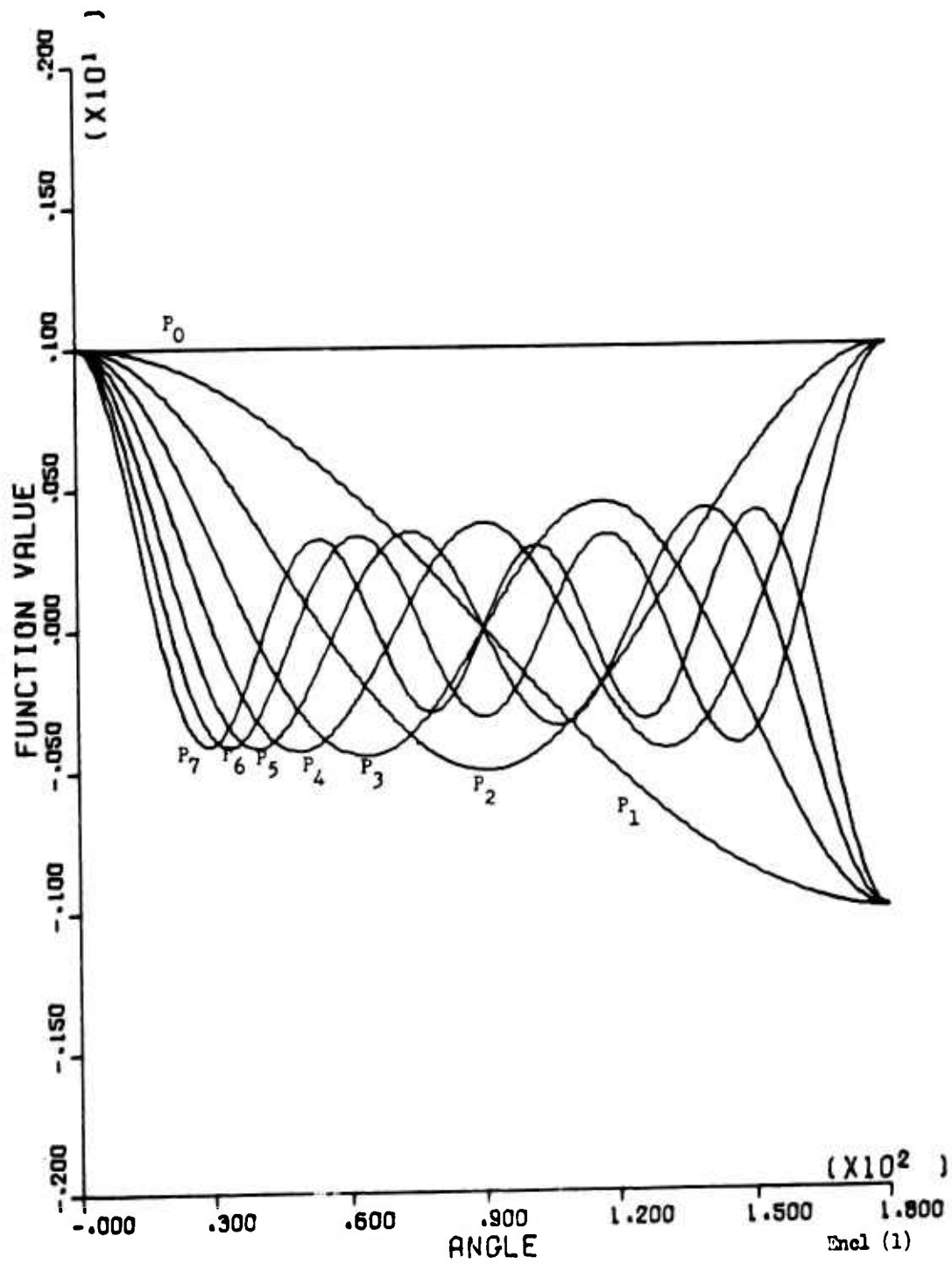


FIG. 2. ASSOCIATED LEGENDRE POLYNOMIALS $P_{7,m}(\cos\theta)$ FOR $m = 0$ THROUGH 7

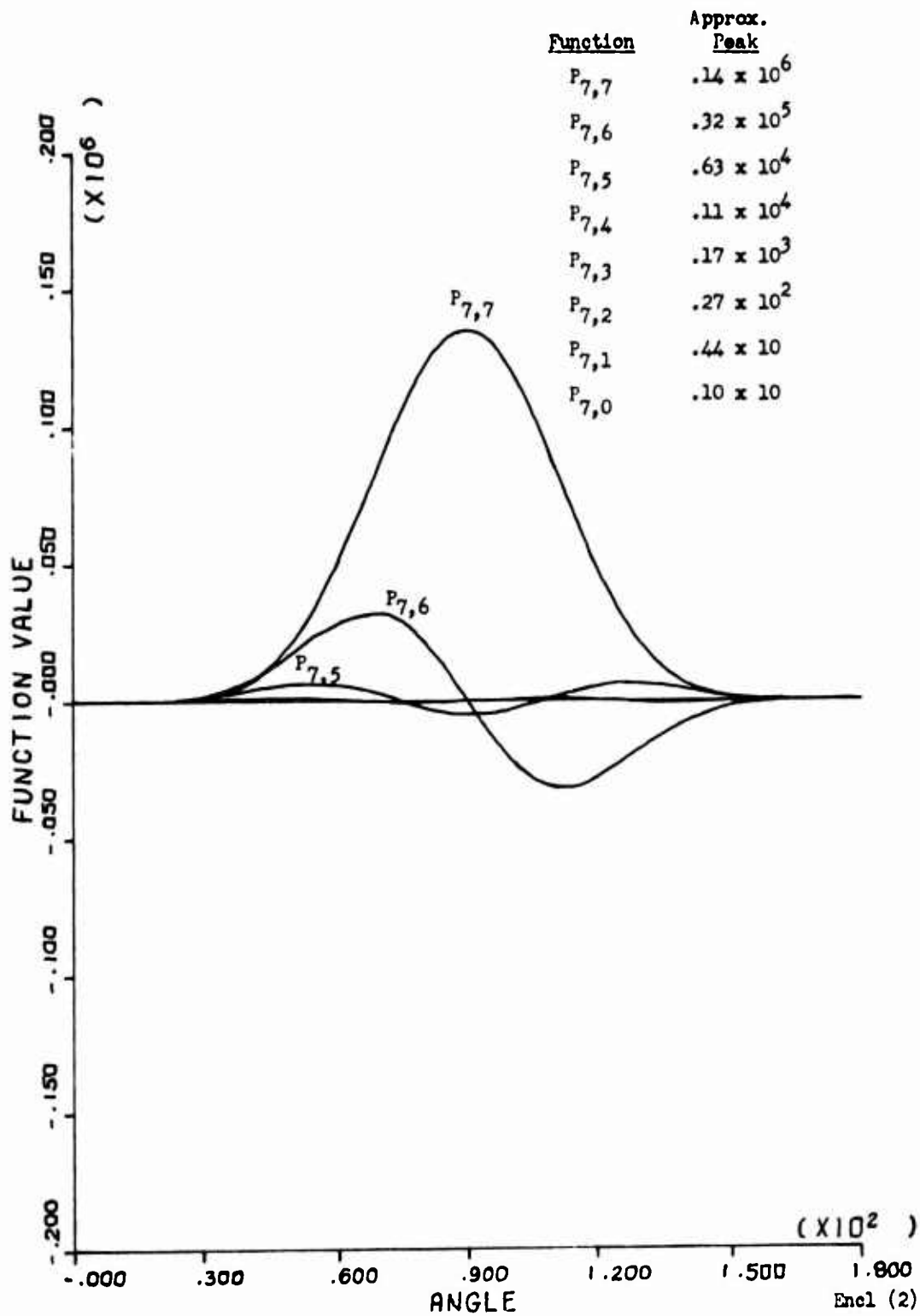


FIG. 3. SCHMIDT POLYNOMIALS $P_7^m(\cos\theta)$ FOR $m = 0$ THROUGH 7

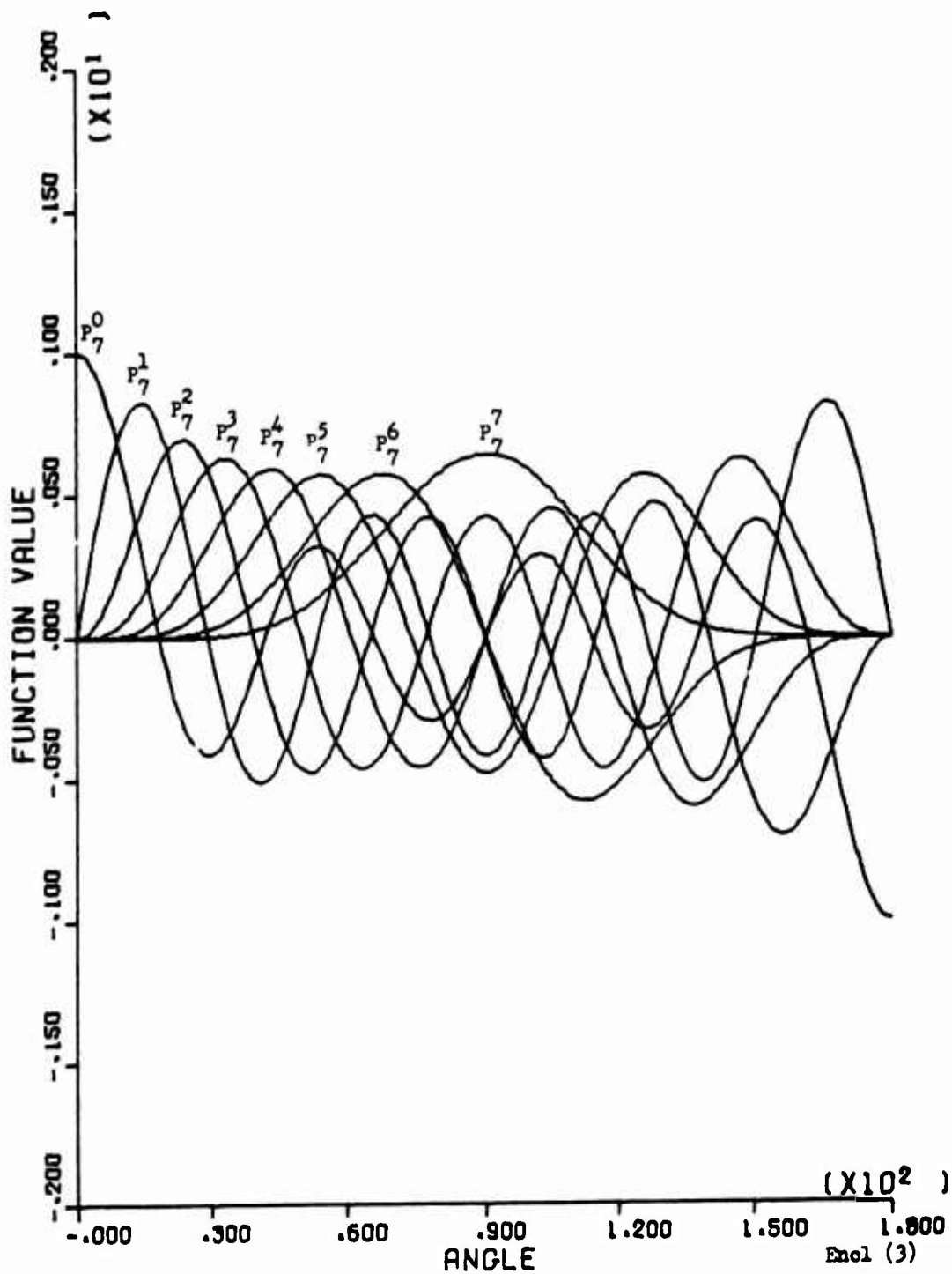


FIG. 4. UNSTABLE PROCEDURE FOR $P_n^{n-1}(\cos\theta)$ FOR $n = 15$ THROUGH 20

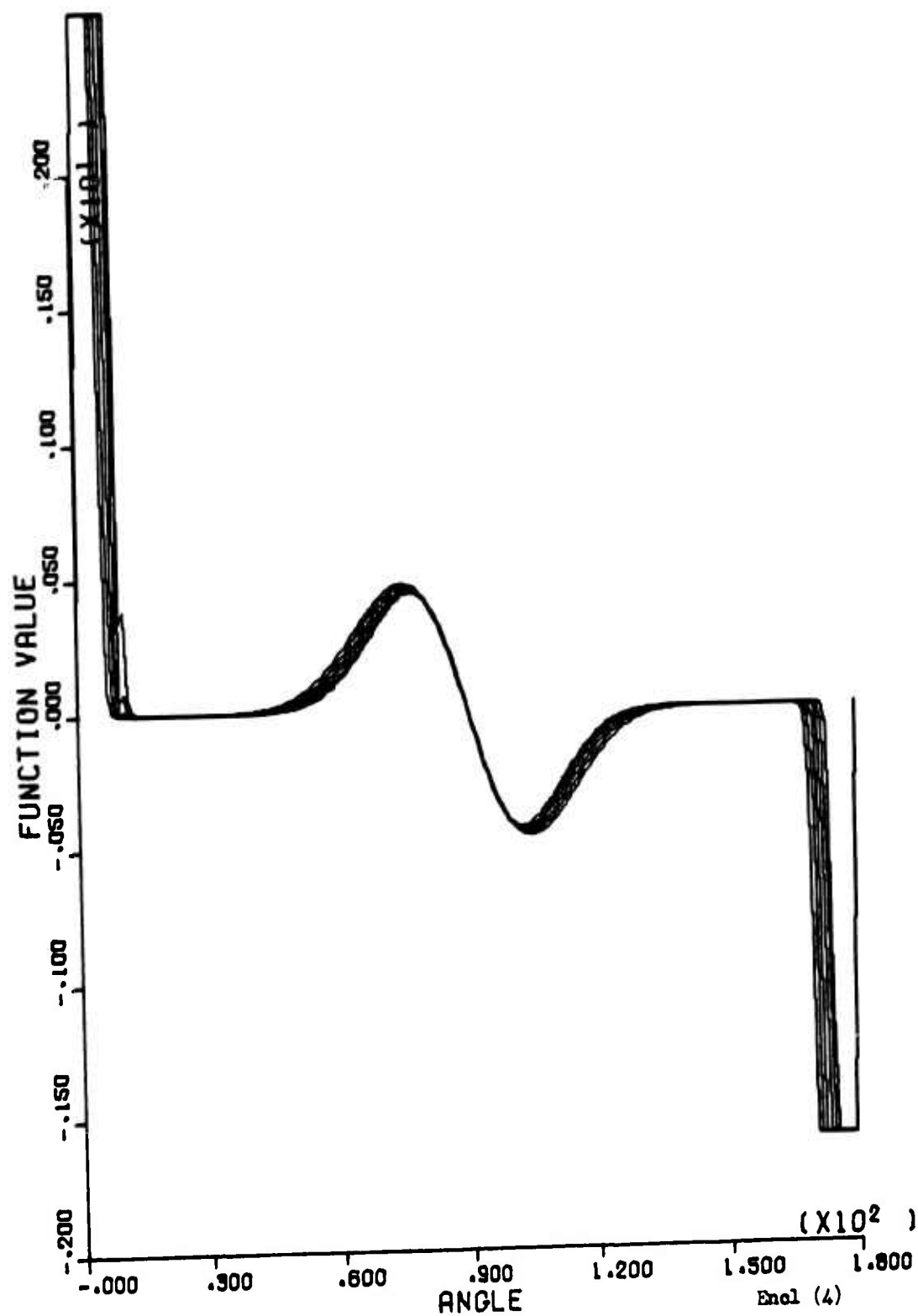
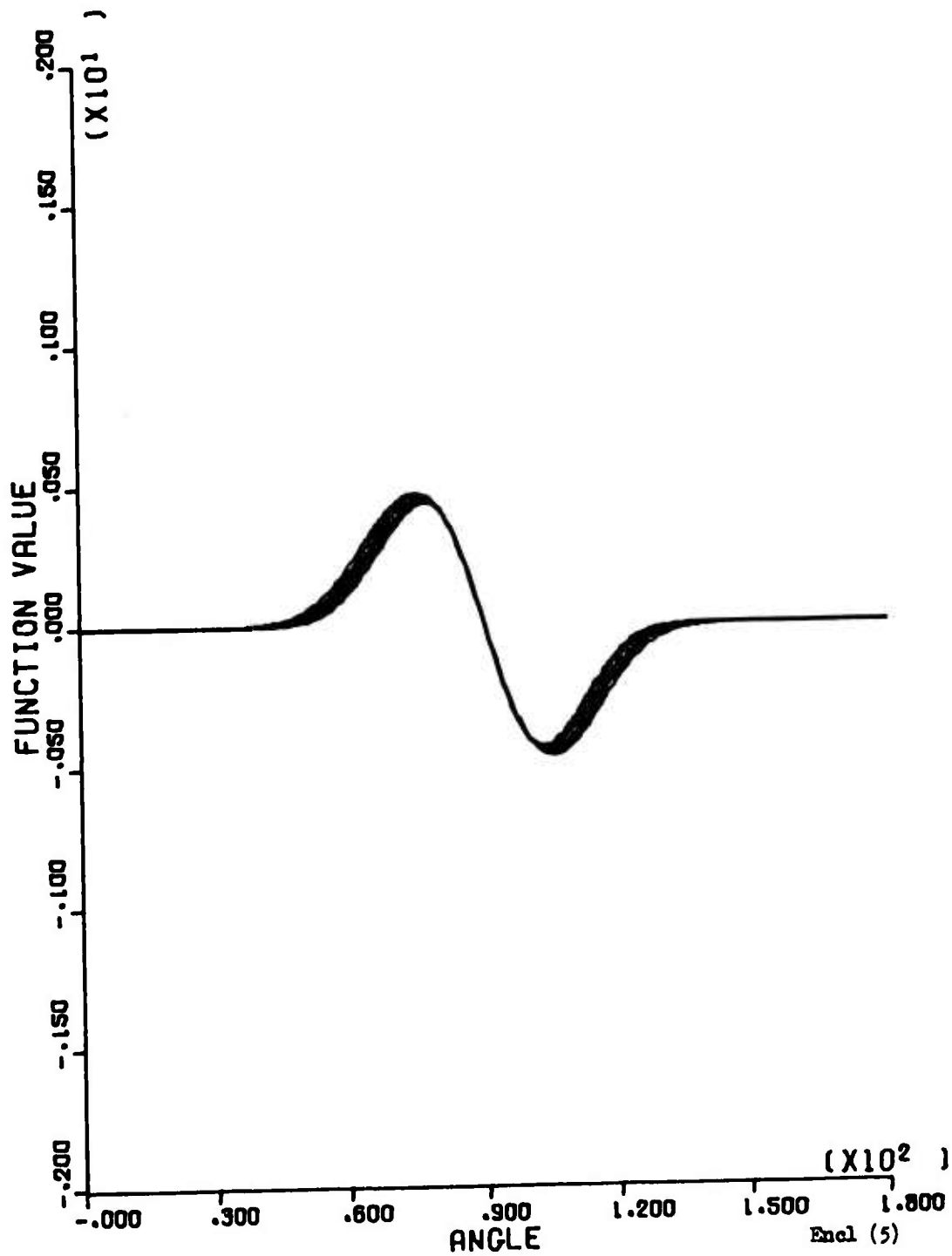


FIG. 5. STABLE PROCEDURE FOR $P_n^{n-1}(\cos\theta)$ FOR $n = 15$ THROUGH 20



APPENDIX A
LISTING OF FUNCTION SUBROUTINES

```

FUNCTION SPNM(N,M,X)
  SPNM=PNM(N,M,X)
  IF (M.EQ.0.OR.M.GT.N) RETURN
  SPNM=SQRT(2.*ANF(N-M)/ANF(N+M))*SPNM
  RETURN
END

```

```

FUNCTION PNM(N,M,X)
  PNM=0.
  IF (M.GT.N) RETURN
  IF ((ABS(1.-ABS(X))).GT.1.E-9) GO TO 20
  IF (M.NE.0) RETURN
  PNM=-1.
  IF (X.GT.0..OR.N.EQ.INT(FLOAT(N)/2.+1)*2) PNM=1.
  RETURN
20 CNM=(2.**N)*ANF(N)*ANF(N-M)
  PNM=1.
  IF (M.NE.0) PNM=SQRT((1.-X**X)**M)
  CNM=ANF(2*N)*PNM/CNM
  PNM=1.
  IF (M.NE.N) PNM=X**(N-M)
  IF (N-M.LE.1) GO TO 40
  PRD1=1.
  NT=(N-M)/2
  DO 30 I=1,NT
    AN1=FLOAT(N-M-2*I+2)
    AN2=FLOAT(2*I)
    AN3=FLOAT(2*N-2*I+1)
    PRD1=-PRD1*AN1*(AN1-1)/(AN2*AN3)
    NE=N-M-2*I
    AN1=1.
    IF (NE.GT.0) AN1=X**NE
    PNM=PNM+PRD1*AN1
  30 CONTINUE
  40 PNM=CNM*PNM
  RETURN
END

```

```

FUNCTION ANF(N)
DOUBLE PRECISION AN
ANF=1.
AN=1.00
IF(N.LT.0) PRINT 01
IF(N.LT.2) RETURN
DO 10 I=2,N
10 AN=AN*DBLE(FLOAT(I))
ANF=SNGL(AN)
RETURN
01 FORMAT(37H1FACTORIAL INTEGER IS LESS THAN ZERO.//)
END

```